Saturn's Rings: the Large and Small

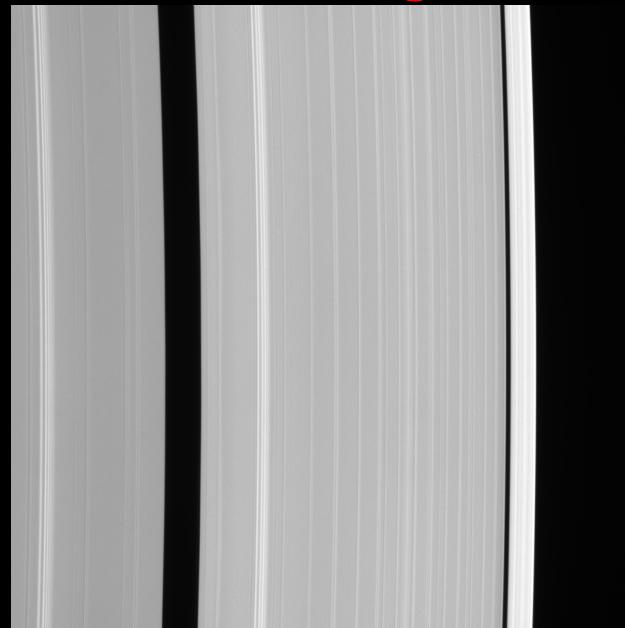
Colin Mitchell John Weiss (Carolyn Porco) 31 July 2007 CICLOPS



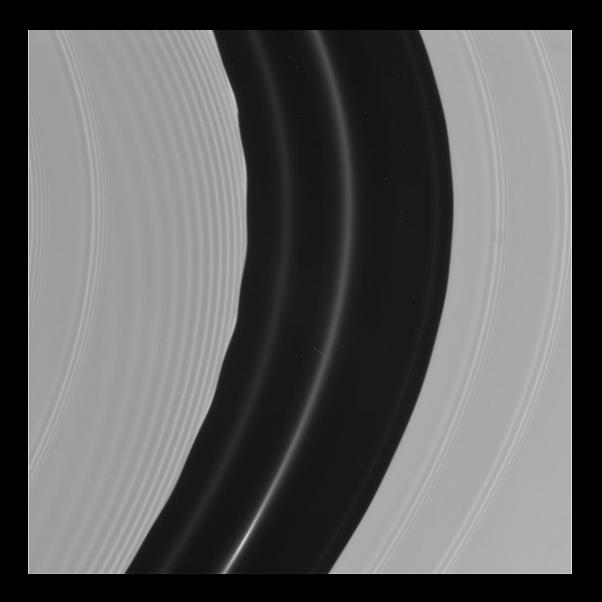
CENTRAL LABORATORY FOR OPERATIONS

Those Tiny, Little Moons

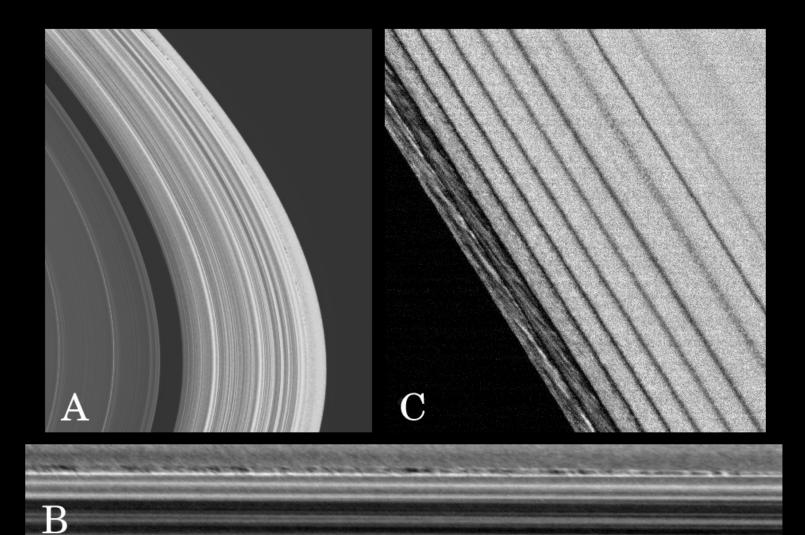




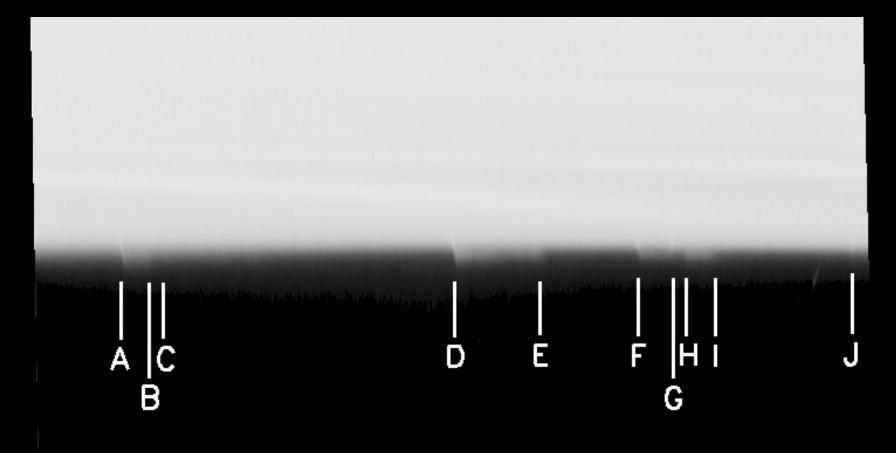




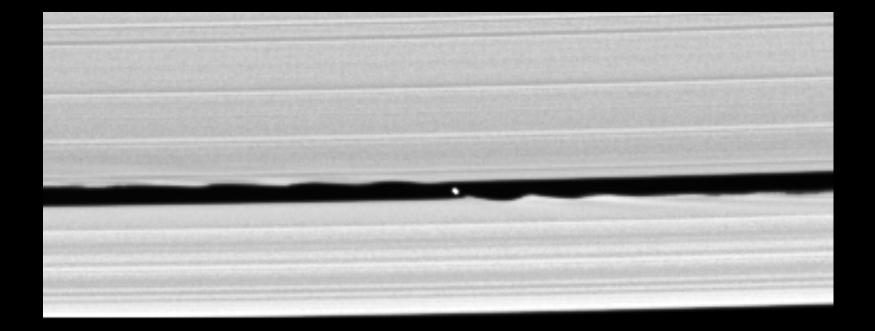
Edge of the Encke Gap







Keeler Gap with Daphnis



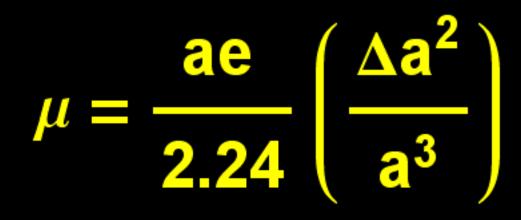
The Main Question:

 How to determine the mass of an embedded moon from the effects on the nearby ring edges?

Other Questions:

- Can the moonlet induce a time-variable structure?
- How does the eccentricity of the moon change these results? What about the particles' eccentricities?

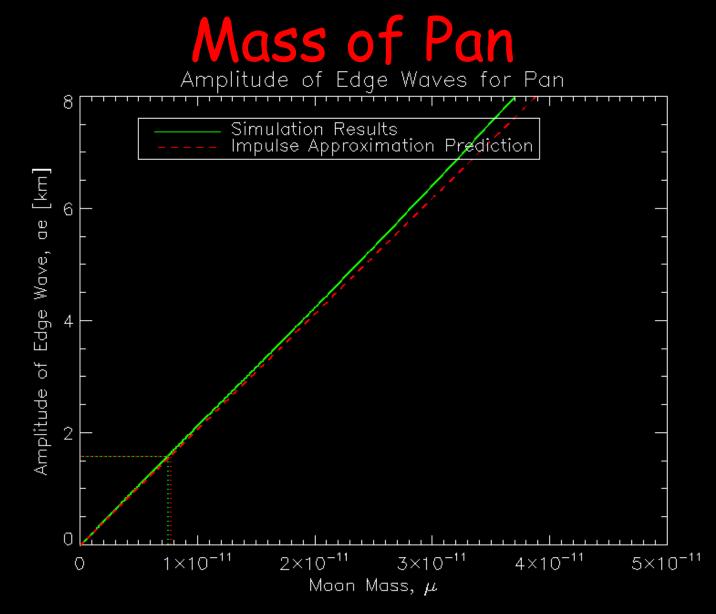
Previous Work



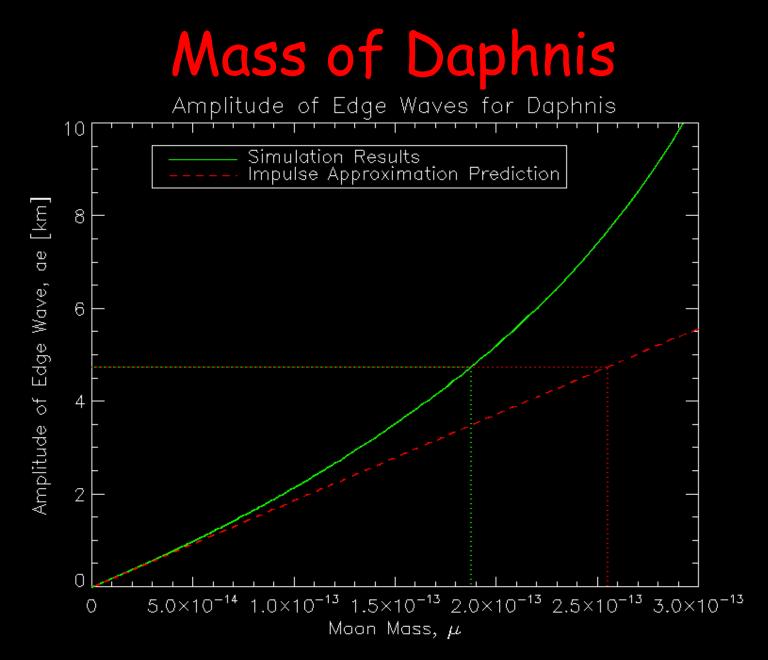
- Nice, analytic expression
- Uses an impulse approximation
- Assumes particle and moon are initially on circular orbits

Approach

- Integrate orbits of particles passing near a moon.
- Particles and moons can have eccentric orbits and arbitrary phase.
- Particles do not interact with each other.



 Analytic theory agrees with integrations to about 3%

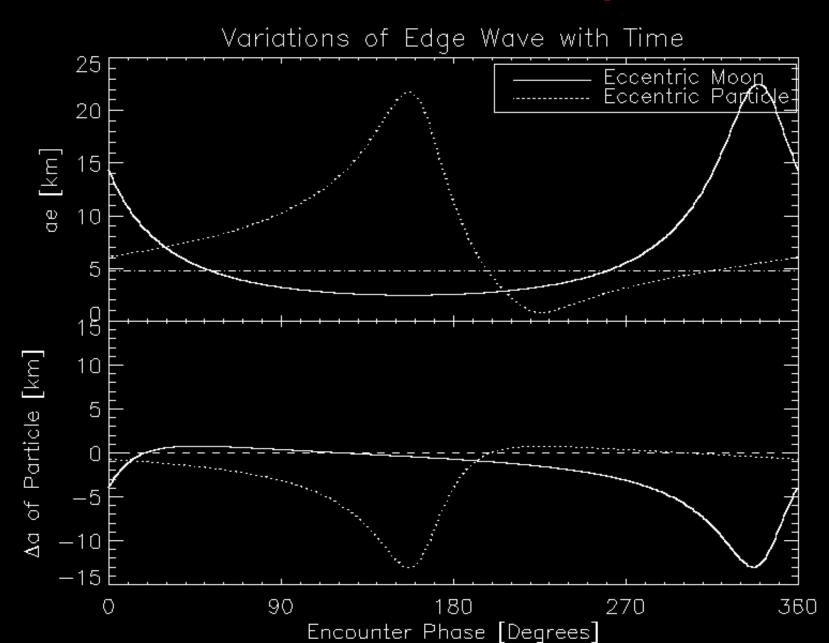


Analytic theory disagrees with integrations by 35%

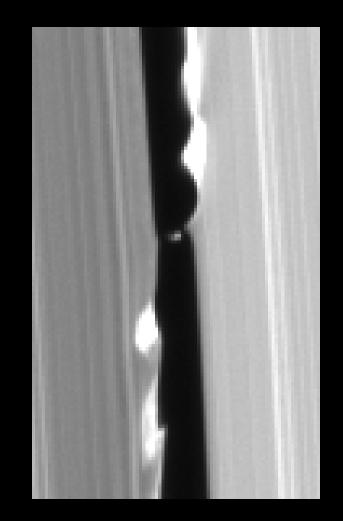
What Does Eccentricity Do

- Following runs give either moon OR the particle eccentricity
 - e is the same for both case, equal to what is measured for the moon
- Observe the variations in dX and ae for the particles
 - Note ae is not equal to the edge wave amplitude! (I'll show you what I mean later)

That Eccentric Daphnis



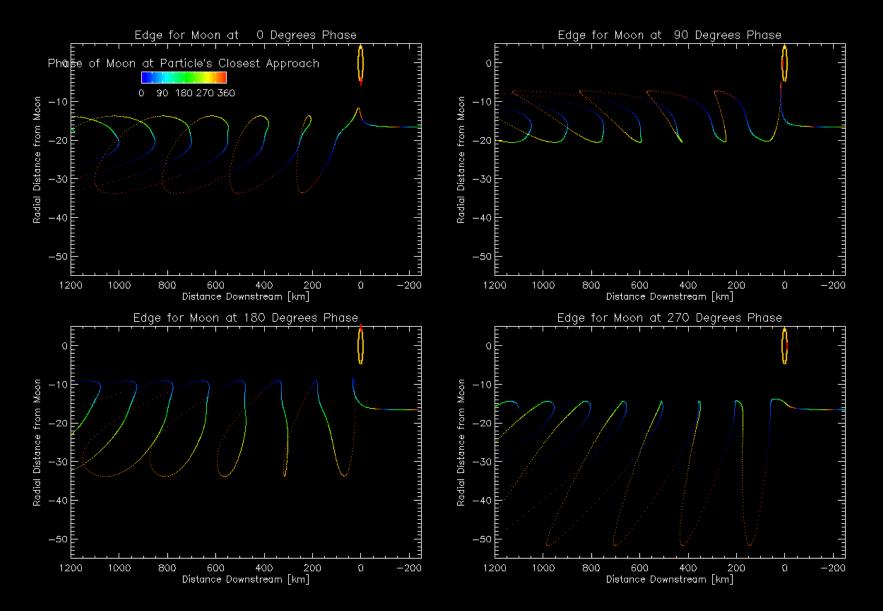
Daphnis Again



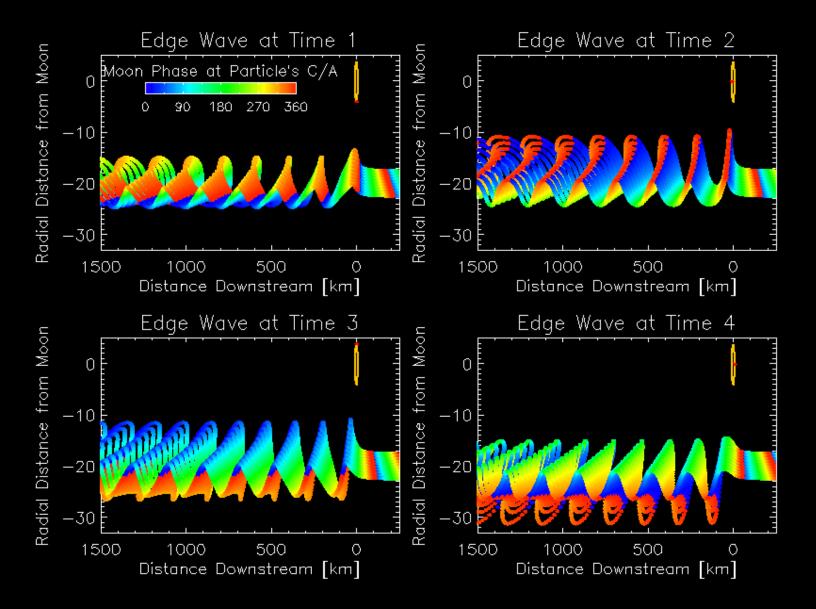
Why "ae" isn't Amplitude

- Because you see particles at all different phases making up the edges.
- This complicates things a lot!

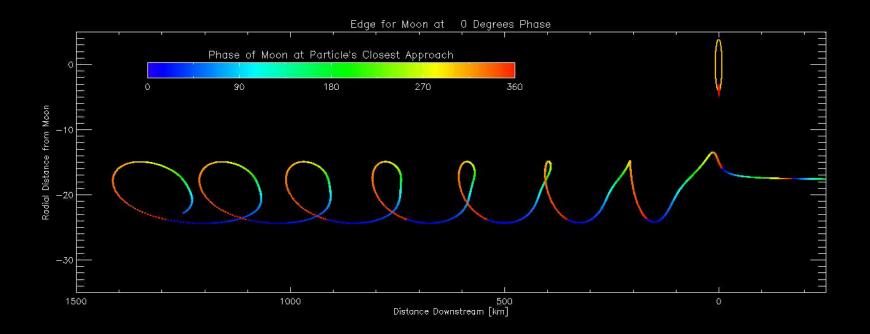
Graphically...



More Interesting Still



Movie



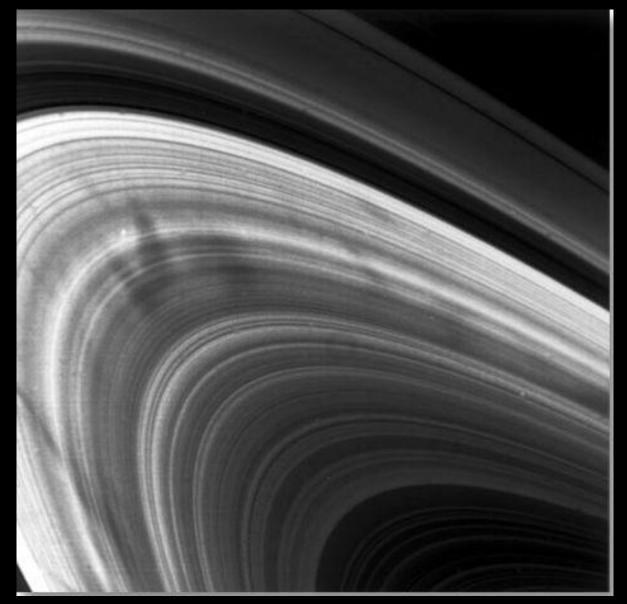
So What?

- The analytic formula works for distant moon-edge encounters, but not for nearby ones.
- Eccentric moons cause time-variable edges
- Need better simulations (*i.e.*, include particle collisions and gravity) to improve mass estimates

Who Cares?

- What all of this means is that small moons can have big effects on rings.
- This is a kind of "open hunting license" for Cassini to go hunting for moons which might have been too small to see yet.
- Knowing the masses of the moons can help us understand their origins... but that's a different part of the talk.



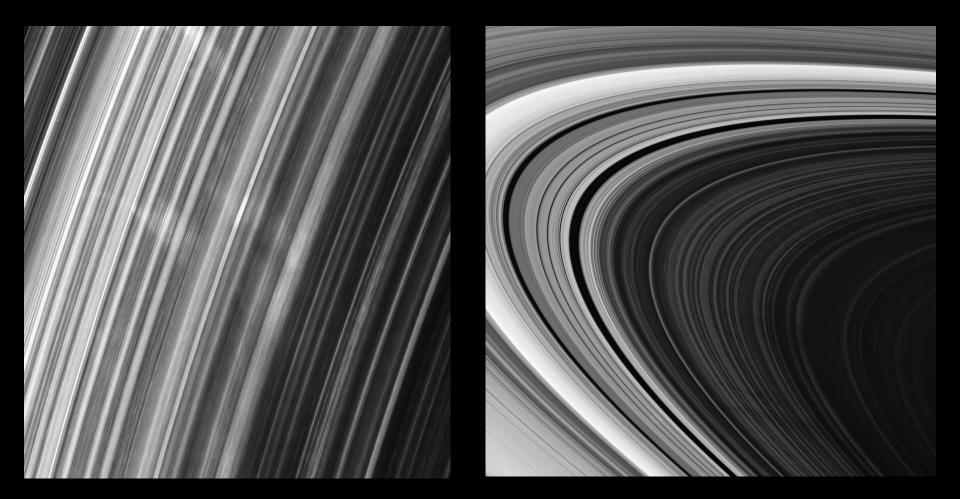


Introduction

Spoke Observations:

Voyager (1980 & 1981) : Discovery Hubble (1994-1998) : Faded out Cassini (2005-???) : Recovered Composed of small, charged dust particles (~0.5µm) Start off radially aligned Most exhibit Kepler shear Preferentially form at the SKR active region during a particular magnetic field orientation

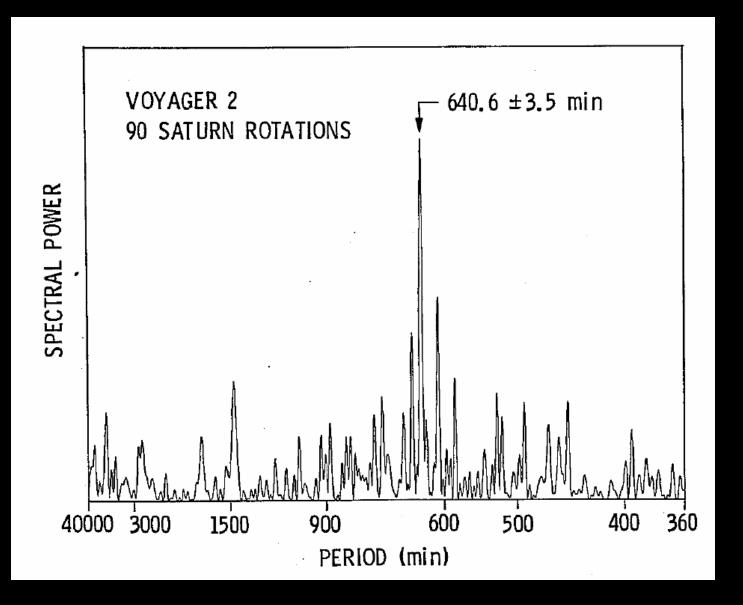
Contrast Phase Dependence



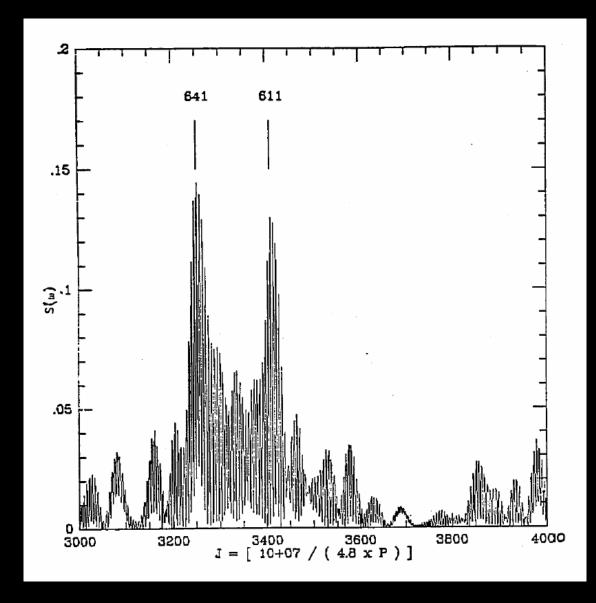
High Phase



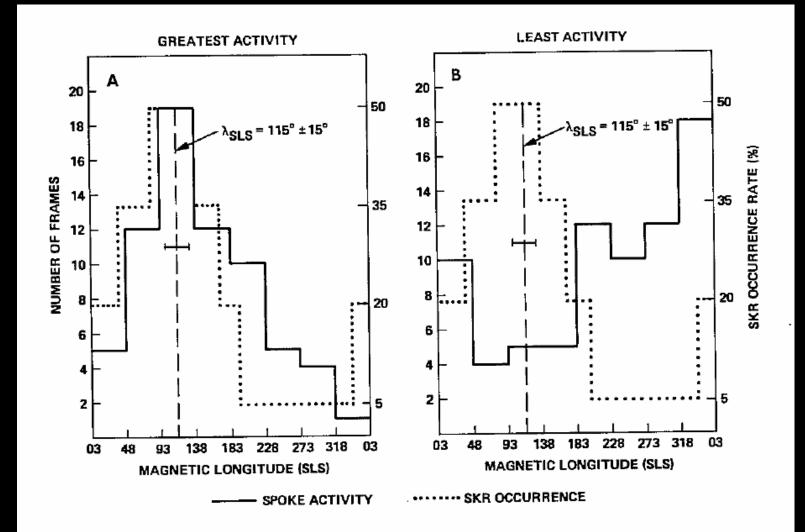
Spoke Periodicity



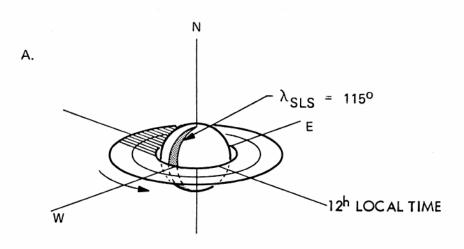
Spoke Periodicity



Spoke Periodicity

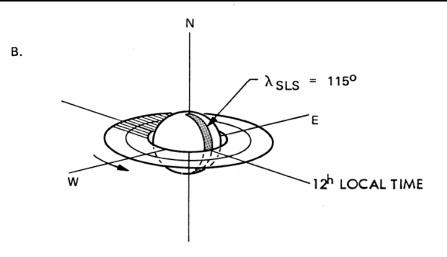


Preferred Geometry

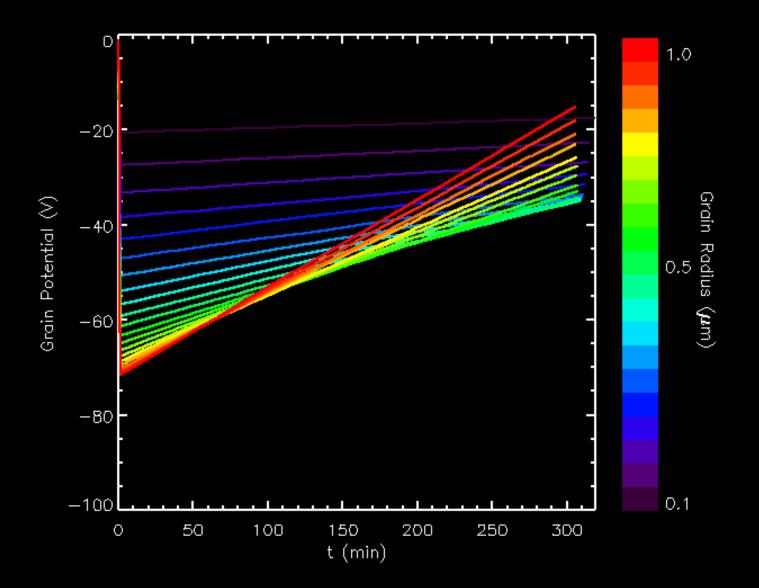


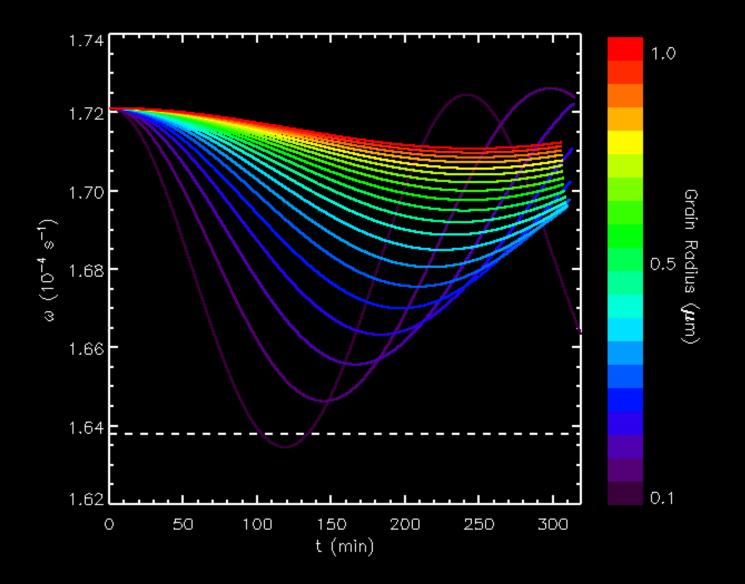
Preferential Geometry for high Spoke Activity

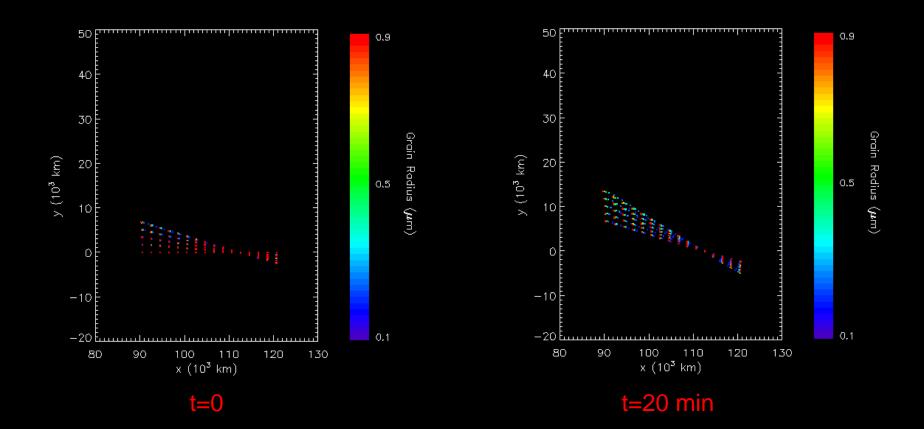
Peak SKR Emission Geometry

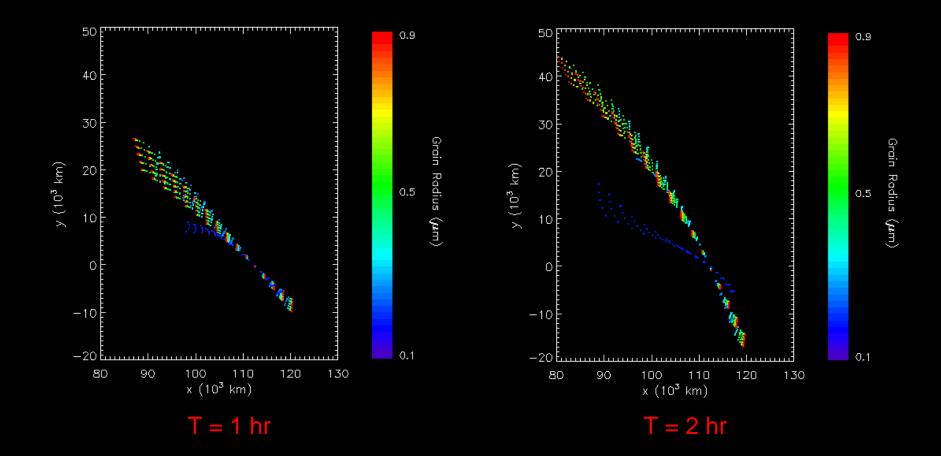


Porco and Danielson, 1982









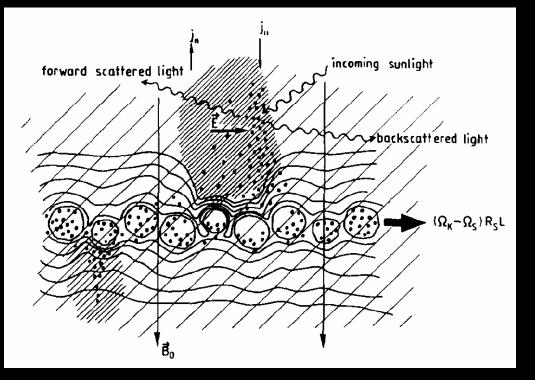
Formation Theories

Need to explain:

Charging and levitation of small grains Morphology (some long and narrow, others broad) Multiple spokes at nearly same location Periodicity with SKR and SED Rapid formation (perhaps < 5 mins)

Two possibilities: Impacts Lightning

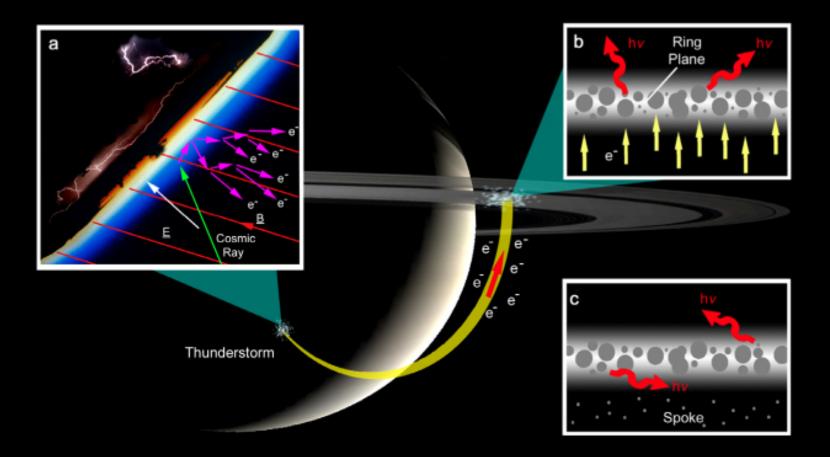
Impacts



Impact produces plasma Plasma charges dust **Dust levitates** Electric field is produce as dust leaves plasma E x B drift pushes plasma radially

Goertz and Morfill, 1983

Lightning



Jones et al. 2006

Model Features

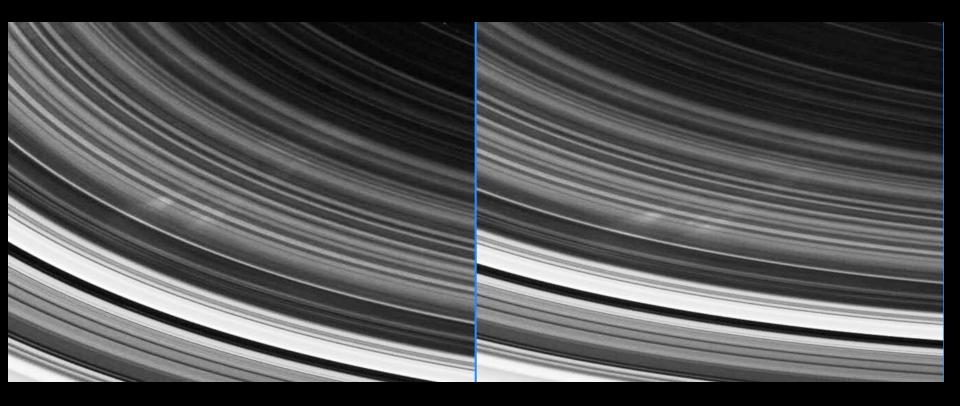
Impact:

Explains long narrow spokes well Clusters of spokes would imply multiple impactors Radial motion of plasma may not be as fast as once thought

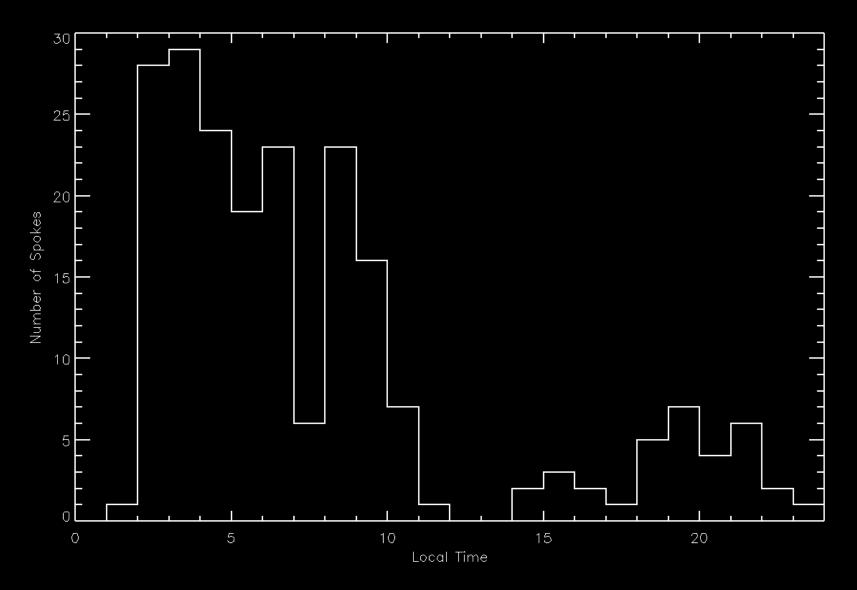
Lightning: Explains groups of spokes well Has difficulty with long narrow spokes

Cassini Results

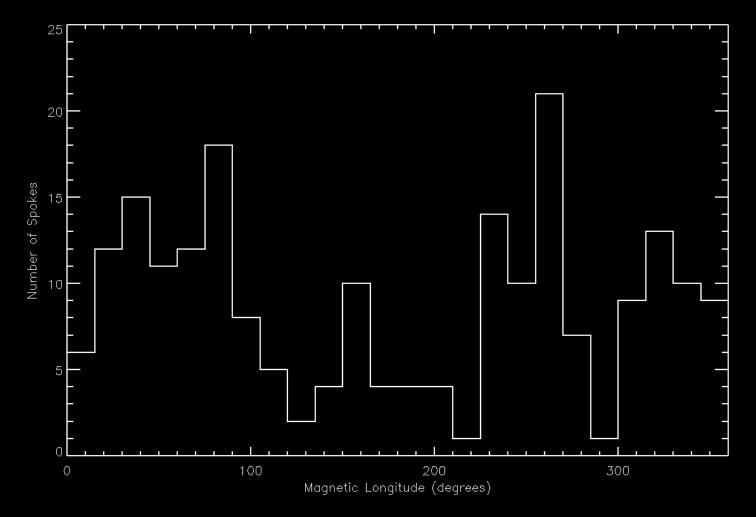
Didn't see any spokes until Sept. '05



Local Time

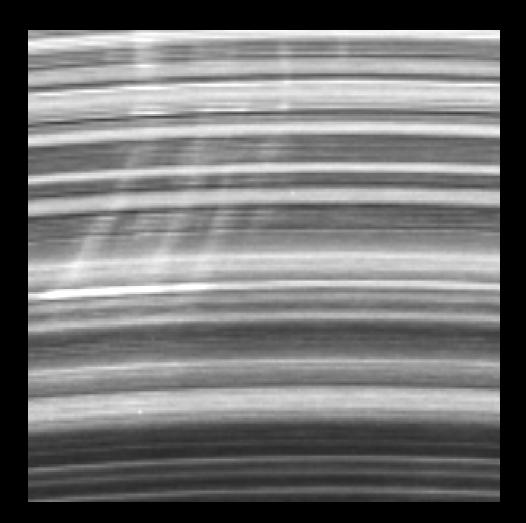


Magnetic Longitude



Mag. Coordinate system by Kurth et al. 2006

Morphology



Spoke Conclusions

The spokes have returned

They appear most abundant on the morning ansa, as expected from Voyager results

They have yet to display any significant dependence on the magnetic field sector

We have yet to see a spoke form

Much work to be done...

Growth of Ring Moons





(Not a lemon)



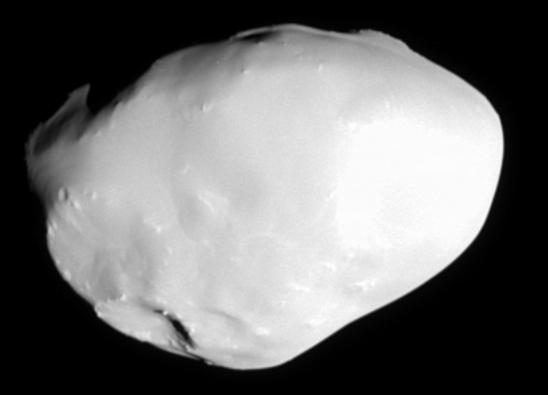
Prometheus



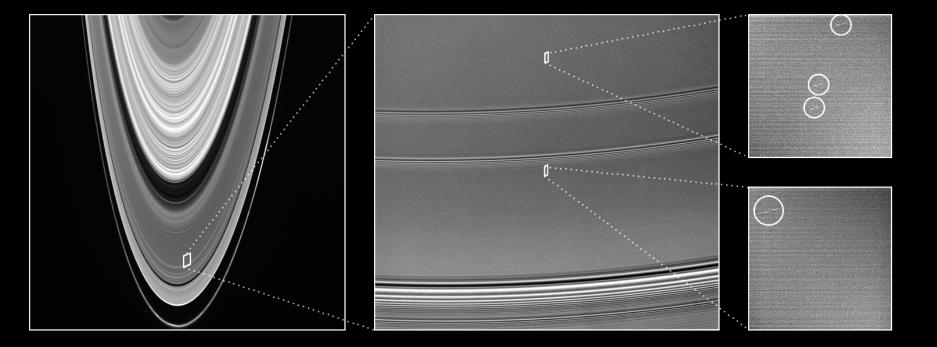
Pandora



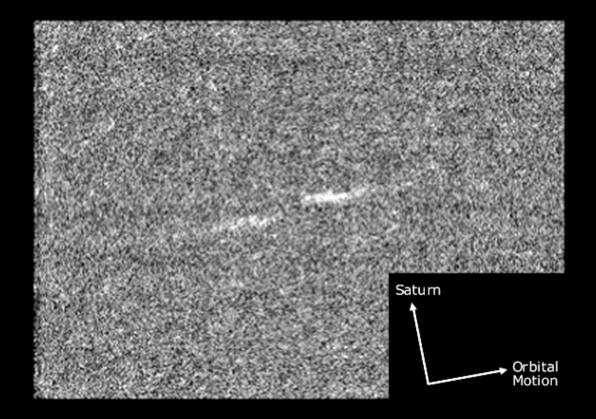




Serendipity

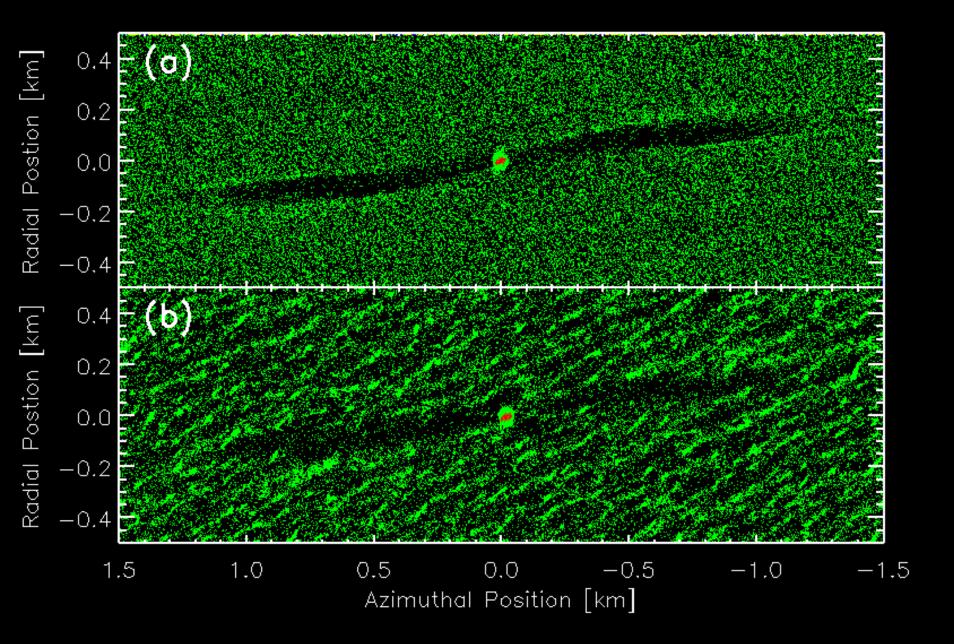


A Propeller in the Rings?

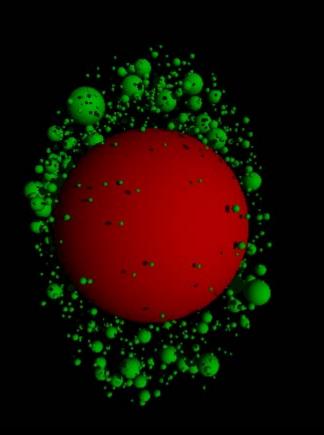


What Was That?

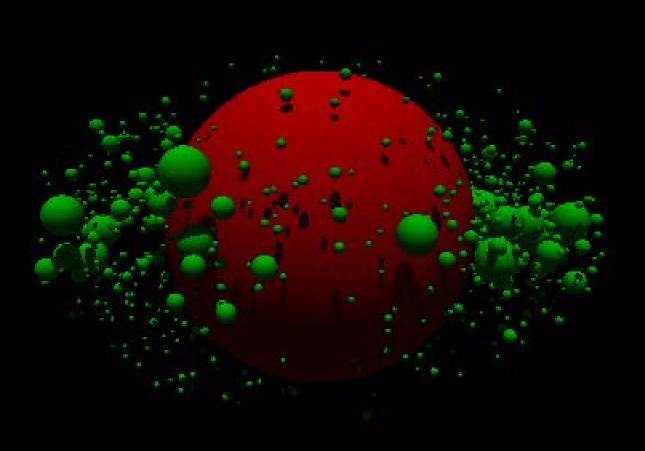
- Probably the signature of a moonlet (~50 m) embedded in the ring.
- What are we actually seeing? Only one way to find out!



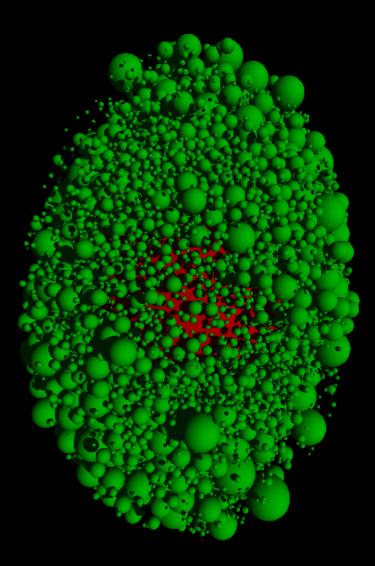
Start of Simulation



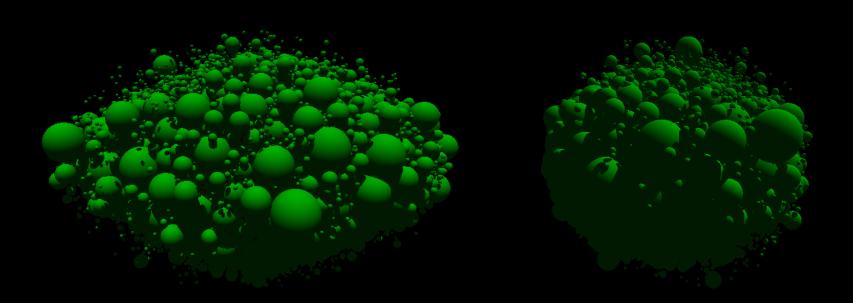
Side View



15 Orbits Later...



Side Views



From along y-axis From along x-axis

Why Does Growth Stop?

- Roche Zone = Region around the moon where the moon's gravity dominates
- If one assumes a particular shape for the moonlet, we can work out the critical density at which the moonlet exactly fills its Roche zone: ~ 0.5 g cm⁻³
- This density is independent of size of the moon

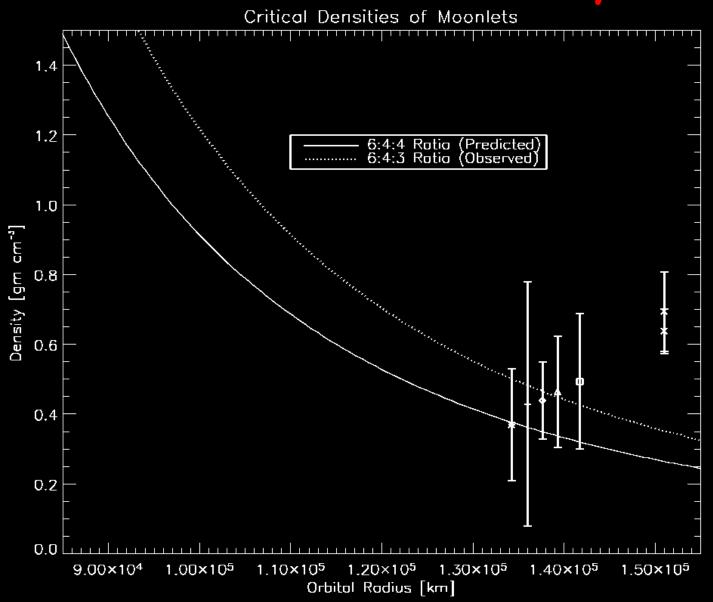
How Long Does it Take?

- Actually, just a few orbits to grow to the final size
- Particles don't always stick around, though
- Given lots of orbits into and out of sunlight, particles might "weld" themselves together, making a more solid moon

Does Growth Always Stop?

- No.
- It depends on how far you are from the planet and how dense the infalling material is
- This theory predicts that the planets, for example, can grow as large as they want. (Which is good, because we're not shaped like lemons!)

How Dense Are They?



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