



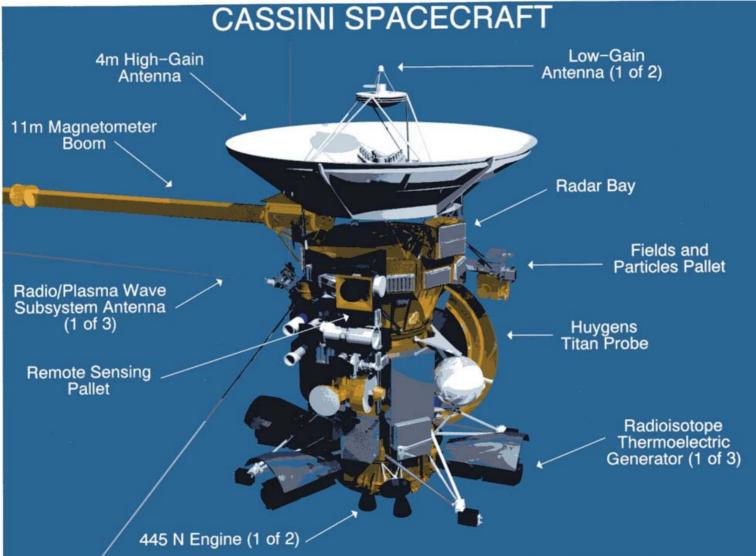
Cassini Spacecraft Engineering Tutorial

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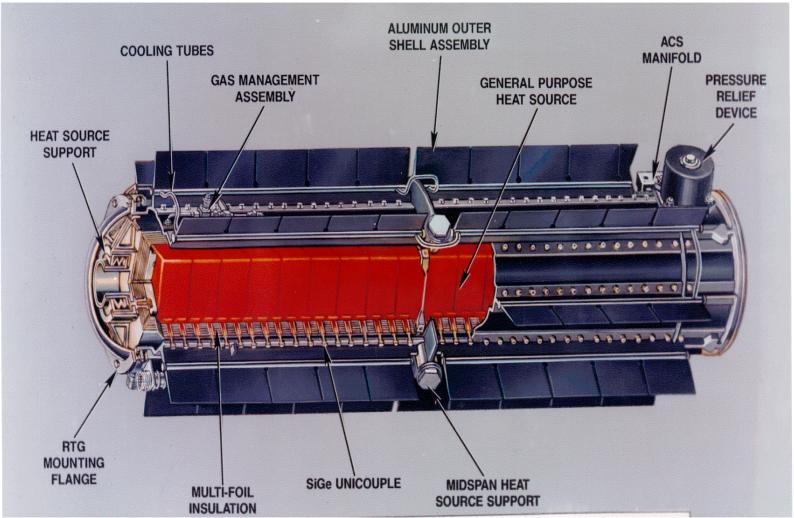
Cassini/Huygens Spacecraft Overview

- Largest outer planetary spacecraft ever built
- Weighed 5574 kg at Launch (3132 kg Propellant)
 - Today's weight is around 2710 kg with 422 kg Propellant left
 - 3 maneuvers (Deep Space Maneuver, Saturn Orbit Insertion, Periapsis Raise Maneuver) used 2083 kg of propellant
- 58 microprocessors
- Twelve Science Instruments
 - Fields and Particles
 - CAPS, CDA, INMS, MAG, MIMI, RPWS
 - Optical Remote Sensing
 - CIRS, VIMS, ISS, UVIS
 - Radar
 - Radio Science
- Huygens Probe















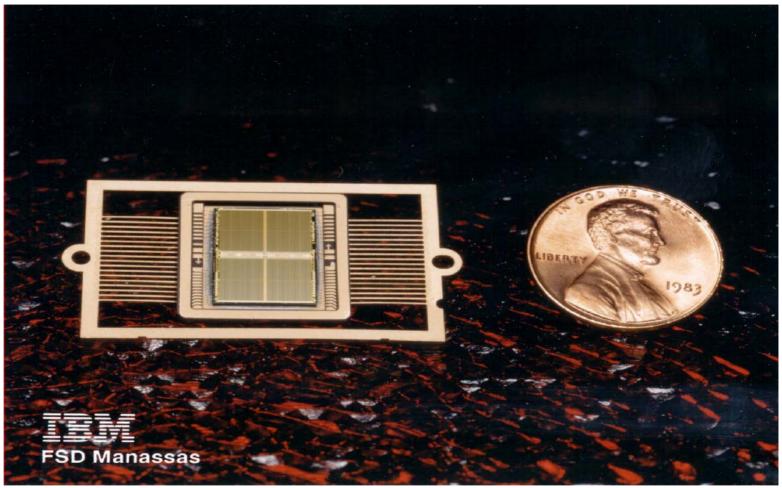
Power System

- Power supplied by three Radioisotope Thermoelectric Generators (RTGs)
 - 875 Watts at Launch, 692 Watts at End of Mission (July 2008)
 - Loses about 1 Watt/month now, will slow to 9 Watts/year by 2007
 - Excess power distributed to shunt regulator/radiator
- Power distribution using 192 Solid State Power Switches
 - Switches act a circuit breakers in over load condition
 - Critical loads have redundant switches and automatic turn-on circuitry
 - Special Fault protection in case of SSPS "trip"
- Not enough power to keep all instruments on all the time, so Operational Modes are used
- Picture shows RTG schematic















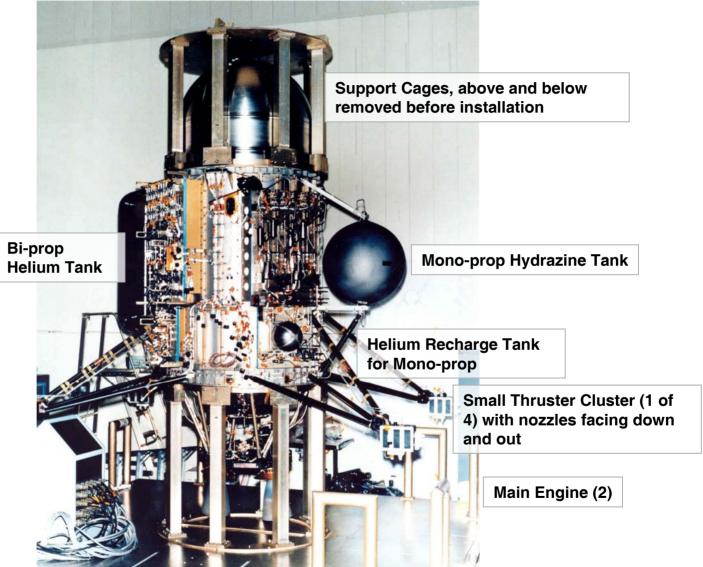
Command and Data System

- Two redundant 1553 Buses with fully redundant hardware
- Flight computer (EFC) IBM 1750A architecture
 - 512k RAM
 - ~48,000 lines of Ada and assembly source code
- Processes all commanding for the spacecraft engineering sub-systems and instruments
 - Background sequences uplinked every 3-6 weeks
 - "Real-time" commands are uplinked during downlink passes
- Processes all telemetry for the spacecraft engineering sub-systems and instruments
 - On-board data storage and playback using 2 Solid-State Recorders
 - 2.3 Gbits total, 2.1 Gbits available for telemetry (each SSR)
- Picture is the IBM 1750A processor chip















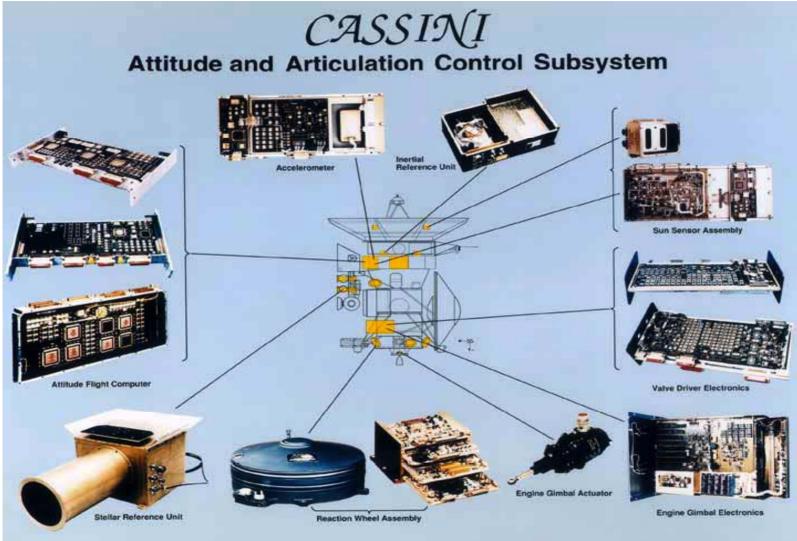
Propulsion System

- Two completely isolated systems
 - Monopropellant System Hydrazine (H₂N₄)
 - 16 Redundant ~1 Newton thrusters, 4 per "cluster"
 - Used for attitude control and small (< 0.3 m/s) trajectory control maneuvers
 - Blow-down mode
 - Single Helium pressure recharge will be done 4/10/06
 - Bipropellant system- Nitrogen Tetroxide (NTO)/Monomethyl Hydrazine(MMH)
 - Main (445 Newton) engine for propulsive maneuvers
 - Burns can be blow-down or pressurized (with Helium)
 - Engines are protected from micrometeoroid particles by articulating cover
 - Redundant main engine (never used in flight)
- Picture shows Propulsion Module as delivered from Lockheed-Martin Astronautics















Attitude Control System

- Two redundant 1553 Buses with fully redundant hardware
- Flight computer (EFC) IBM 1750A architecture (same as CDS)
- Attitude determination
 - Attitude determined by Stellar Reference Units
 - ~3500 Star on board catalog, picking the 4 or 5 brightest stars in the field of view at any given time
 - Attitude rates supplied by Hemispheric Resonator Gyros (HRGs), aka Inertial Reference Units

Attitude Control

- 3 axis stabilized using thrusters or 3 reaction wheels (RWAs)
 - Small (1 Newton) thrusters used when pointing isn't as important or when faster turns are needed
 - RWAs provide pointing stability better than 100 microradians over 1000 seconds

Propulsive maneuvers

- Uses small thrusters or main engine (445 Newton) for trajectory control
- Picture shows various components of Attitude Control System

















The spacecraft turns to Earth for a typical 9 hour tracking pass, to downlink telemetry data and receive commands.

The DSN radiates an X-Band uplink signal which the spacecraft locks to (just like your FM radio receiver) and then the spacecraft multiplies the uplink signal by a exact (880/749) amount to form the coherent downlink carrier. This gives Navigation and Radio Science very precise two-way coherent Doppler and ranging tones for spacecraft location.

Deep Space Network – 3 complexes: Barstow, Ca, Madrid, Spain and Canberra, Australia Each complex has a 70 meter antenna, and multiple 34 meter antennas.









Telecommunications System

Antenna

- 4 meter High Gain Antenna (HGA) used for main communications
 - HGA is a multi-aperture antenna used jointly for Telecom (X-Band), Radio Science (S, X (same as Telecom) and Ka-Band), RADAR (Ku-Band) and Probe mission (S-Band, different frequencies than Radio Science)
 - HGA was used as a Sun shield during Inner Cruise
- 2 Low Gain Antennas used for emergency X-Band emergency communications
 - LGA-2 located below Probe was used only during Inner Cruise

Cassini Deep Space Transponder

- Receives uplink X-Band signal (and commands) from DSN
- Downlinks
 - Ultra-Stable Oscillator (USO) used for 1 Way Doppler
 - Voltage Controlled Oscillator (VCO) used for 2 Way Coherent Doppler
- Ranging channel for Navigation
- 20 Watt X-Band downlink Traveling Wave Tube Amplifier
- Data Rates
 - Uplink nominal rates at 250 or 500 bps
 - Emergency rate is 7.8125 bps
 - Downlink nominal tour rates from 22.12 165.9 Kbps
 - Depends on Deep Space Network antenna and Earth-Saturn range
 - Emergency downlink is 5 bps
- Picture 1 taken during Assembly, Test and Launch Operations shows HGA, LGA-1 and Upper Equipment Module.















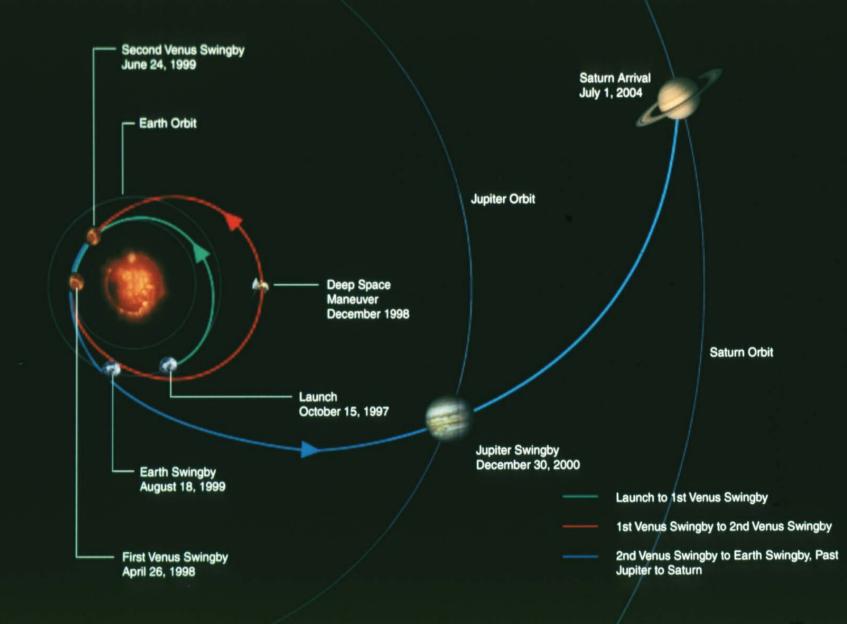
Thermal System

- Thermal System designed to withstand the Sun's heat at Venus distances and the cold of 0.1 Sun at Saturn distances
- Passive components
 - Blanketing and surface treatments
 - Radio-isotope Heater Units (RHUs) used in some locations
 - Instrument radiators that face "cold sky"
- Active components
 - Louvers on some of the electronic bays
 - Variable Radio-isotope Heater Units (VRHUs) used in some locations
 - Autonomous (thermostatic) Thermal Control of key selected components
 - Replacement and supplemental heaters for the instruments necessary because the instruments are turned on or off by Operational Modes
- Picture show spacecraft being readied for Solar Thermal Vacuum Test





Cassini Interplanetary Trajectory







Getting to Saturn

Launch and TCM-1

- Performed flawlessly on 10/15/97
- CDS and AACS were nominal
 - Switched prime CDS and AACS computers in first week (to change from launch to cruise parameters)
- Telecom performance was above predict for LGA-2
 - HGA used as sunshade, so low bit rate early mission was necessary
- TCM-1 executed within 3% error (well within errors for uncalibrated TCM)
 - MEA Cover closed after launch, opened for TCM-1, closed after TCM-1

Early Checkout 10/97 - 3/98

- CAPS, CIRS, MIMI, CDA launch latches released
- UVIS membrane puncture
- RPWS antennas deployed
- Performed AACS checkout
 - Inertial Reference Unit-B and Engine Gimbal Assembly-B
- Probe Checkout
- Instrument checkout
 - Only INMS and RFIS (requiring HGA) not seen
 - All checkouts were nominal







Getting to Saturn (continued)

- Venus 1 flyby on 4/26/98
 - Altitude 284 km
 - RADAR and RPWS performed science calibrations
- Deep Space Maneuver (TCM-5) performed on 12/3/98
 - 90 minute regulated burn, just like SOI!
- Instrument Checkout on 1/99
 - 25 days on HGA allowed instruments to perform real calibrations with high rate science data
 - RFIS saw Ka-Band and S-Band for first time
 - ORS imaged Spica
 - DFPW rolls performed
- Venus 2 flyby on 6/24/99
 - Altitude 603 km
 - CAPS, CDA, MAG, MIMI, RPWS, ISS, UVIS and VIMS performed science calibrations







Getting to Saturn (continued)

- Earth Flyby on 8/18/99
 - 4 TCMs prior to flyby done for Earth avoidance
 - VIMS Cover release
 - MAG Boom Deploy
- Switch to permanent HGA on 2/1/00
 - Real science data rates now could be used!
- New AACS FSW loaded March 00
 - RWA capability finally in place
- New CDS FSW loaded August 00
 - Dual SSR ping-pong in place for Jupiter activities
- Jupiter Encounter (closest approach 12/30/00)
 - 6 month campaign with all instruments but INMS participating
 - Extensive science data taken
 - Demonstrated many tour science sequencing capabilities







Getting to Saturn (continued)

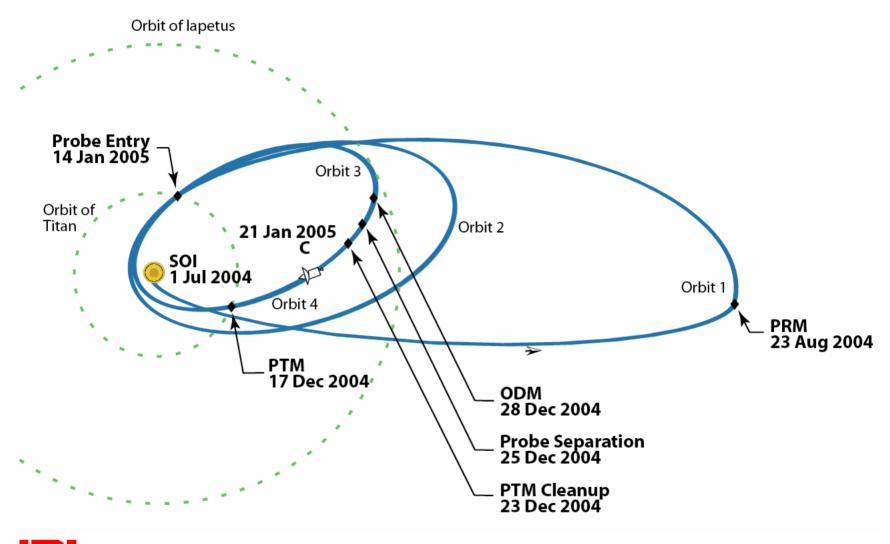
- New AACS FSW (A8) loaded February 03
 - Rotating Coordinates Probe tracking
 - Tracking for SOI (yaw steering)
 - IRU-B and SSA-B checked out
- New CDS FSW (Version 9) loaded March 03
 - Dual record for Probe Mission enhancements
 - SSR IEB capability for instruments
 - HGA Response algorithm
- Switched to Articulated RWA-4 to replace RWA-3 07/03
 - RWA-1,2,4 prime for "rest of Mission"
- A8.6.5 FSW uplinked 02/04
- A8.6.7 FSW uplinked 04/04
- TCM-20
 - First fully pressurized burn since DSM 5/27/04
- SOI 07/01/04!







SOI to Huygens Entry













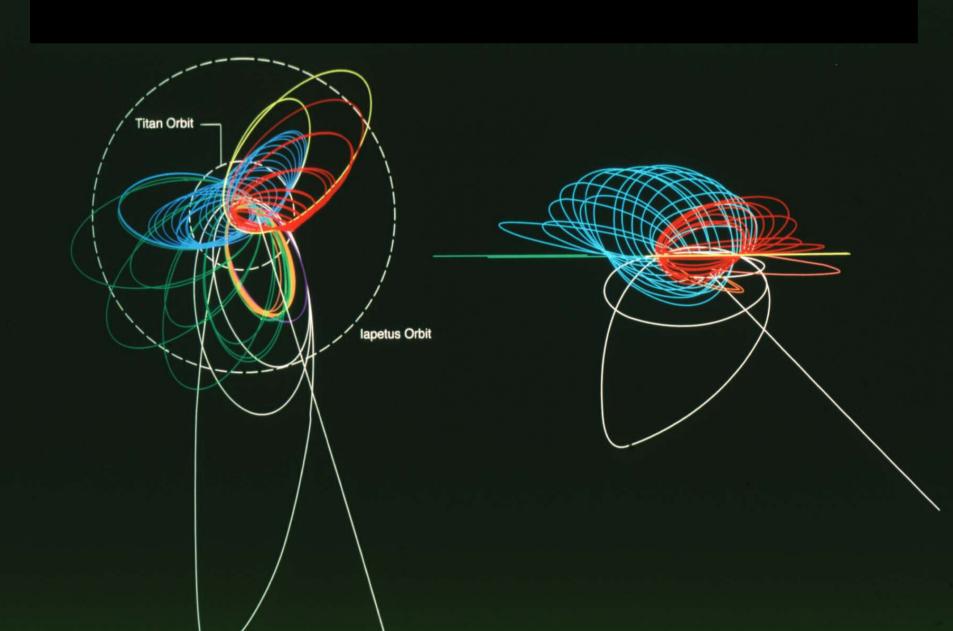




Huygens Probe

- European Space Agency Mission
 - Six instruments
 - Designed to study the surface and atmosphere of Titan
 - Released on December 25, 2004
 - Landed on Titan January 14, 2005
- Orbiter services to Probe Mission
 - Power for Probe Receiver
 - S-band telemetry link
 - Probe Relay Data storage and playback
- Picture is Huygens probe engineering test model after being drop tested









Staying on Tour

- Each orbit around Saturn takes 3 maneuvers (Orbit Trim Maneuvers)
 - Titan 3 days (Approach maneuver)
 - Titan + 3 days (Clean-up maneuver)
 - Apoapsis (used for next Titan and/or to change orbit size/angle)
- "Live Updates" for Saturn or moon encounters
 - Better position knowledge means we update the vectors
- Periodic Engineering Maintenance
 - Exercise "moving" parts to keep things lubricated
- Dust hazards
 - If the dust hazard is great enough we close the engine cover and turn the HGA to dust hazard
- Conjunction
 - Once per year, spacecraft and Earth are on opposite sides of the Sun
 - Quiet period on spacecraft because telecommunications is disrupted







For more information about the Cassini Huygens mission

http://saturn.jpl.nasa.gov

Also, see "Basics of Space Flight" available for JPL Public website

