



ASTROBIOLOGY

The Story of our Search for Life in the Universe



Produced by the NASA
Astrobiology Program to
commemorate 50 years of
Exobiology and Astrobiology
at NASA.

Astrobiology

A History of Exobiology and Astrobiology at NASA

This is the story of life in the Universe — or at least the story as we know it so far. As scientists, we strive to understand the environment in which we live and how life relates to this environment. As astrobiologists, we study an environment that includes not just the Earth, but the entire Universe in which we live.

The year 2010 marked 50 years of Exobiology and Astrobiology research at the National Aeronautics and Space Administration (NASA). To celebrate, the Astrobiology Program commissioned this graphic history. It tells the story of some of the most important people and events that have shaped the science of Exobiology and Astrobiology. Beginning shortly after the establishment of NASA, this field is relatively young. However, as you will see, the questions that astrobiologists are trying to answer are as old as humankind.

Concept & Story

Mary Voytek
Linda Billings
Aaron L. Gronstal

Artwork

Aaron L. Gronstal

Script

Aaron L. Gronstal
Jordan Rizzieri

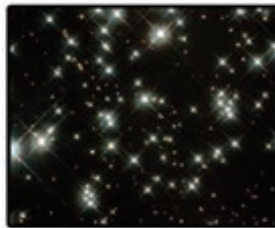
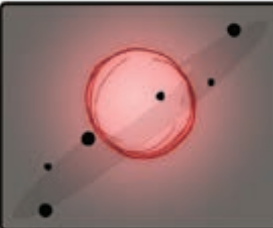
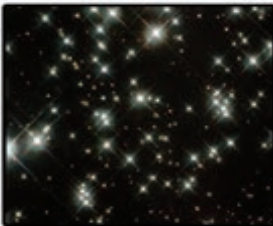
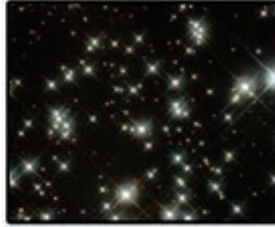
Editor

Linda Billings

Layout

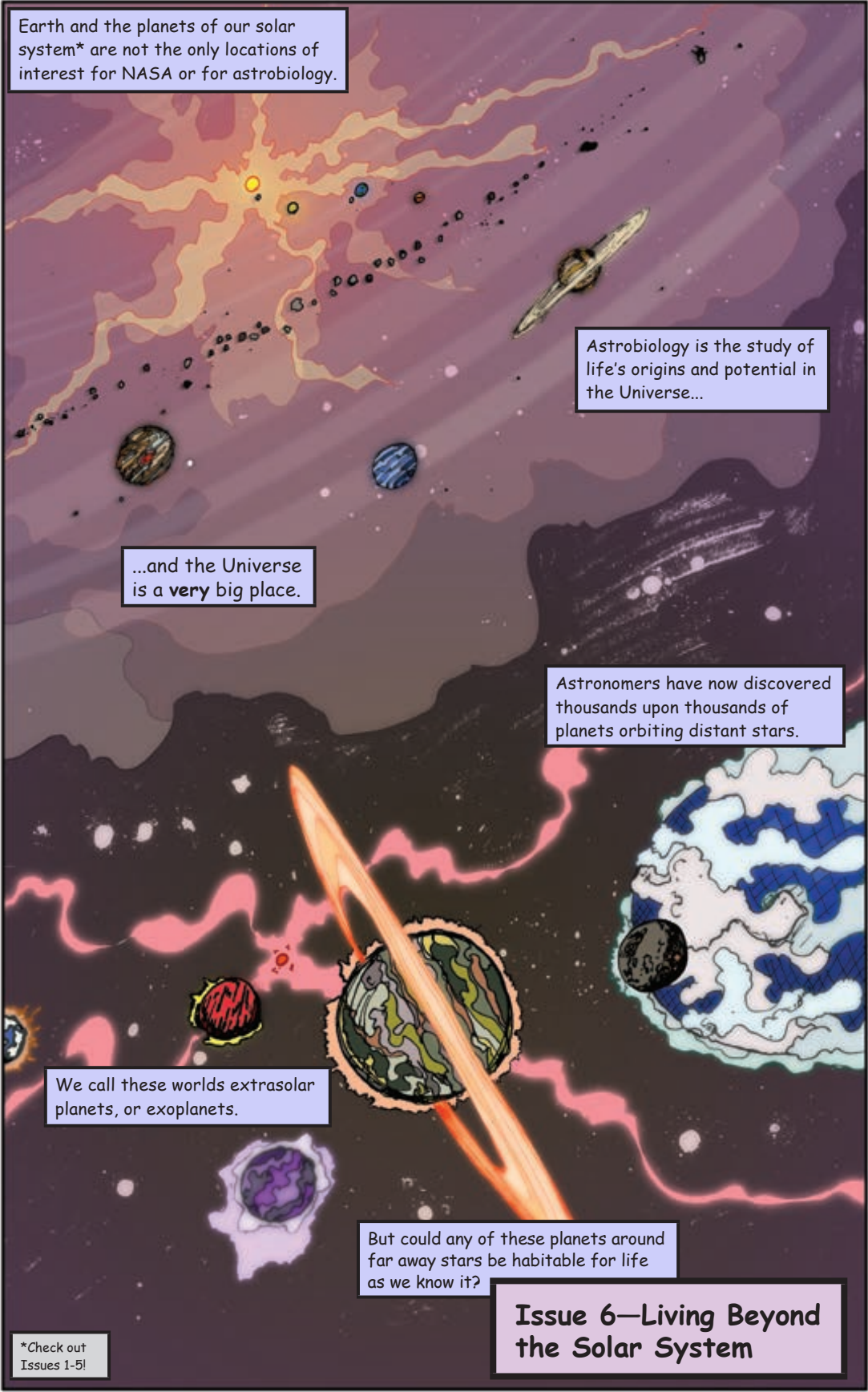
Aaron L. Gronstal
Jenny Mottar

Issue #6—Living Beyond the Solar System



The year 2010 marked the 50th anniversary of NASA's Exobiology Program, established in 1960 and expanded into a broader Astrobiology Program in the 1990s. We are telling the story of how this field developed and how the search for life elsewhere became a key component of NASA's science strategy for exploring space. This issue is the sixth in a series of graphic history books. Though not comprehensive, the series has been conceived to highlight key moments and key people in the field as it explains how Astrobiology came to be.

-Linda Billings, Editor



Earth and the planets of our solar system* are not the only locations of interest for NASA or for astrobiology.

Astrobiology is the study of life's origins and potential in the Universe...

...and the Universe is a **very** big place.

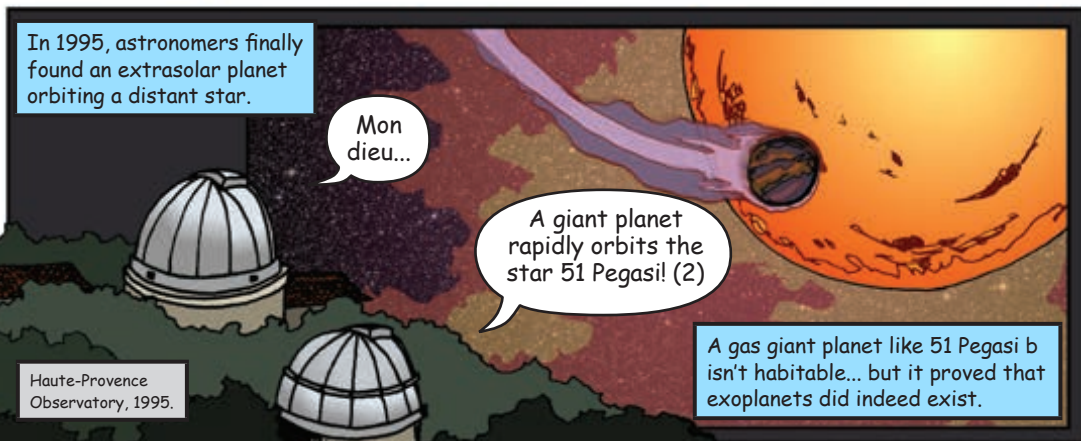
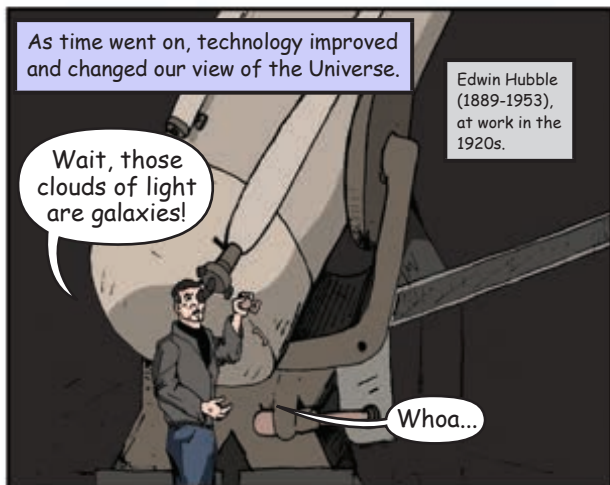
Astronomers have now discovered thousands upon thousands of planets orbiting distant stars.

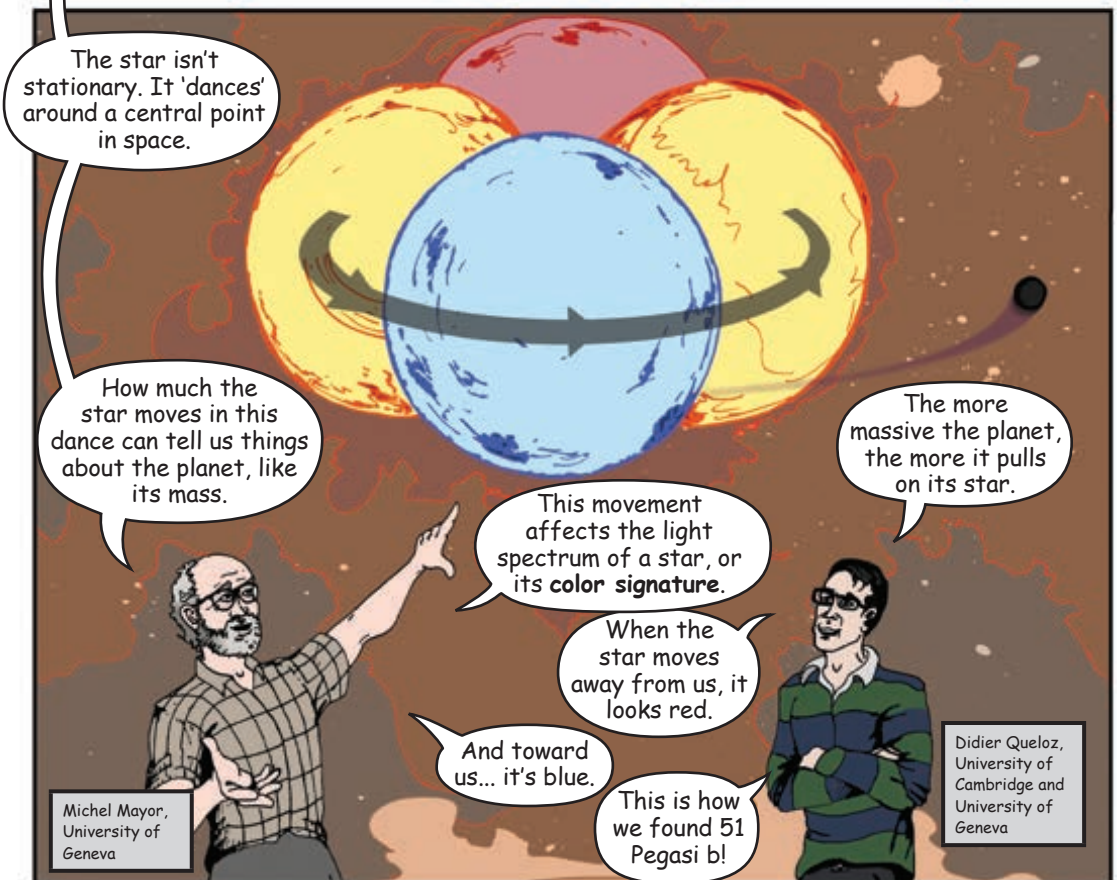
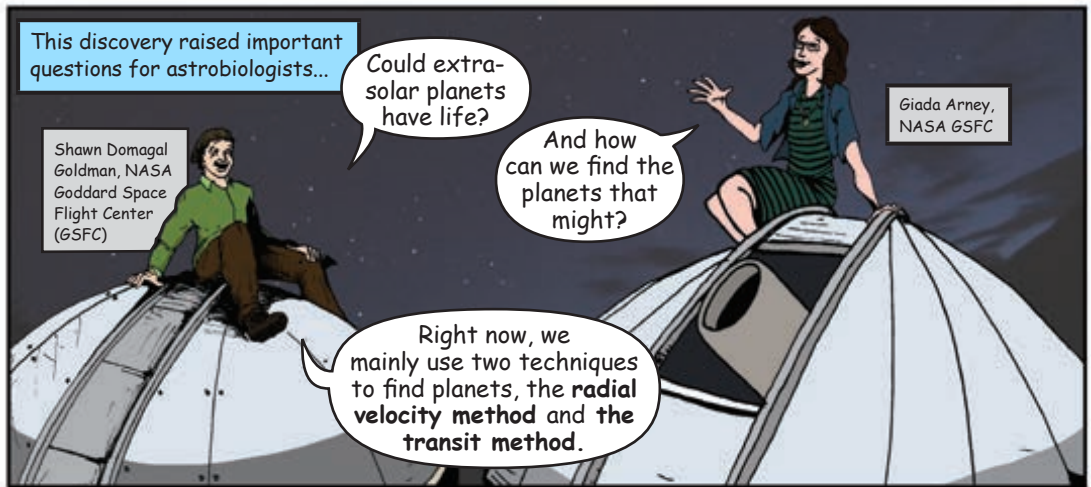
We call these worlds extrasolar planets, or exoplanets.

But could any of these planets around far away stars be habitable for life as we know it?

*Check out Issues 1-5!

Issue 6—Living Beyond the Solar System





The transit technique works when a planet passes in front of its star.

When a planet moves in front of its star, it blocks part of the light that we see.

Natalie Batalha, NASA Ames Research Center (ARC)

The amount of light blocked, and how often it happens, can tell us about the planet's size, its orbital period (or "year"), and its distance from the star.

If you know the radius of the star, that is.

Planets orbiting close to a star are easier to study, because close planets transit more often, giving us more observations to look at.

With both techniques, it's easier to spot big planets.

Gas giants similar to Jupiter have more gravity to affect their star. And they're big, so they block out more light when they transit.

Radial velocity was the first method to find exoplanets in large numbers, but now it has been overtaken by the transit method, both in the number of planets discovered and in its ability to characterize them. (3)

Bill Borucki,
NASA ARC

After the first exoplanet discovery, astronomers got to work hunting for more.

From 1995 onward, ground- and space-based telescopes were used to add to the list.

Victoria Meadows,
University of
Washington (UW)

New discoveries flooded in, including planets unlike any of the planets we see in the Solar System.

The first planets were the easiest to spot - huge gas giants similar to Jupiter.

But as techniques improved, smaller and smaller planets were discovered.

That's when things started to get very interesting for astrobiologists.

We found more than hot Jupiters close to their stars.

There were stars with multiple planets in orbit!

...and smaller worlds that could be rocky like Earth!

We now know the most common planets in the Universe are mini-Neptunes and the so-called 'Super Earths'.

Finding small, rocky planets raised questions about whether or not any could be habitable.

If we could find an inhabited planet, it could answer so many questions about Earth's place in the Universe.

Example of facility using:
*transit method, or
**radial velocity method

Spitzer Space Telescope (NASA, 2003-2020)

Gaia (ESA, 2013-2025)

W.M. Keck Observatory**,
Hawai'i

European Southern Observatory (ESO),
Cerro Paranal

High Accuracy Radial velocity Planet Searcher (HARPS) spectrograph, ESO, La Silla Observatory**

Anglo-Australian Planet Search (APS)**,
Australia

Convection, Rotation and planetary Transits (CoRoT) (CNES/ESA, 2006-2013)*

Everything we know so far about exoplanets is built on the work of some amazing technology and missions.

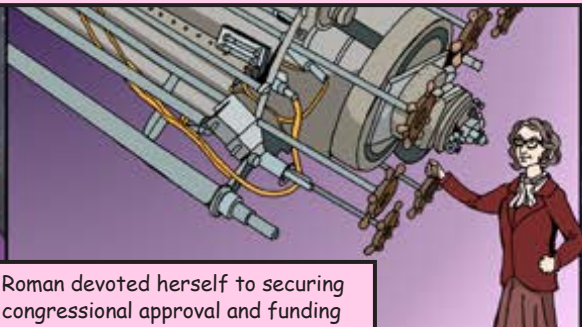
And a key figure in their development was Dr. Nancy Grace Roman (1925-2018).

When NASA was founded 1959, Roman was working as a radio astronomer at the Naval Research Laboratory.



Orbiting Solar Observatory 1
(NASA - 1962)

Six months after NASA formed, Roman was hired to help create a space-based astronomy program. By 1960, she was serving as NASA's Chief of Astronomy and Relativity.



Roman devoted herself to securing congressional approval and funding for many missions.

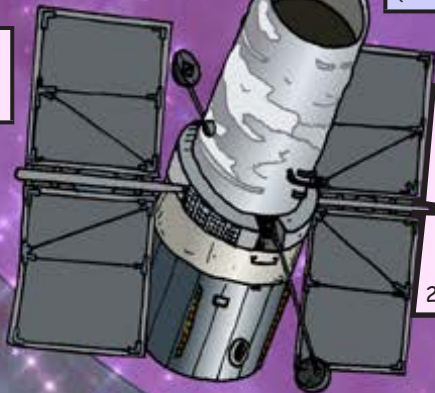
This included an ambitious mission known as the "Large Space Telescope."

Cosmic Background Explorer
(NASA, 1989-1993)

Twenty years after she started work at NASA, Congress approved the "Large Space Telescope," which would later be dubbed the Hubble Space Telescope.

Hubble launched in 1990 and is still returning valuable data after decades of work.

Hubble Space Telescope
(NASA, 1990)



Over its long lifetime, Hubble has made many important observations for exoplanet science. This includes the first atmospheric measurements of both a giant exoplanet (in 2001) and a super-Earth (in 2016). (4, 5, 6)

NASA's history in astronomy and the development of space-based telescopes are the result of the dedication of Roman and her colleagues.

Another incredibly important mission in the history of exoplanet science is NASA's Kepler space telescope. (7)

Kepler was approved in 2001 as NASA's first planet-hunting mission. Launched in 2009, it quickly made a huge impact on exoplanet science.

Kepler was fixed at observing a single spot in the sky. One tiny percent of the entire Universe.

In this one spot, Kepler found a HUGE number of planets, over 6000 candidates! And several hundred were Earth-sized.*

Kepler completed its extended mission in 2018, but for years to come, scientists will be combing through data to find even more.

Kepler has shown us that rocky planets are not rare, and that there are a lot of potentially habitable planets out there.

But Kepler's planets are very far away. The telescope could spot them, but wasn't powerful enough for astrobiologists to study them in enough detail. Bigger, more powerful telescopes would be needed**.

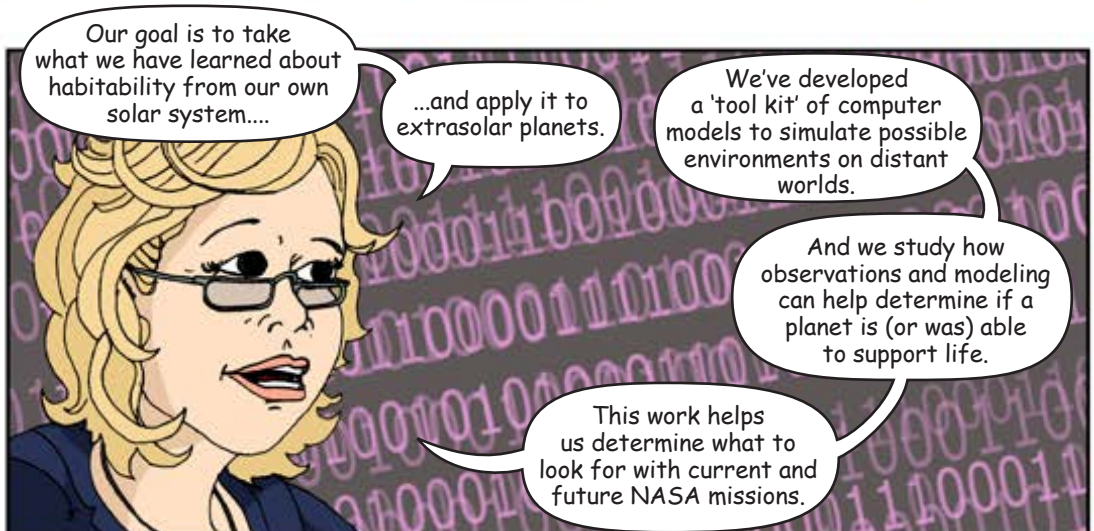
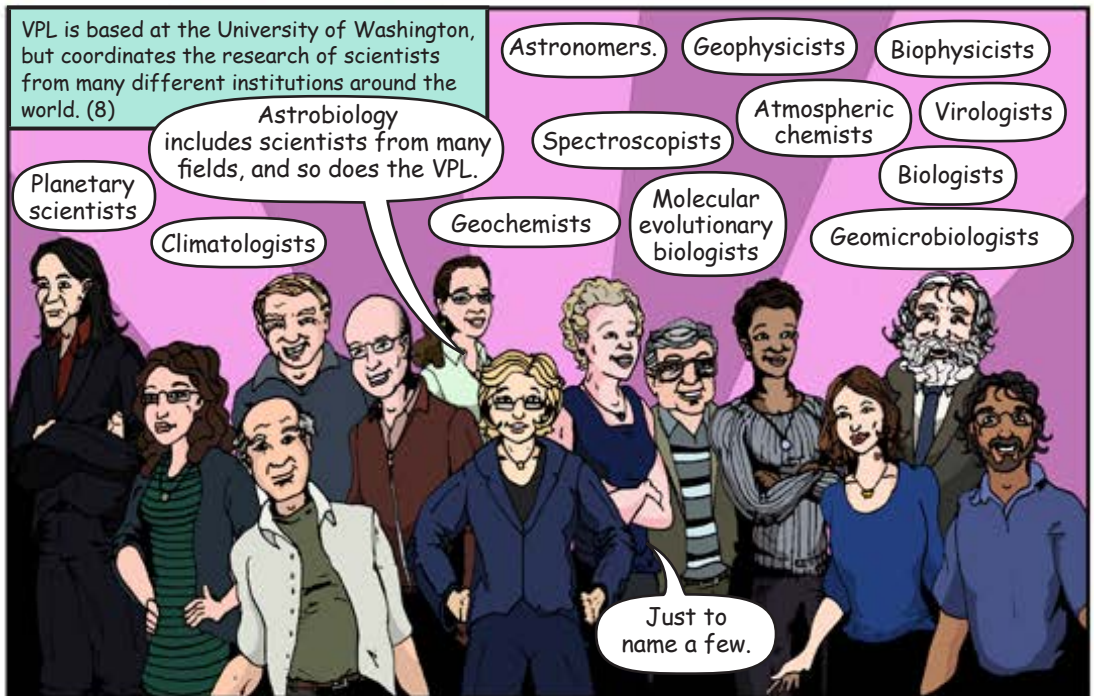
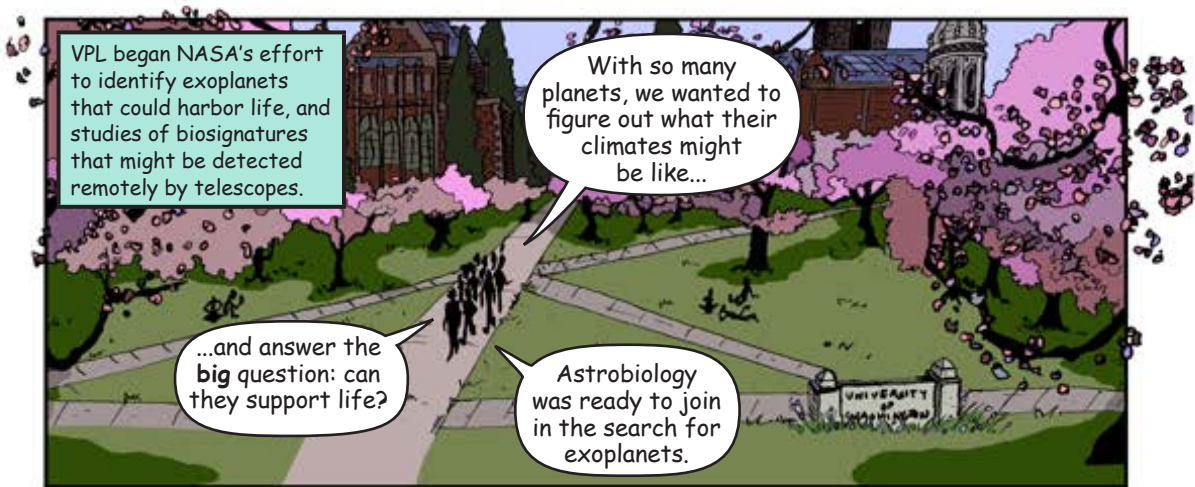
But because of Kepler, we now know that we have a good chance of finding small planets closer to home.

Astrobiologists are hoping to find such planets within 10 parsecs (roughly 32 light years) of the Earth.

At the same time NASA approved Kepler in 2001, the Virtual Planetary Laboratory (VPL) was founded with support from NASA Astrobiology.

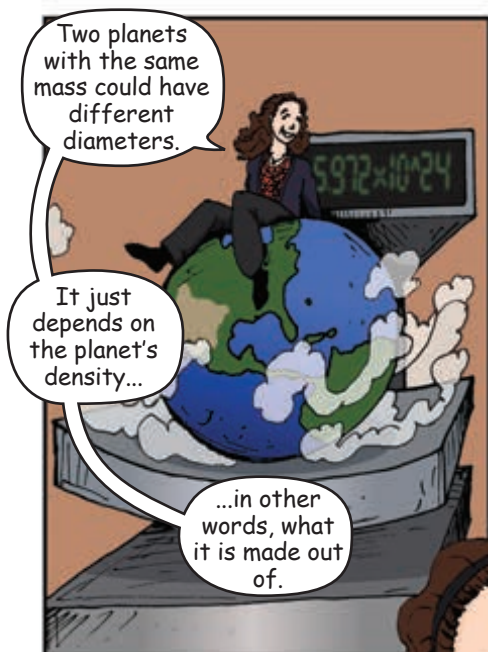
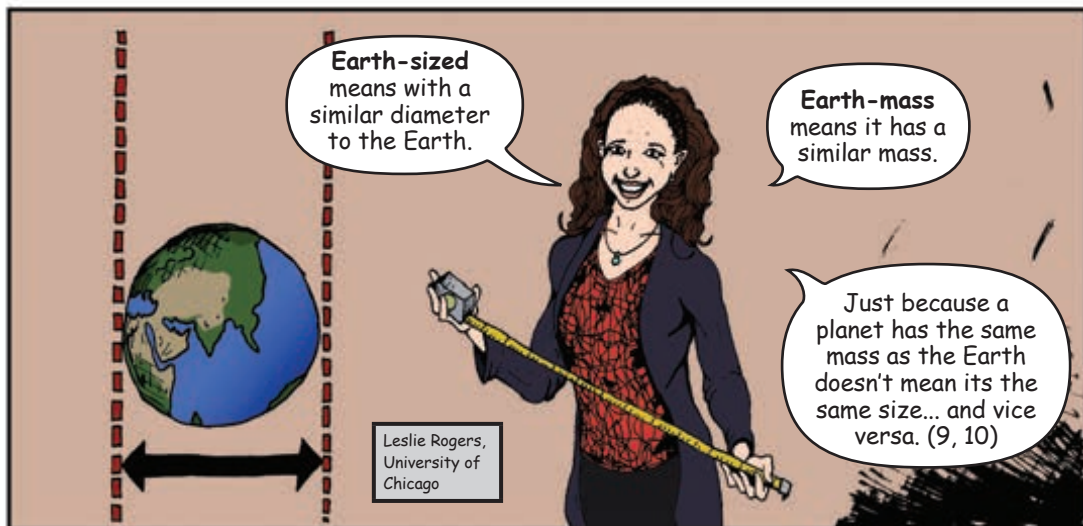
*after a 'candidate' is found, more observations are made over time to confirm that it's actually a planet.

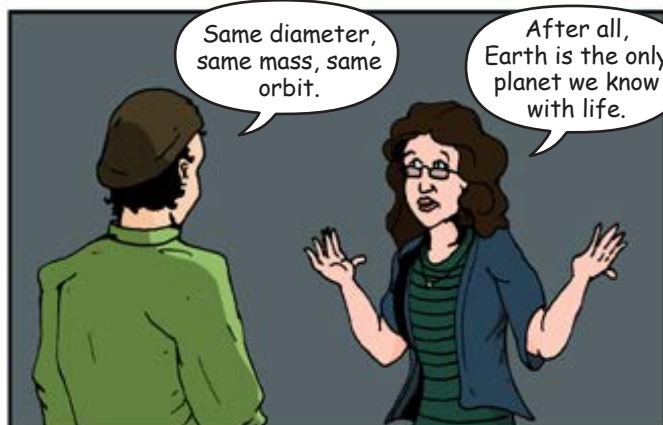
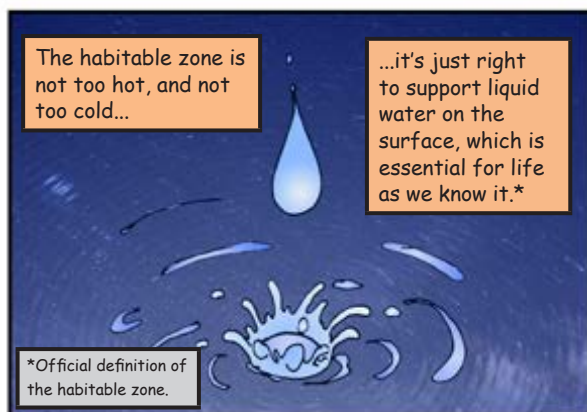
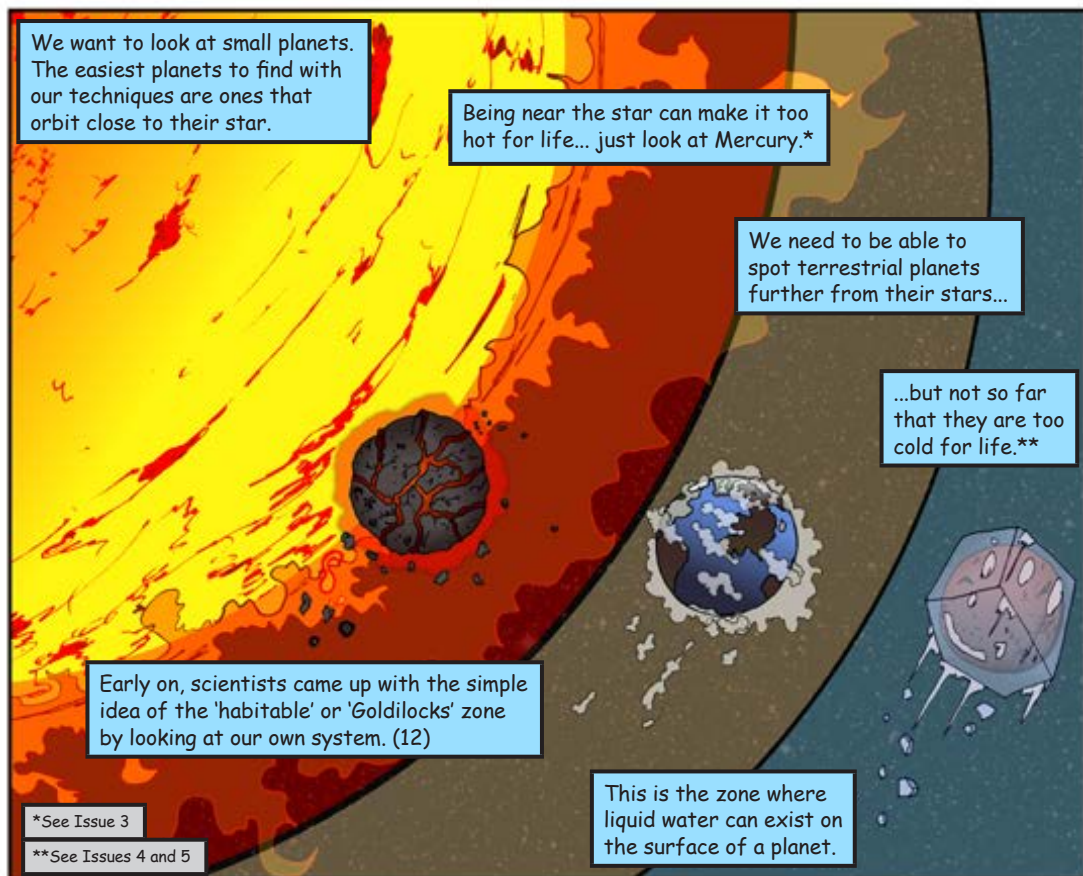
**More on that later.

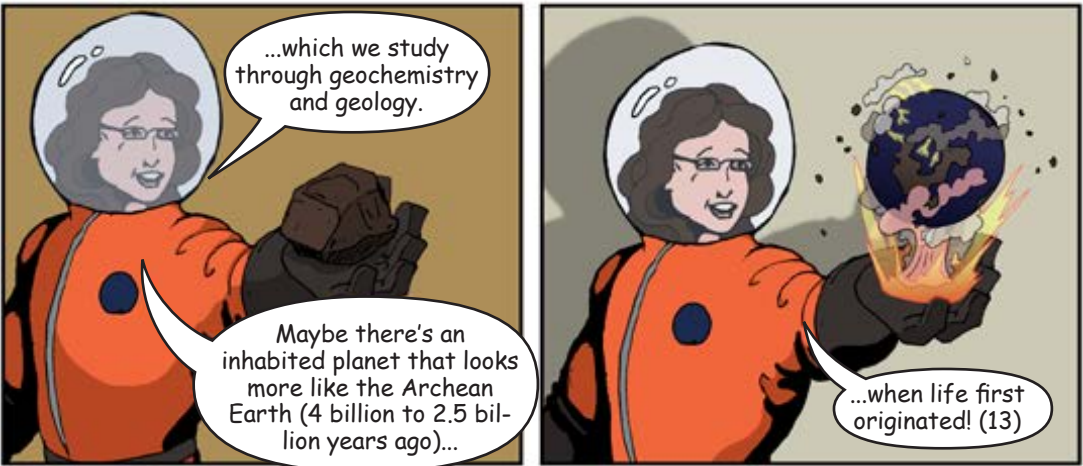
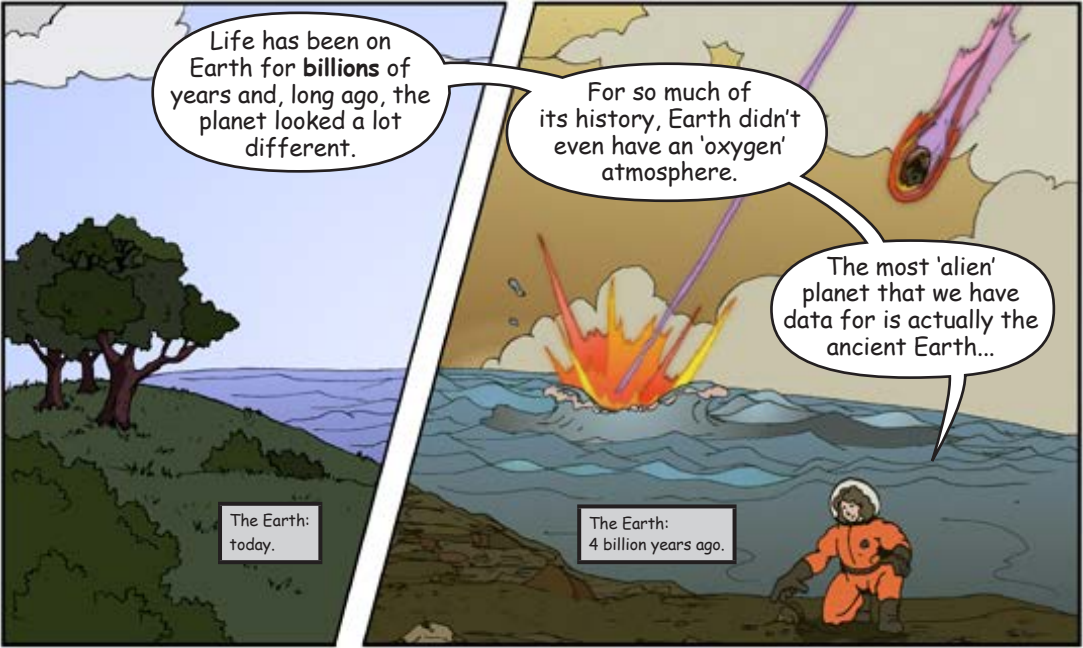


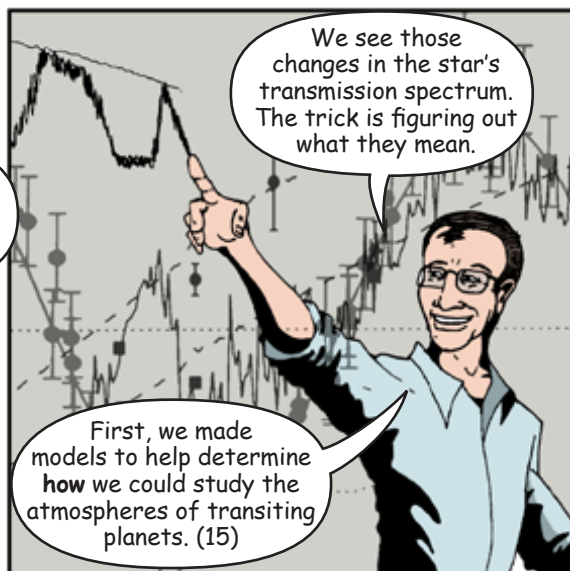
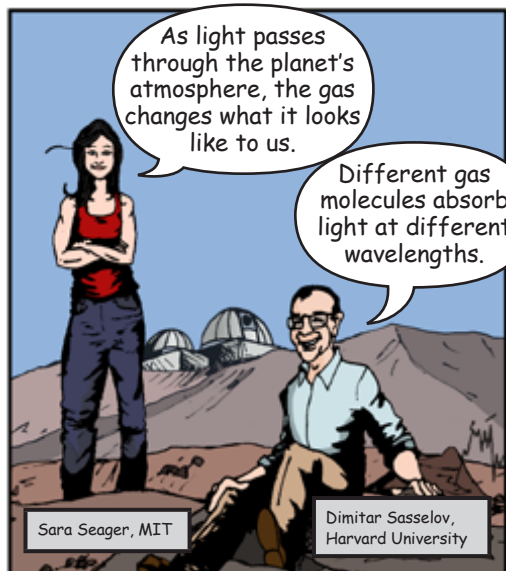
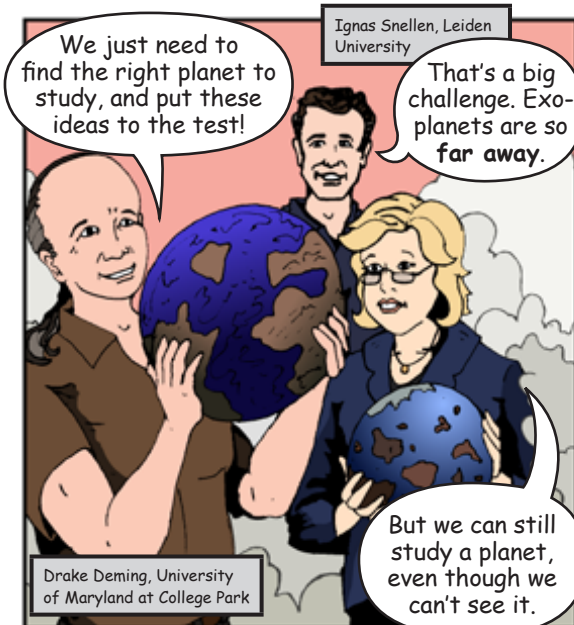


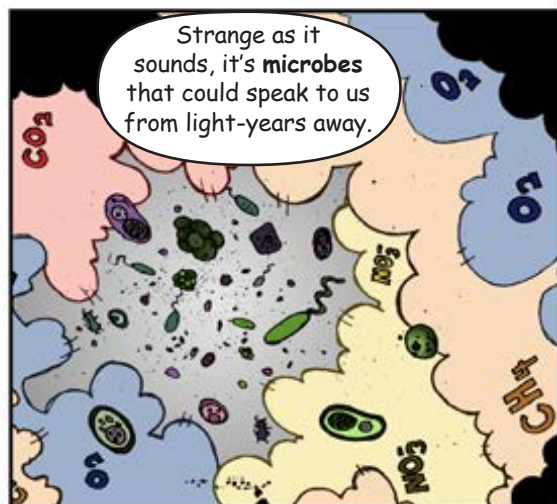
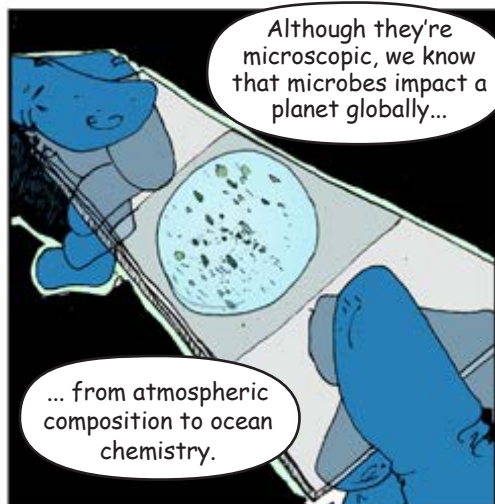
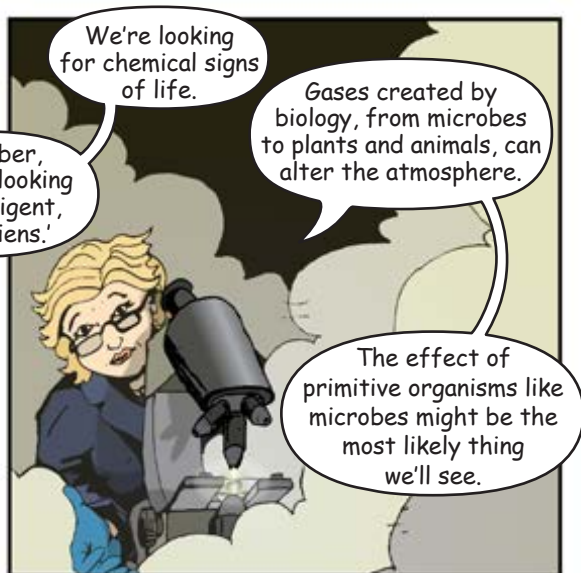
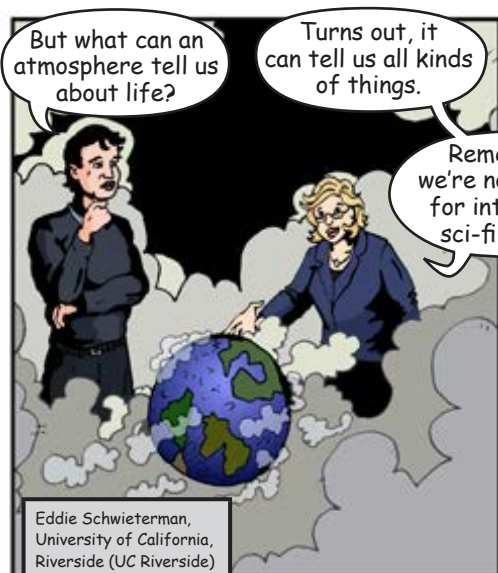
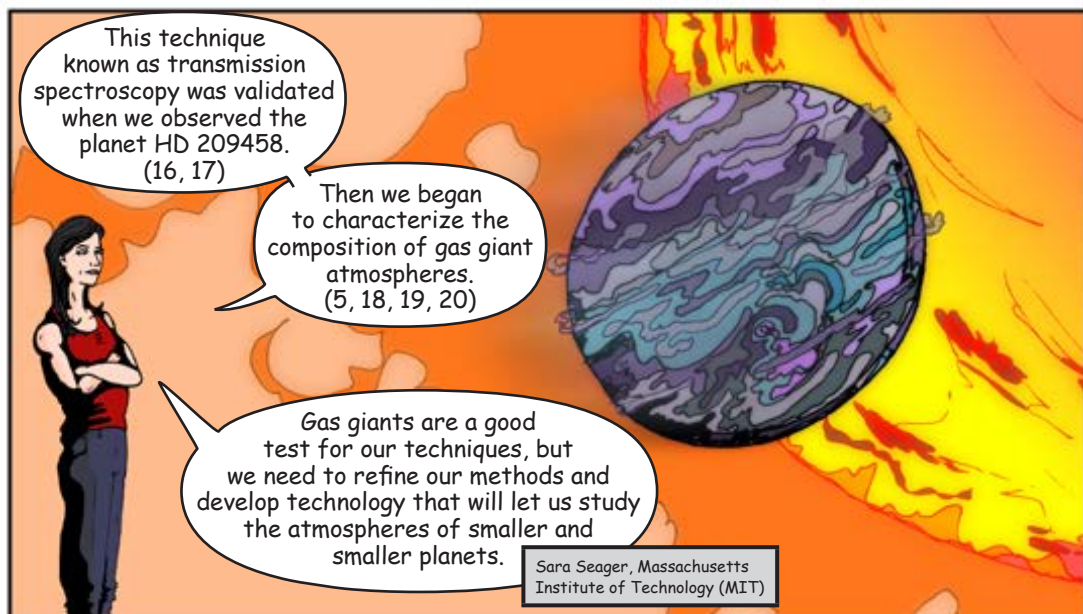
**See Issues 2,3 & 5











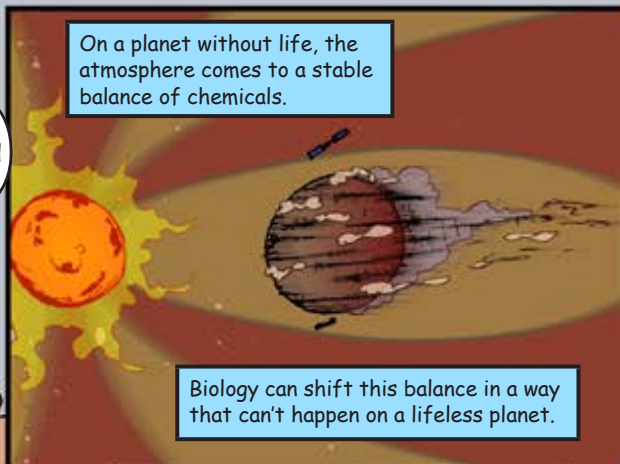
It was James Lovelock* who first suggested that biology could be detected by its effect on the atmosphere.



If Earth had no life, the atmosphere would look different. (21)

*See Issue 2

On a planet without life, the atmosphere comes to a stable balance of chemicals.



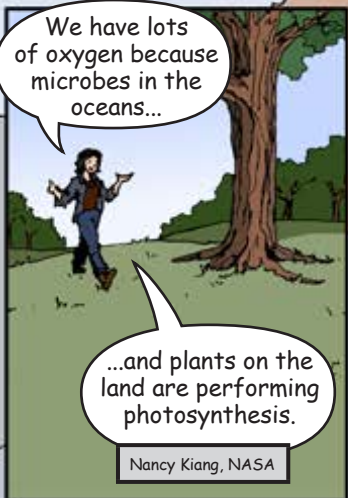
Biology can shift this balance in a way that can't happen on a lifeless planet.

Some kinds of chemicals don't co-exist easily in the atmosphere because of chemical reactions.

For example, oxygen and methane don't normally co-exist easily. But on Earth today, they do...

...because of biology!

We have lots of oxygen because microbes in the oceans...



...and plants on the land are performing photosynthesis.

Nancy Kiang, NASA

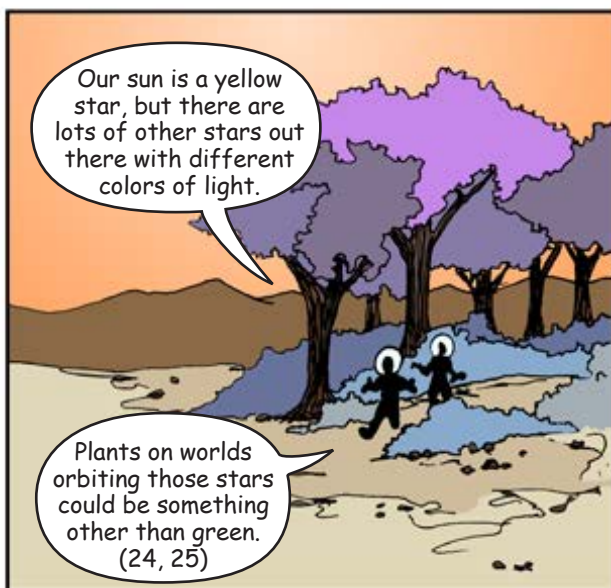
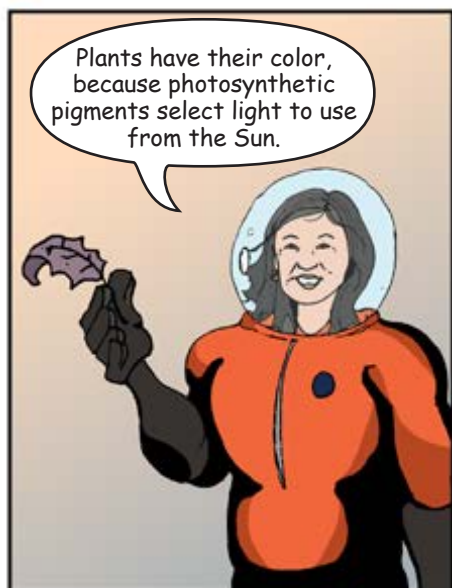
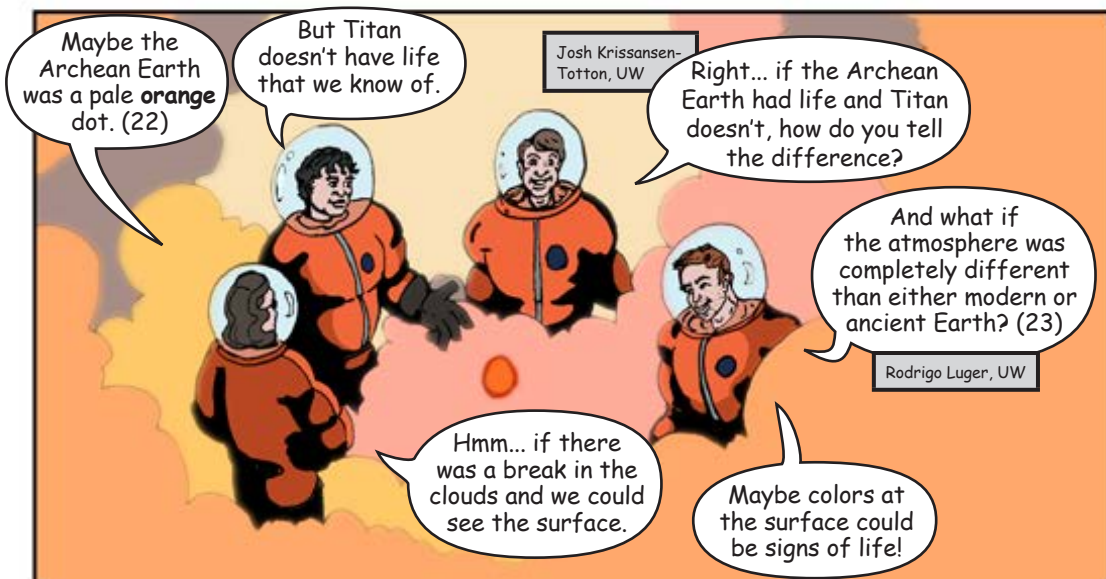
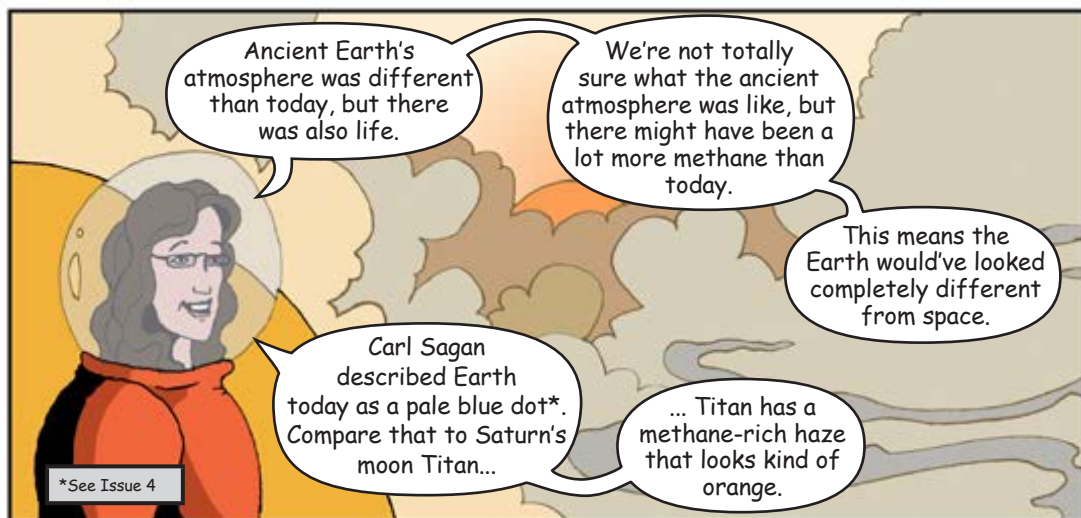
Photosynthesis makes the oxygen, and it also makes cool colors that can be seen from space!

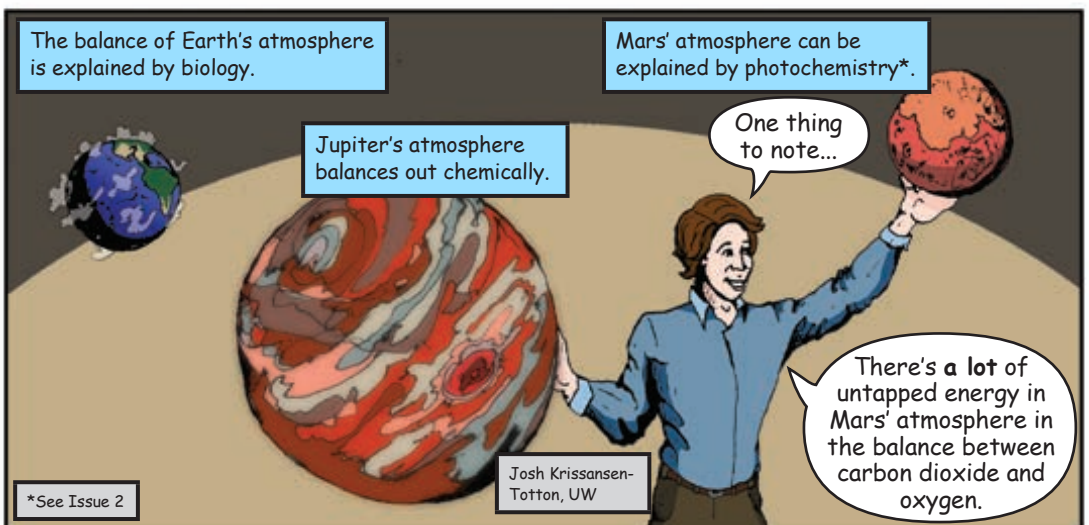
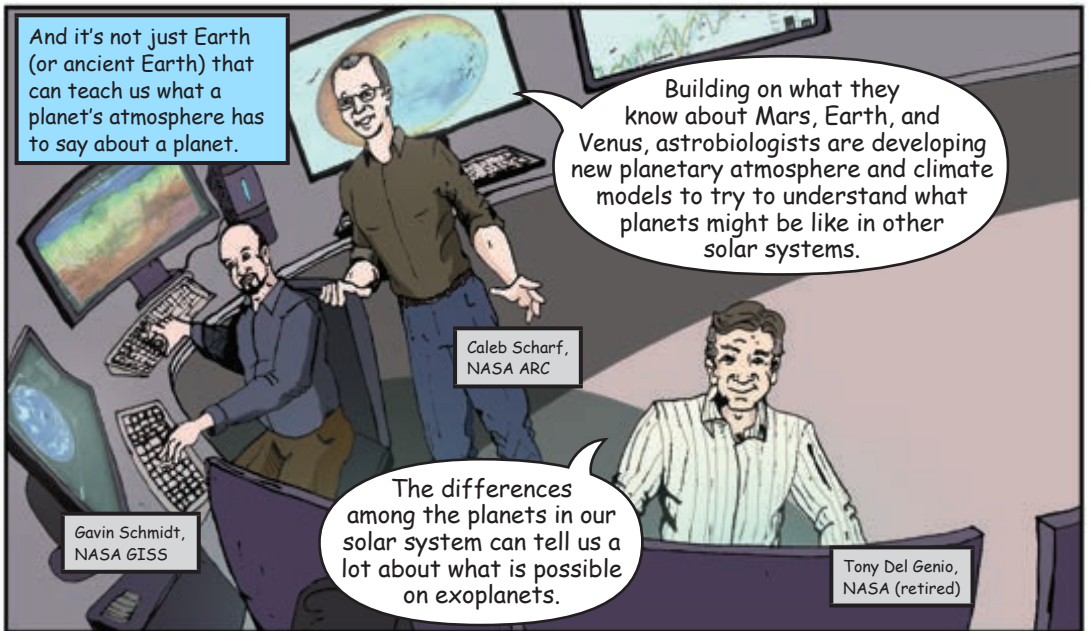
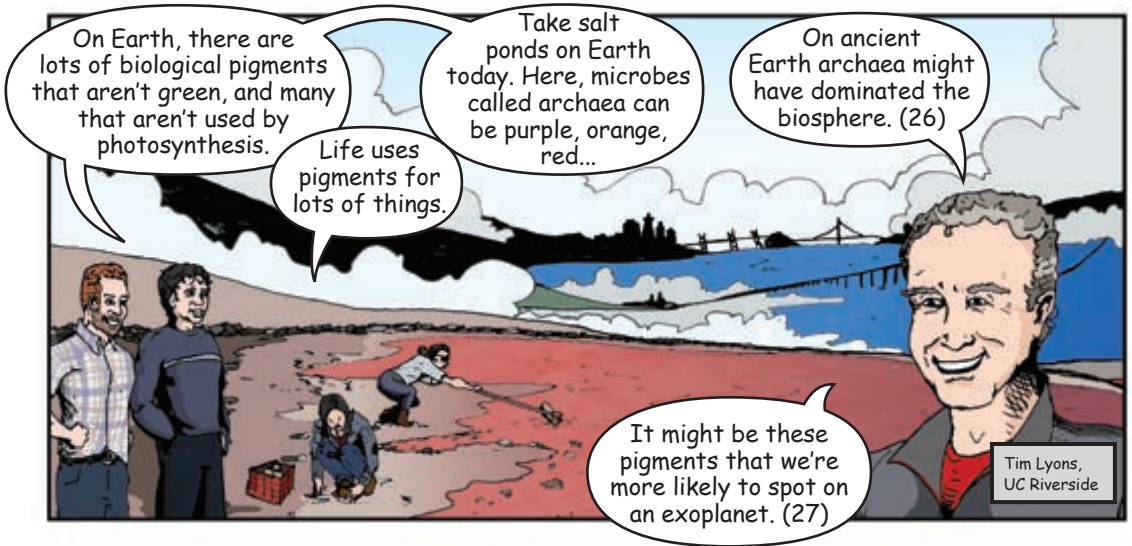


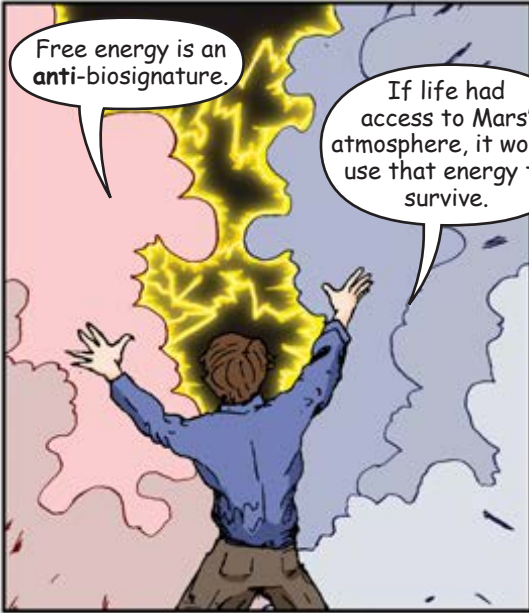
It's especially easy to see things like forests and grasslands covering the land.



* See Issue 5

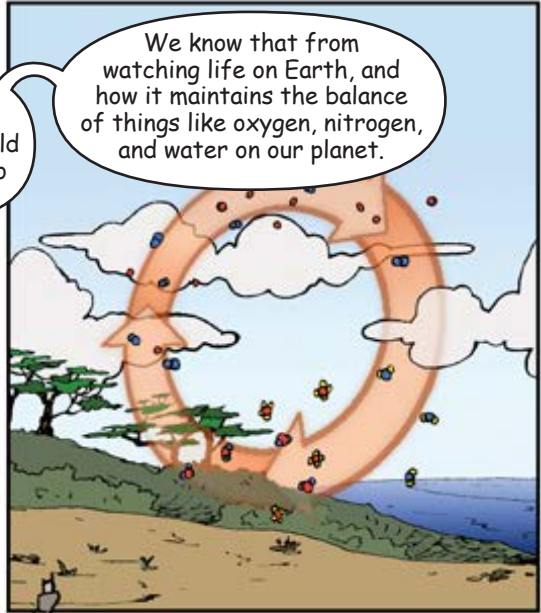






Free energy is an **anti-biosignature**.

If life had access to Mars' atmosphere, it would use that energy to survive.



We know that from watching life on Earth, and how it maintains the balance of things like oxygen, nitrogen, and water on our planet.



Life makes the composition of Earth's atmosphere possible.

"The low oxygen on Mars is what made me think the planet was lifeless back in the days of Viking."
(21)*

*See Issue 2



Once again, it seemed so easy at first...

...look for something small...

...in the habitable zone...

...and an **oxygenated** atmosphere!



But, of course, we realized there was more to it than that.

Ravi Kumar Kopparapu,
NASA GSFC



At a coffee shop
in Seattle, 2010...

Uh-oh!...



What, is
something
wrong?

Uh... yeah.

More
coffee?

No, not
that...

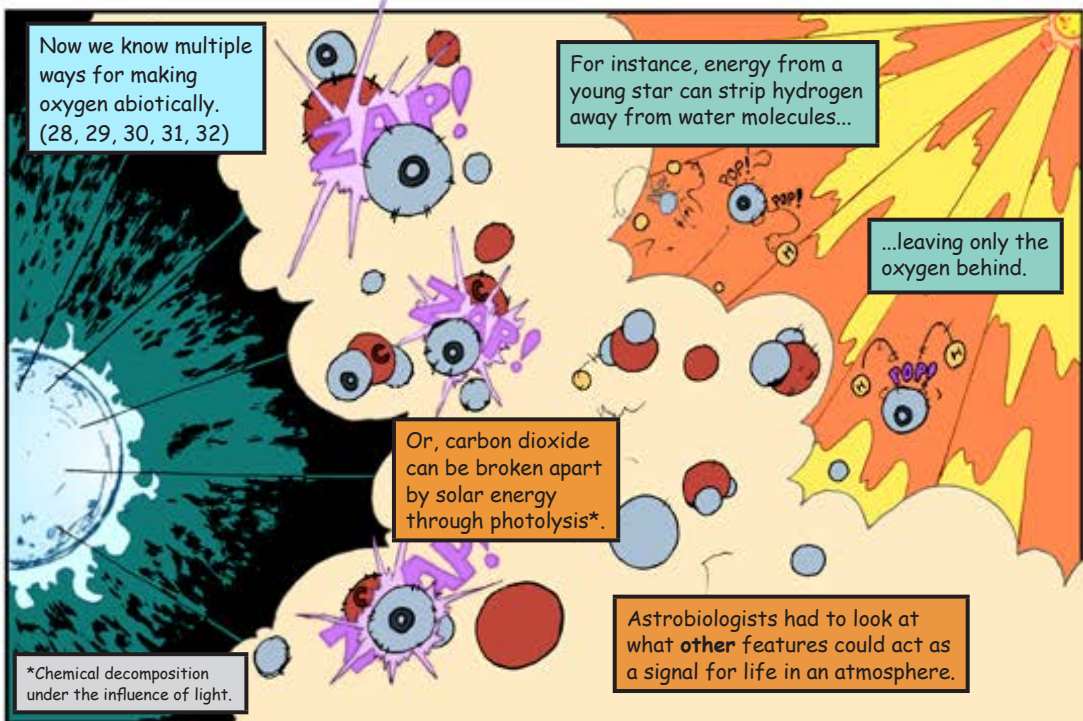
Jen Domagal-Goldman,
American Democracy Project



Oxygen
doesn't have
to come from
life...

Turns out,
you can get lots
of oxygen without
biology. (28)

That would be a
false positive!



Now we know multiple
ways for making
oxygen abiotically.
(28, 29, 30, 31, 32)

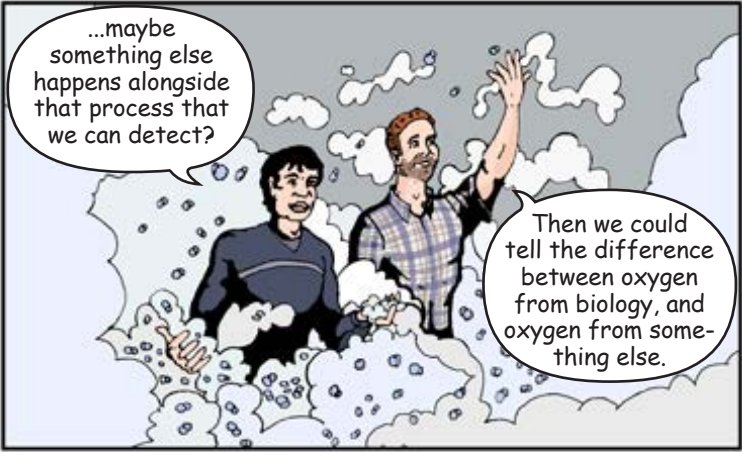
For instance, energy from a
young star can strip hydrogen
away from water molecules...

...leaving only the
oxygen behind.

Or, carbon dioxide
can be broken apart
by solar energy
through photolysis*.

Astrobiologists had to look at
what **other** features could act as
a signal for life in an atmosphere.

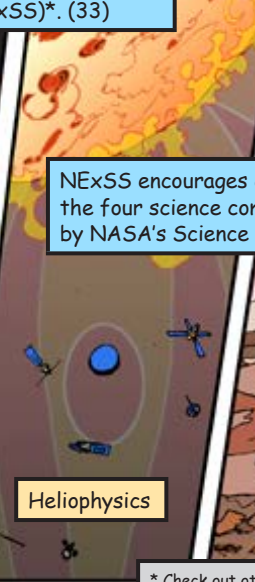
*Chemical decomposition
under the influence of light.



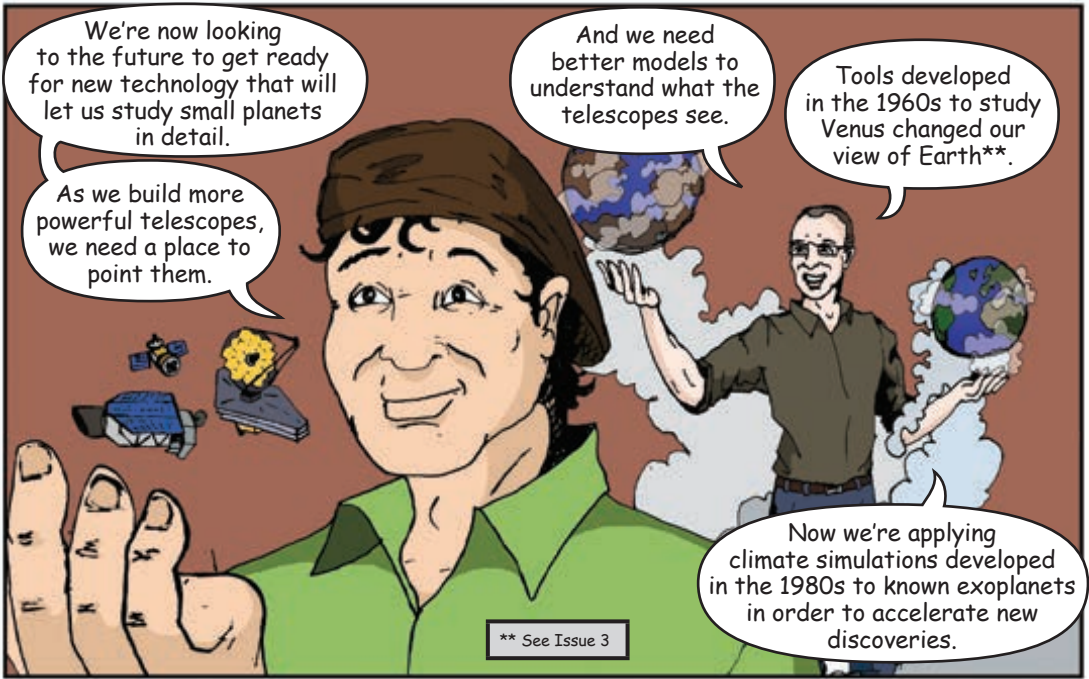
In 2014, NASA Astrobiology led an effort to create a new, coordinated network of exoplanet explorers.

It was dubbed the Nexus for Exoplanet System Science (NExSS)*. (33)

NExSS encourages collaboration among the four science communities supported by NASA's Science Mission Directorate.



* Check out other NASA Research Coordination Networks (RCNs) in other issues!



NExSS includes several research teams, each dedicated to solving a piece of the exoplanet puzzle.



With new ideas from even more scientists, let's look again at the 'habitable zone.'



The definition of "habitable zone" is the distance from a star at which liquid water could exist on surfaces of orbiting planets. (34, 35)

The 'habitable zone' makes things sound easy. We just need a small planet with an orbit similar to the Earth.

If only it were that simple...

Antígona Segura Peralta,
Universidad Nacional
Autónoma de México



The habitable zone is based on an Earth-like planet, with an Earth-like atmosphere.

As soon as a planet deviates from that composition, it can get into trouble.

Mars is at the outer edge of our system's habitable zone.

Earth sits in the middle.

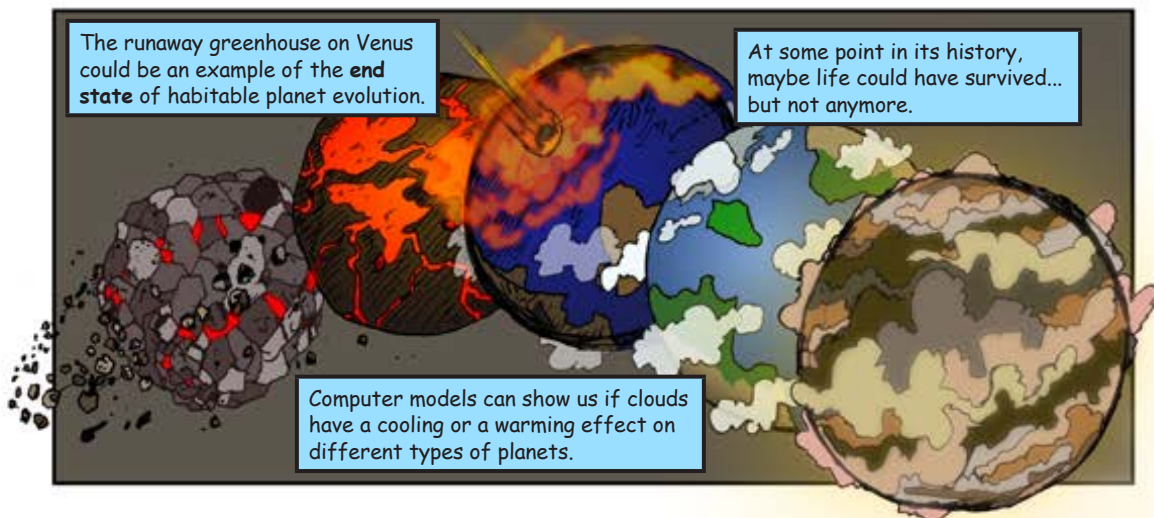
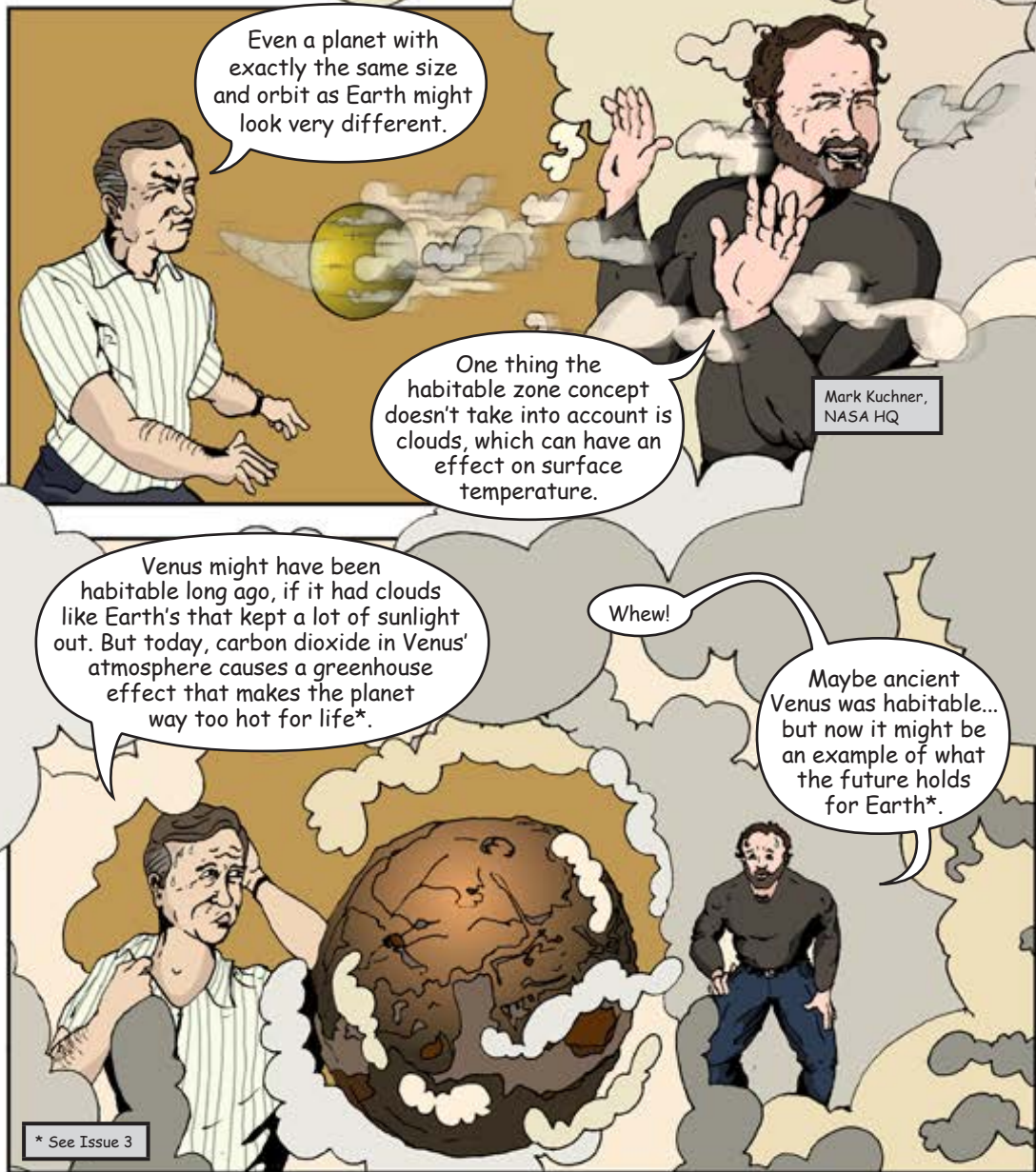
Venus is outside the habitable zone, but it's close to the inner limit.

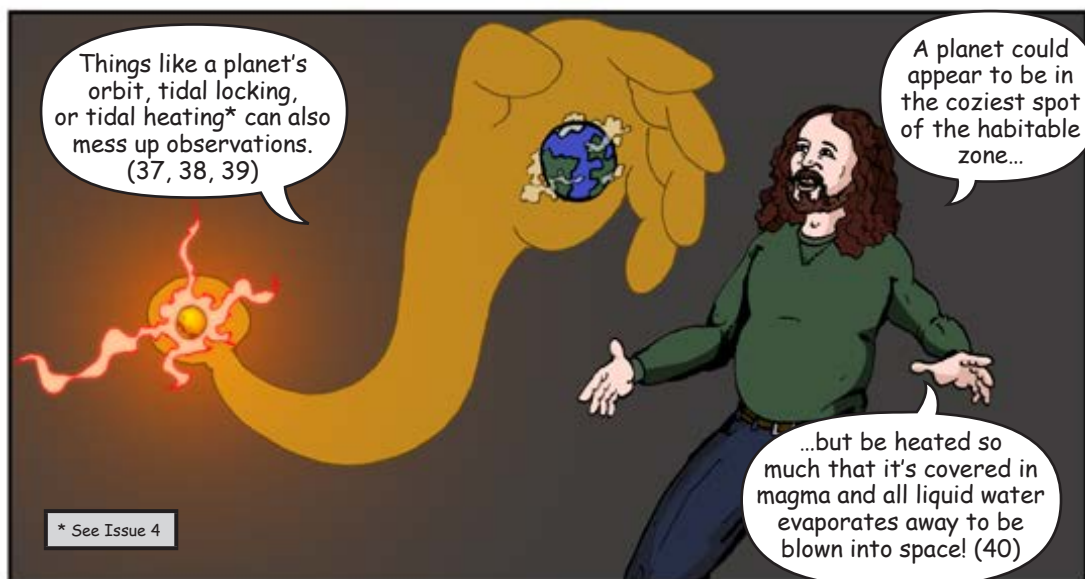
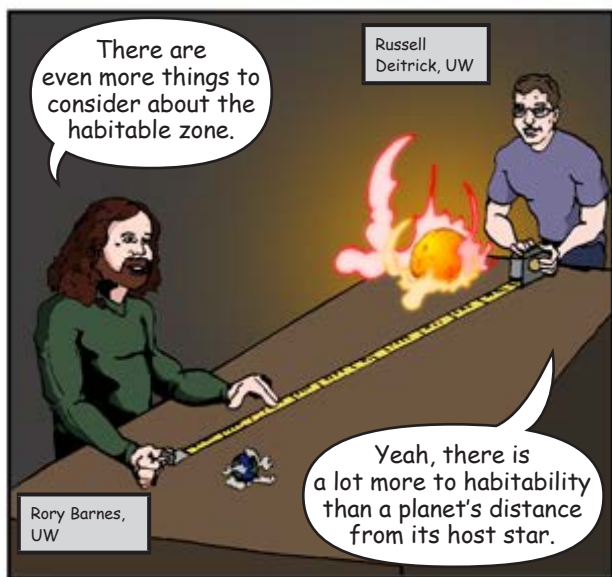
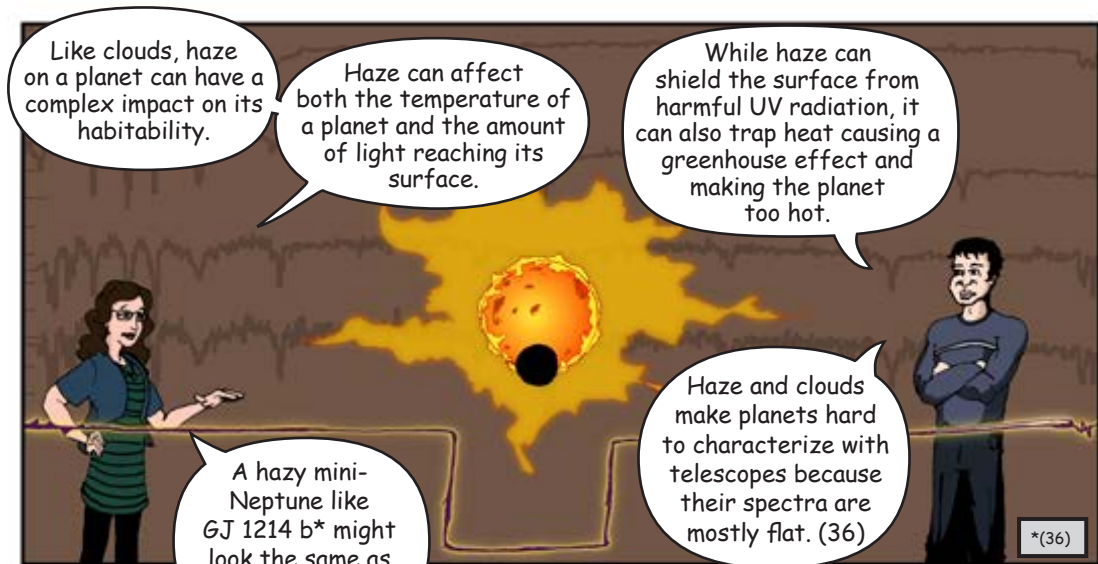
So far, we haven't found evidence of past or present life on either Mars or Venus.

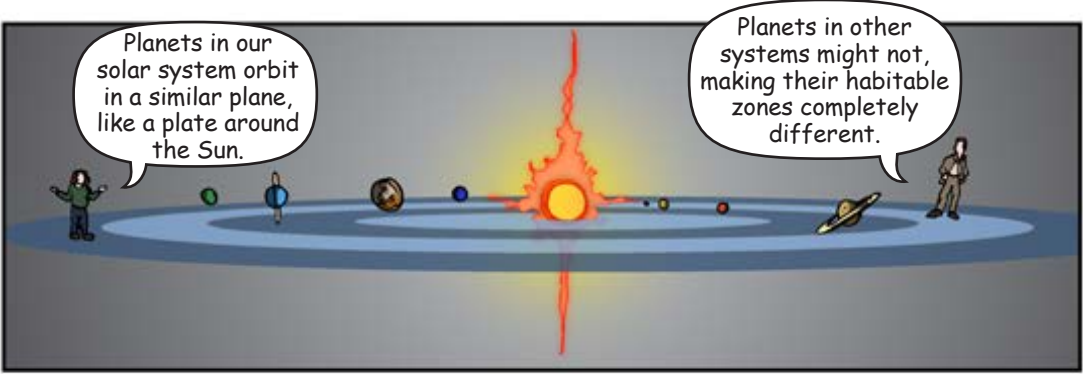
And what about the Moon? It sits within our system's habitable zone, yet it isn't habitable.

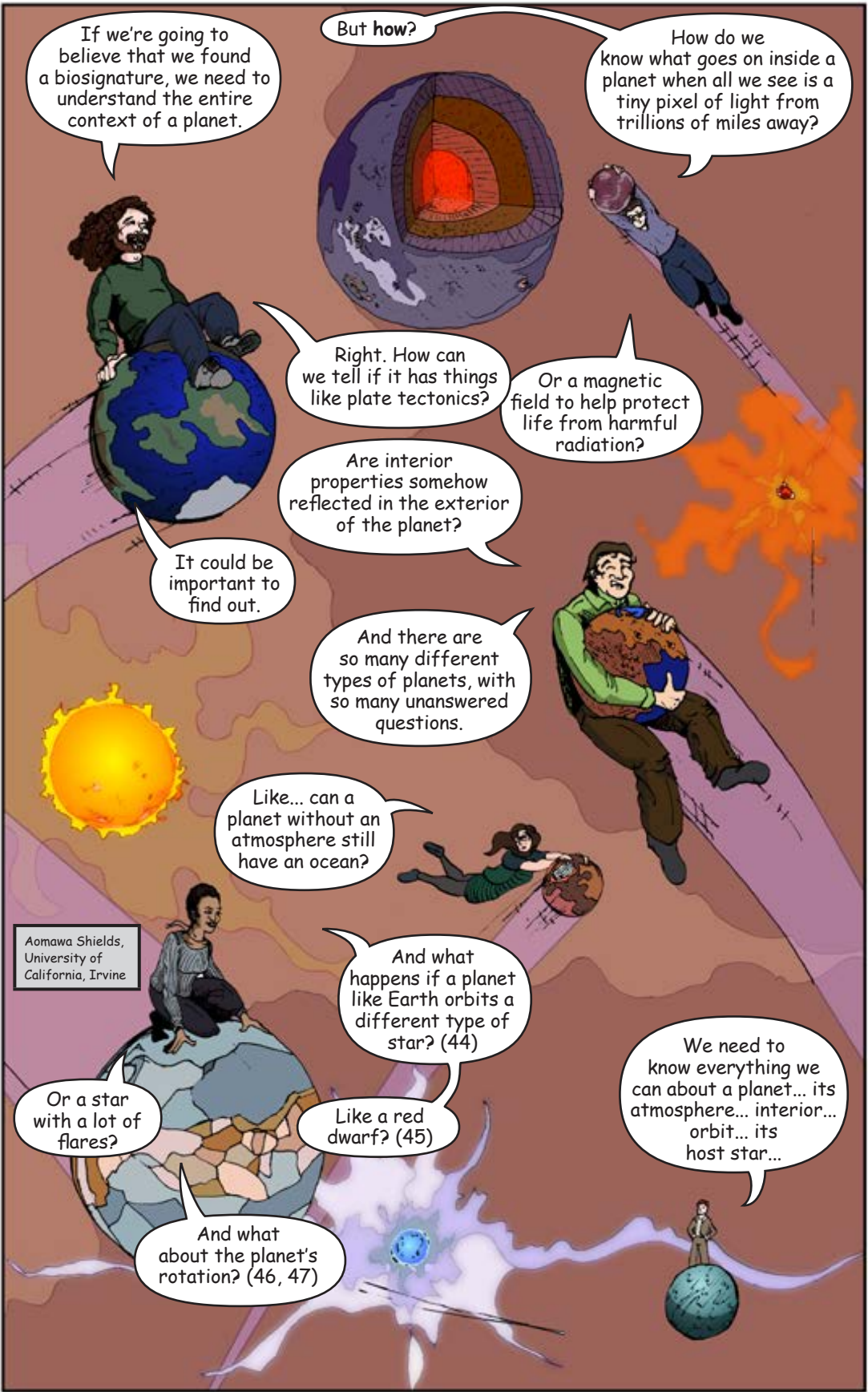
So you see, there's more to it than spotting a world in the habitable zone.

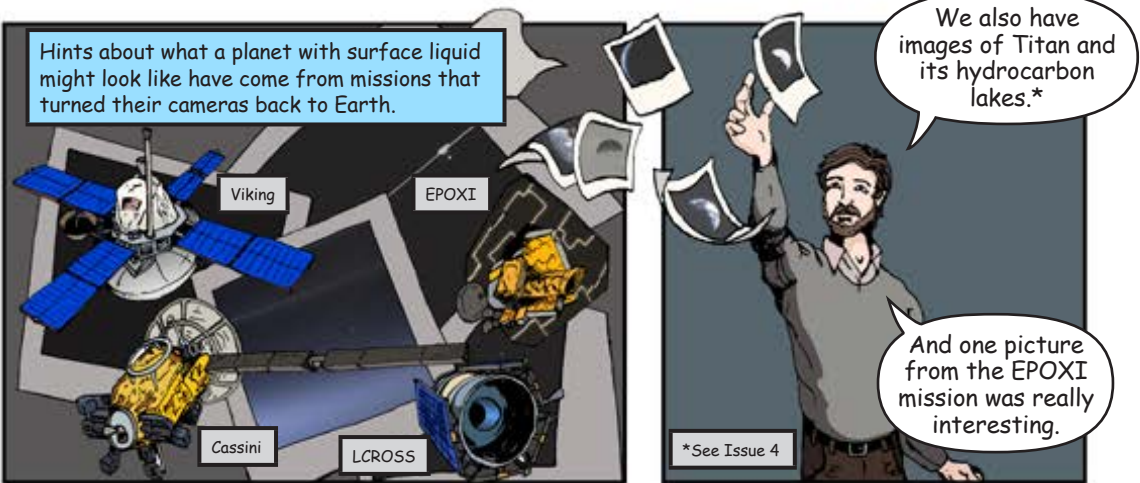
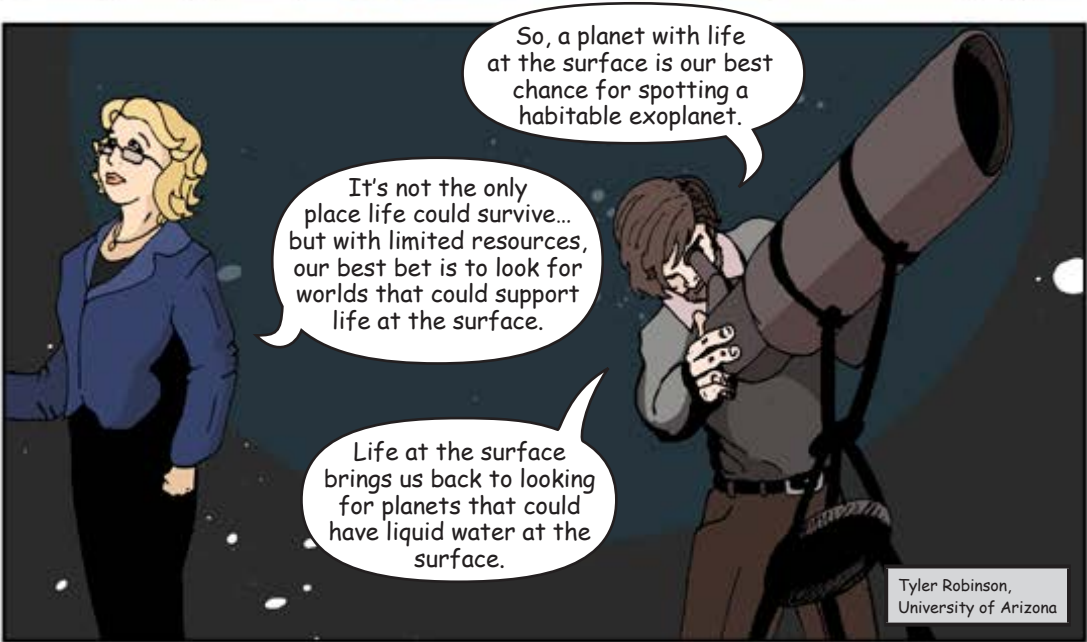
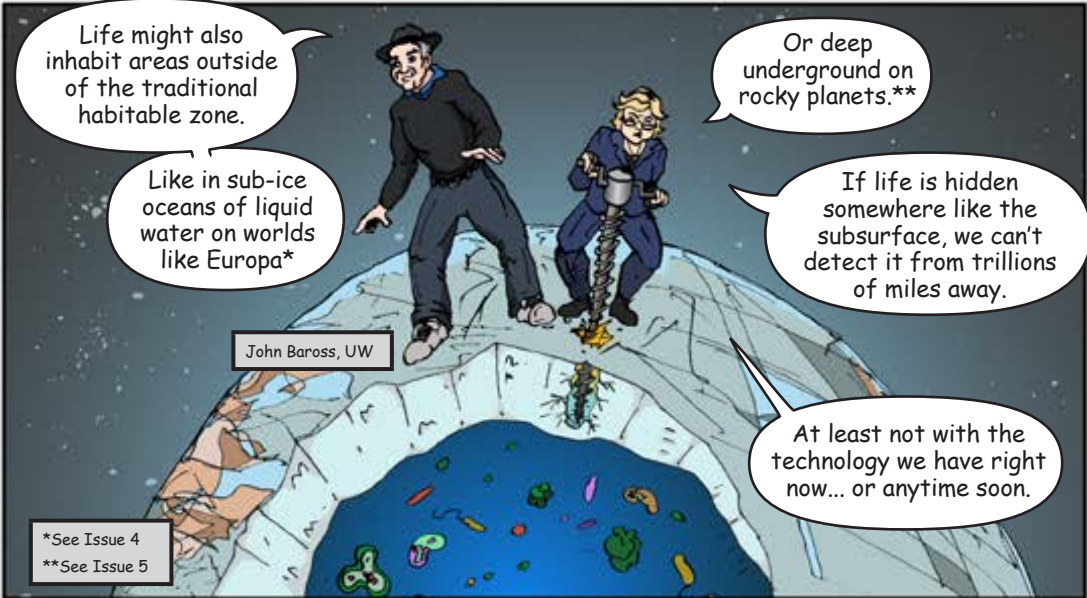


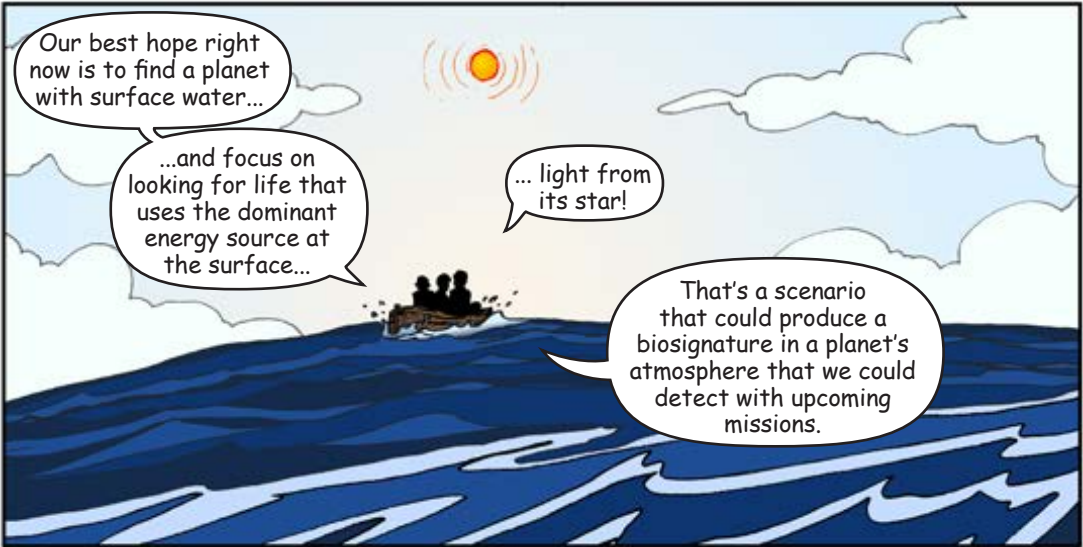
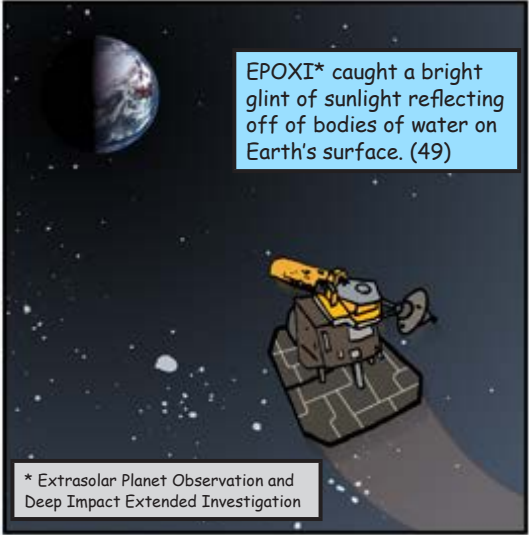
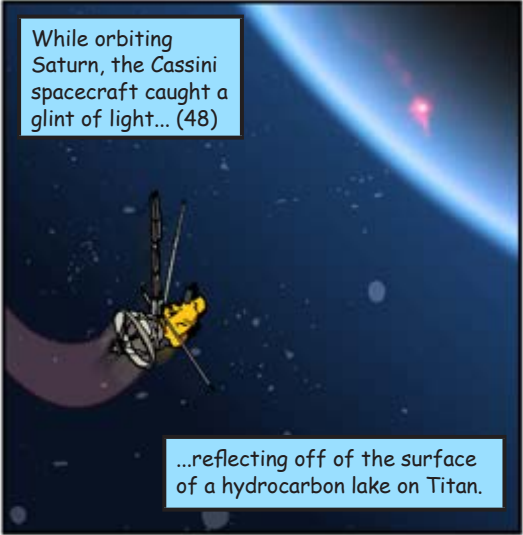


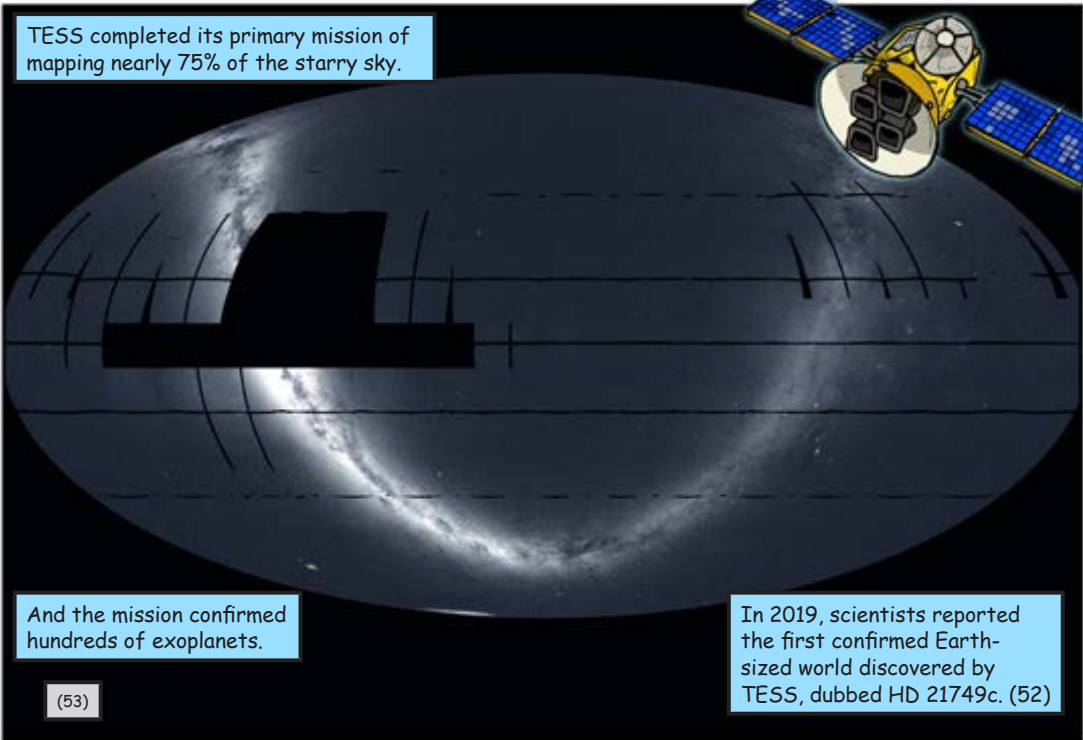
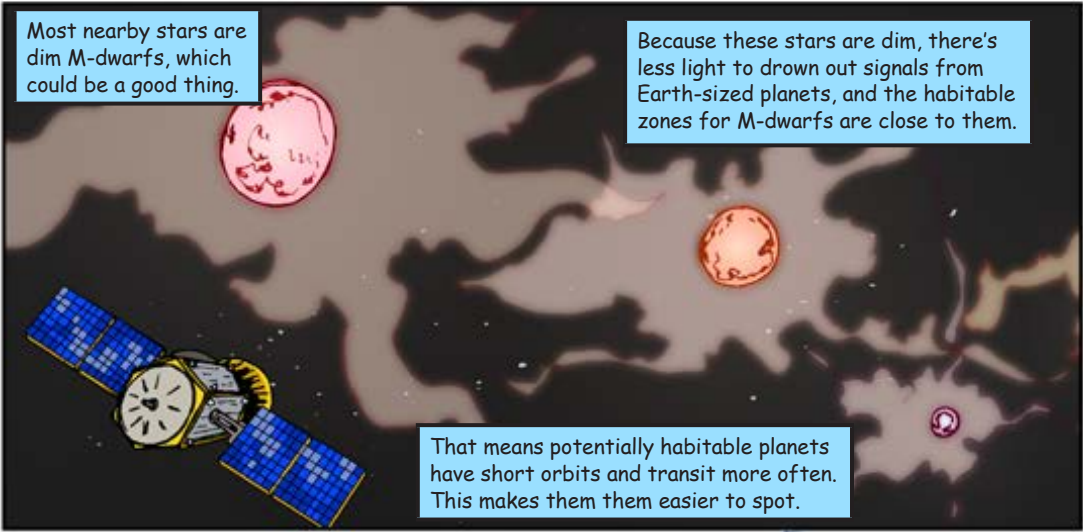
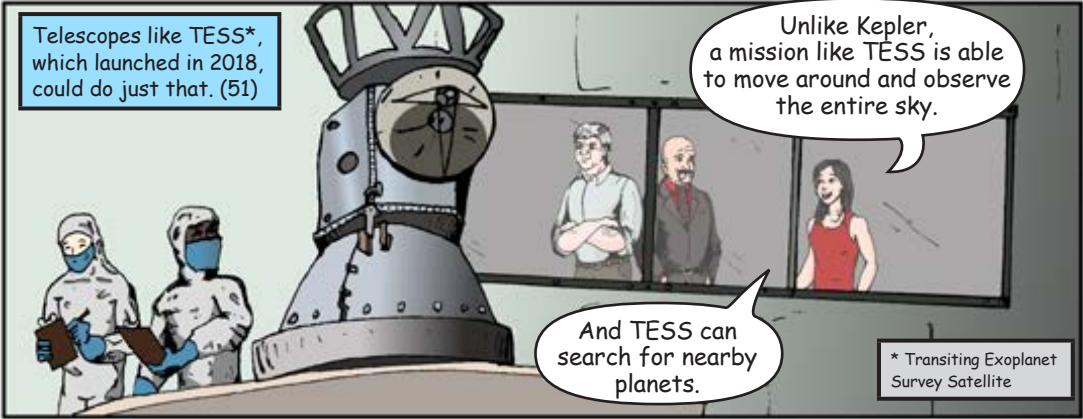


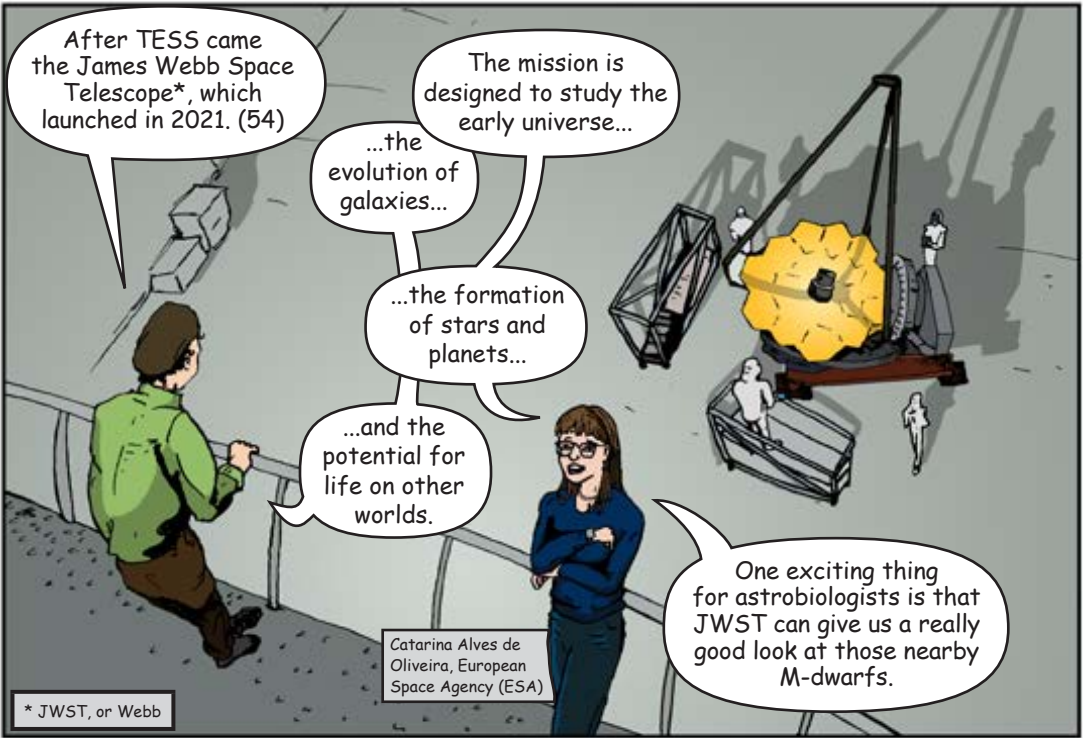












But Webb does have limitations in terms of studying planets.

Importantly, the transmission spectra we get from Webb can only tell us about the uppermost part of the atmosphere.

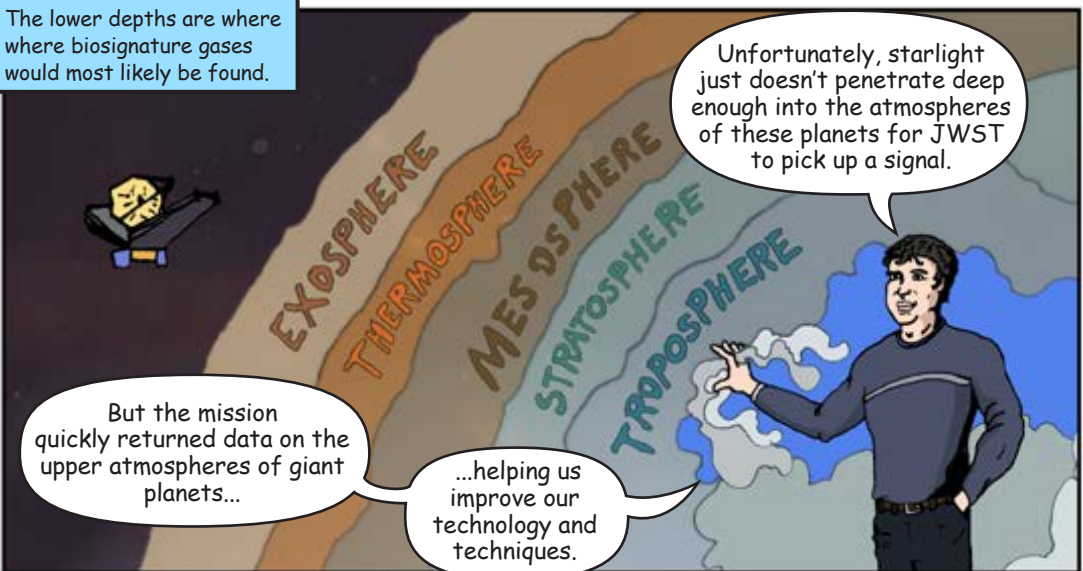
And it's the lower levels of the atmosphere that interest astrobiologists most.

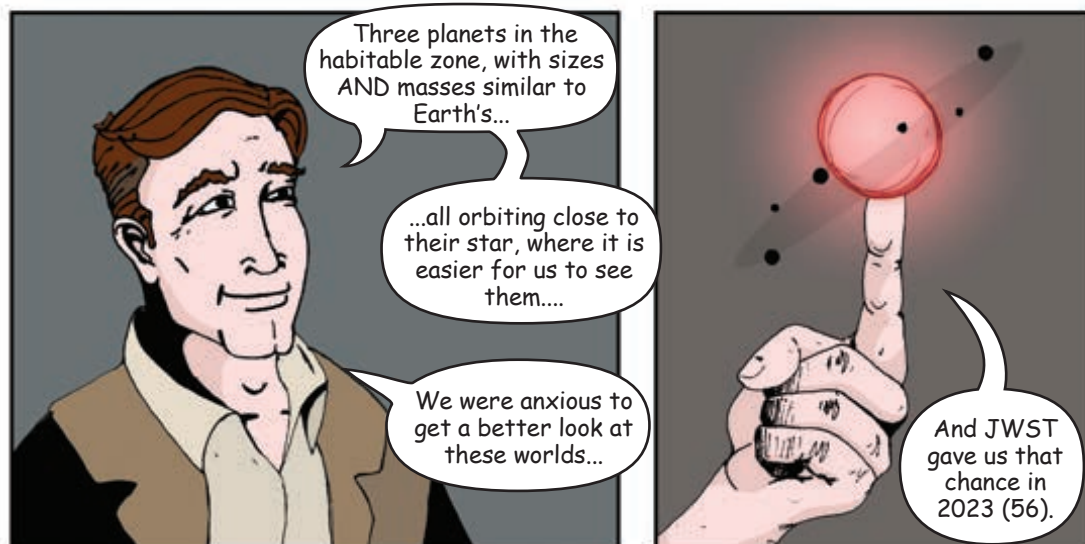
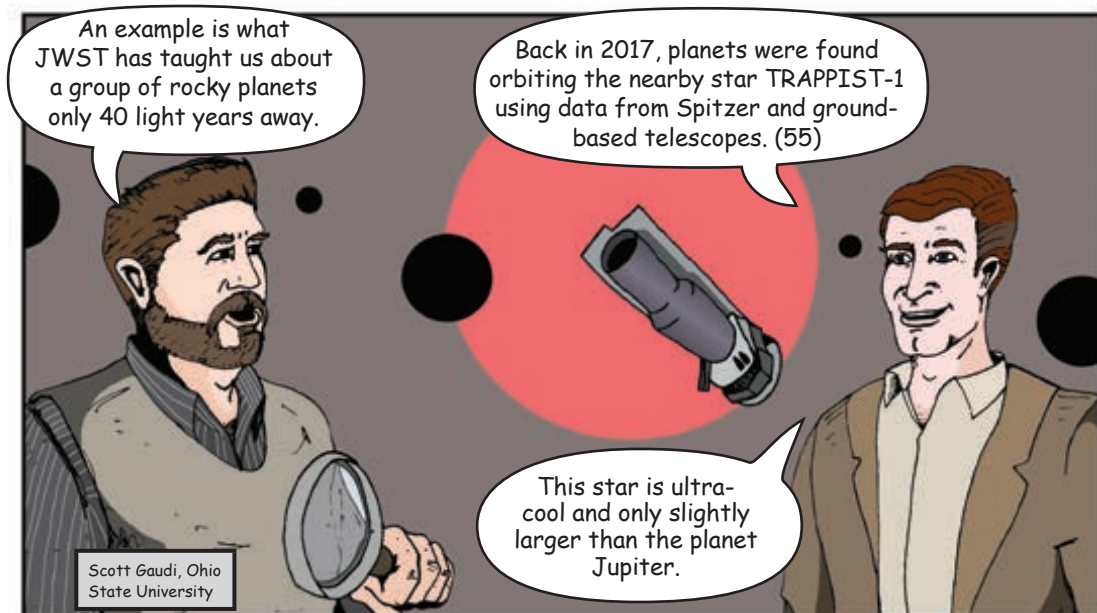
The lower depths are where where biosignature gases would most likely be found.

Unfortunately, starlight just doesn't penetrate deep enough into the atmospheres of these planets for JWST to pick up a signal.

But the mission quickly returned data on the upper atmospheres of giant planets...

...helping us improve our technology and techniques.





The mission was able to measure the temperature of the planet Trappist-1b based on its thermal emission. (57)

Some scientists had speculated that the planet Trappist-1c might have a thick CO₂ atmosphere, but data showed that this wasn't the case. (58)

Aki Roberge,
NASA GSFC

But other studies still leave hope for water vapor or oxygen. (59)

Michaël Gillon,
Université de Liège

To study a small planet, the telescope will have to observe the planet transiting its star many, many times.

Dawn Gelino,
NASA's Jet
Propulsion
Laboratory (JPL)

And getting data from a single planet in this way is going to take a HUGE amount of observation time.

This is why we'll only be able to study a few planets during the mission's lifetime.

Many people need to use the telescope, and it's not just exoplanets that JWST is studying. (60, 61)

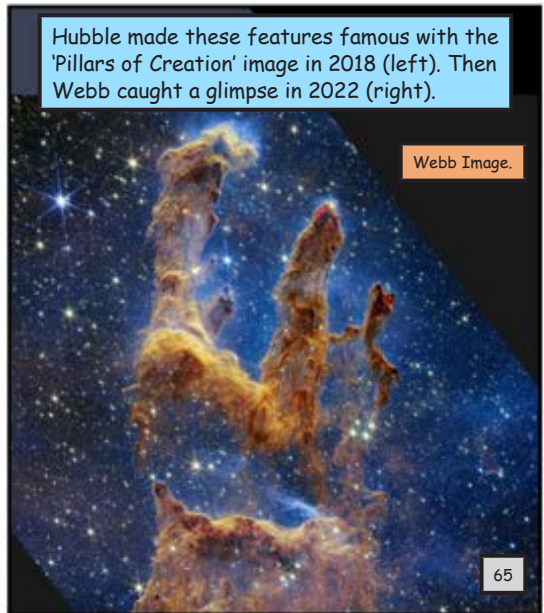
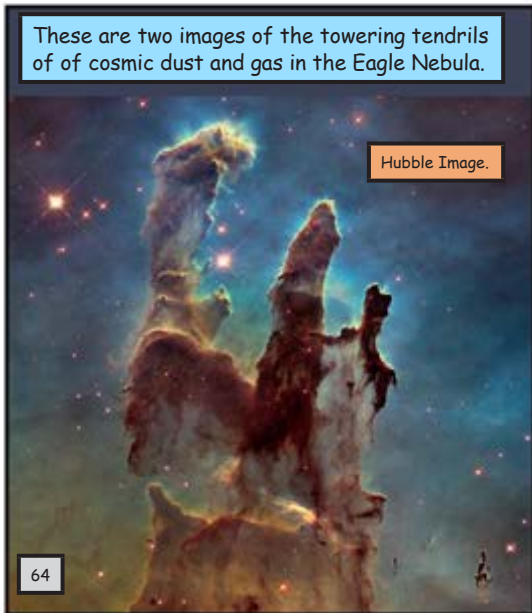
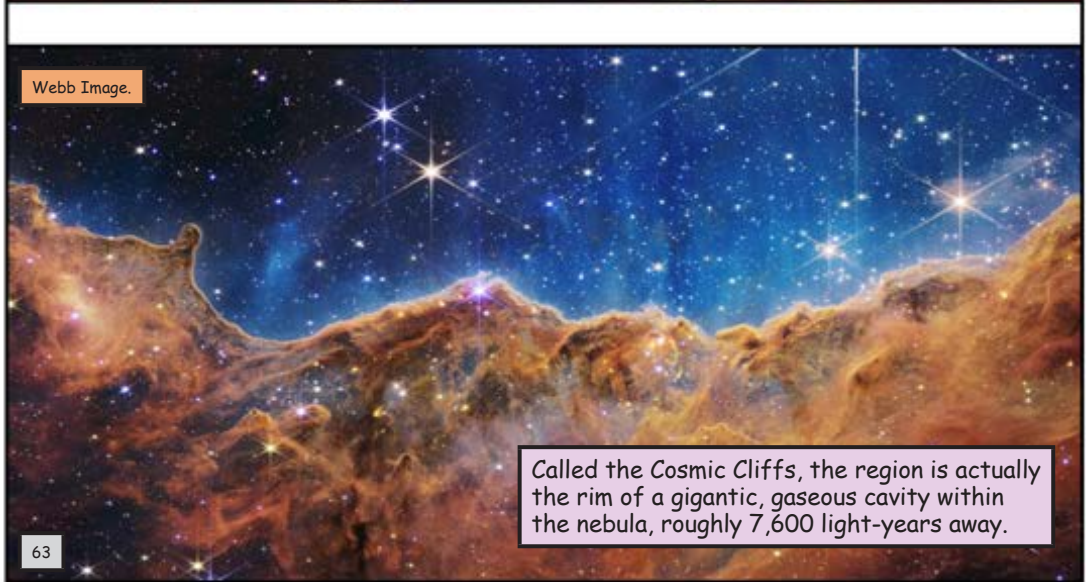
So we have to carefully choose the best planets to look at.

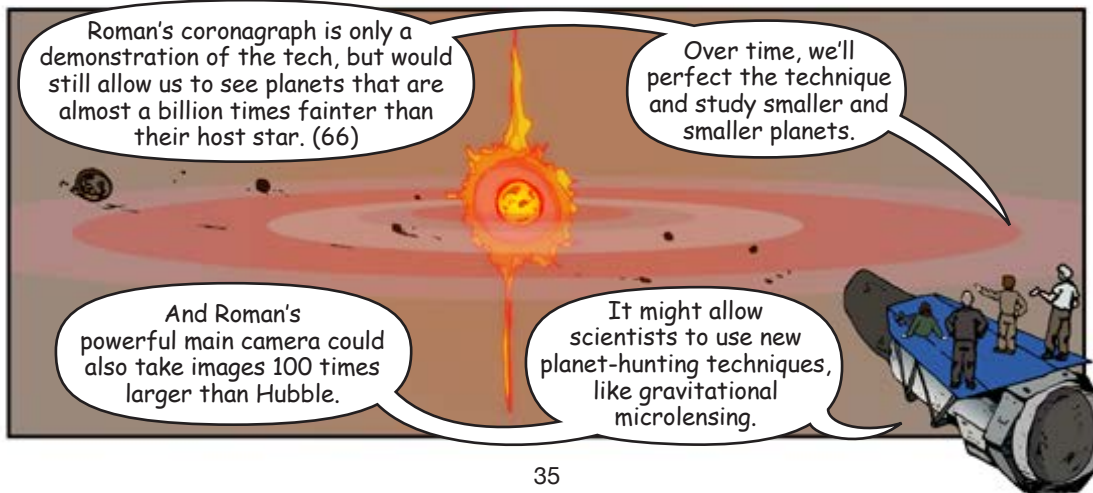
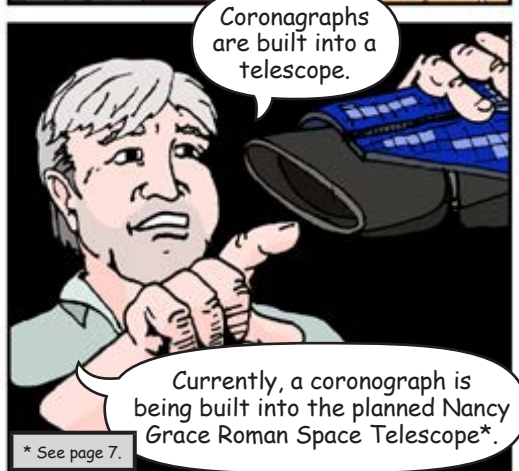
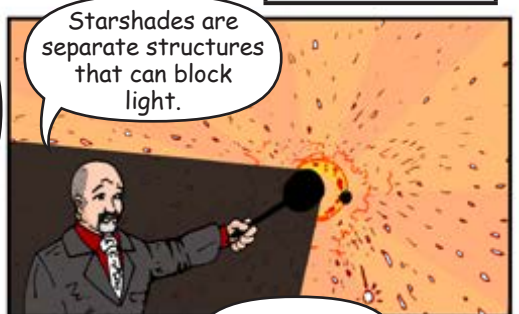
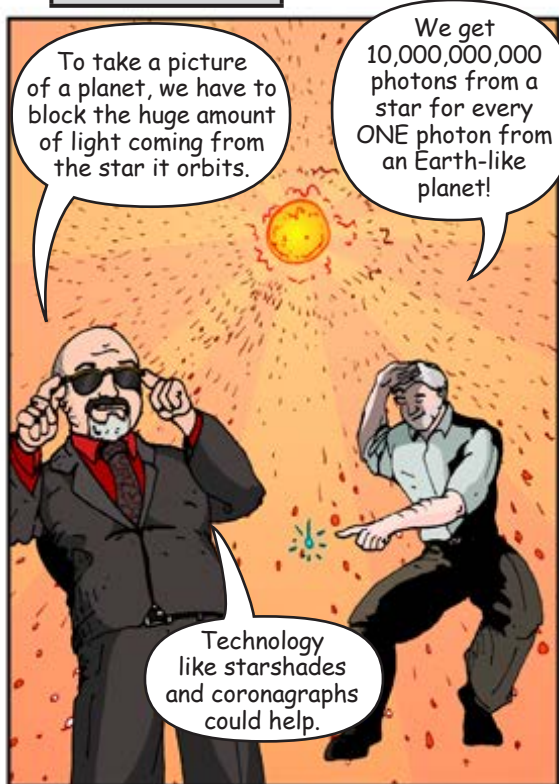
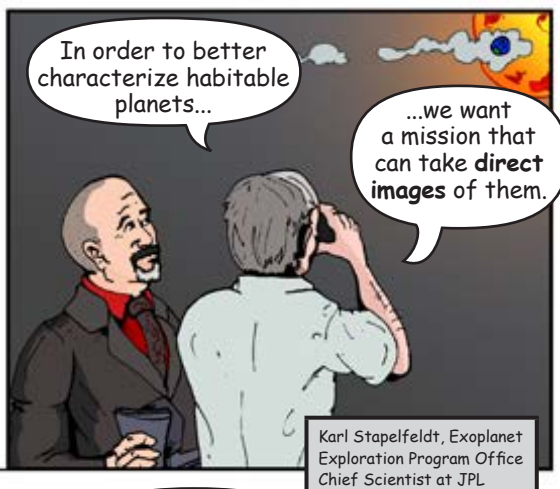
We'll only be able to study a handful of planets out of thousands of possibilities.

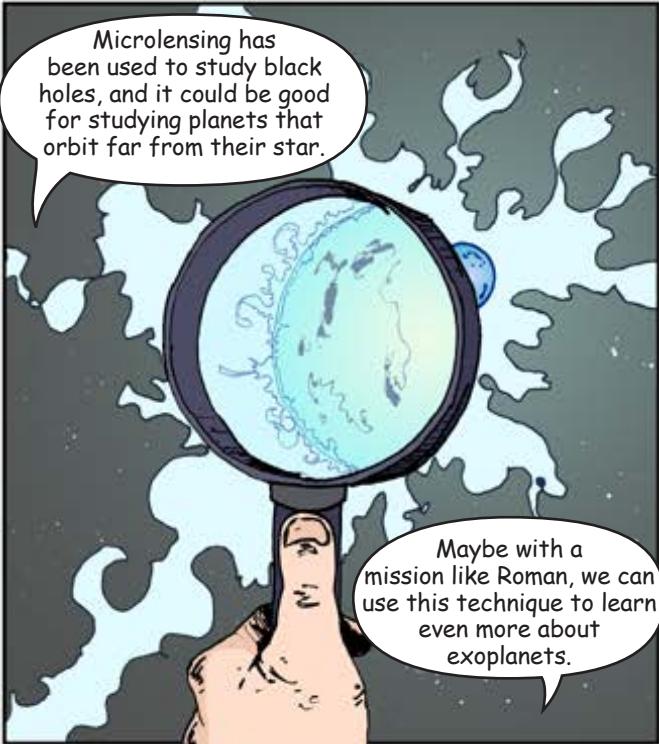
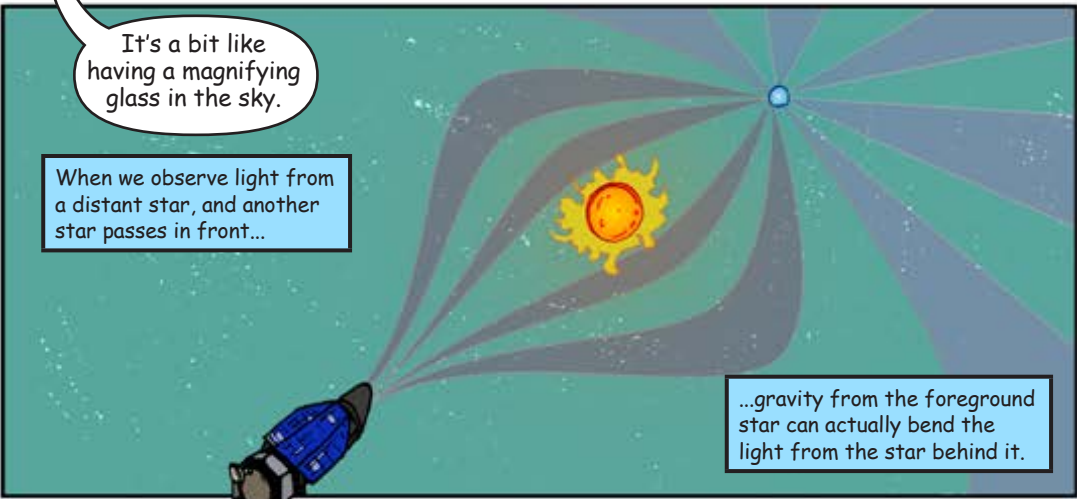
Still, JWST is a huge advancement in our capabilities. After all, it's **100 times** more powerful than Hubble!

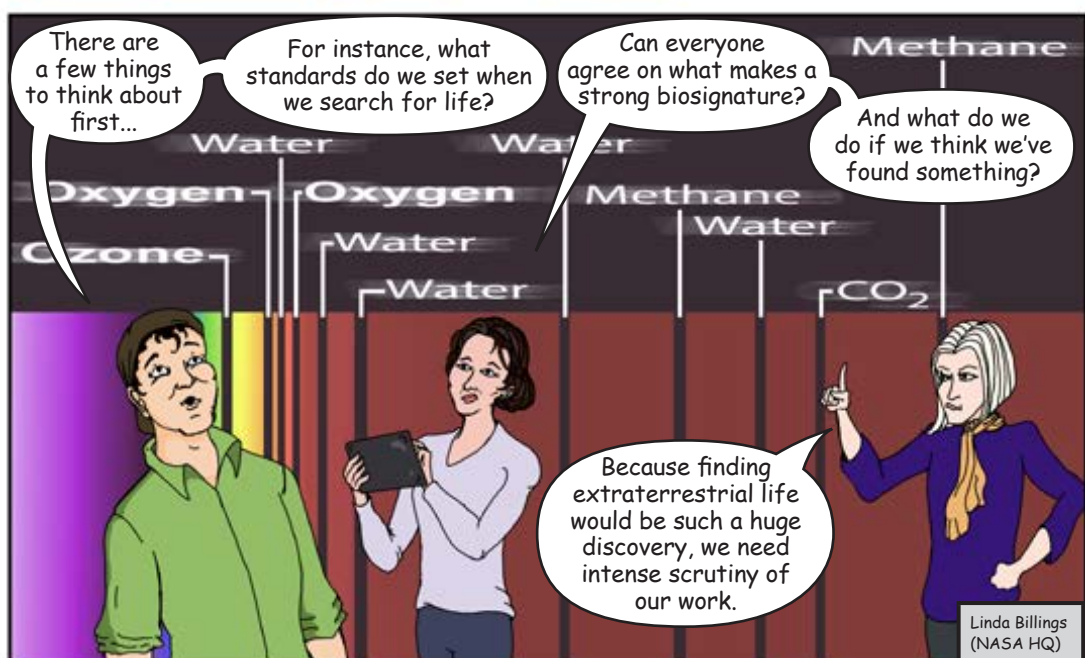
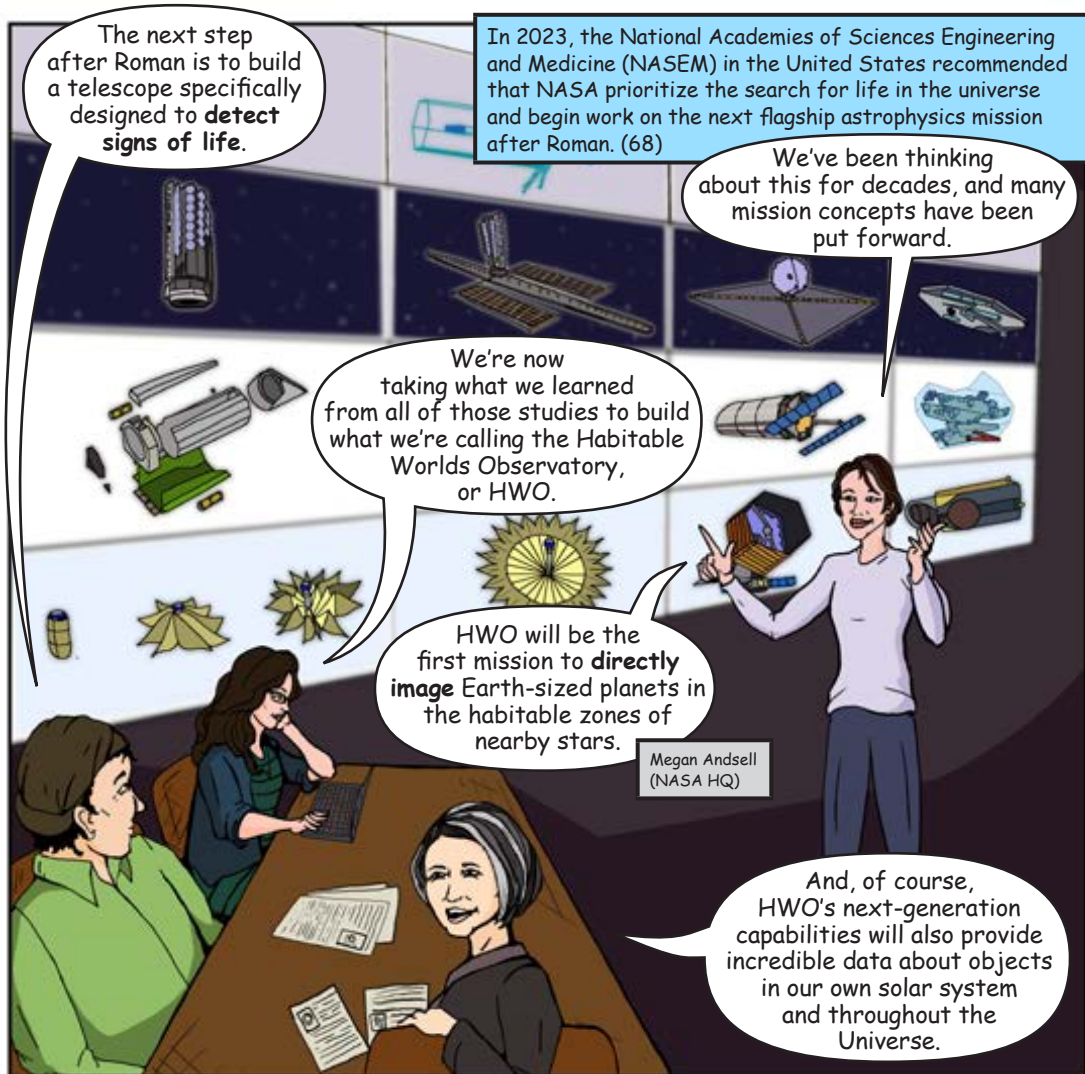
Mary Voytek, NASA
Astrobiology Program

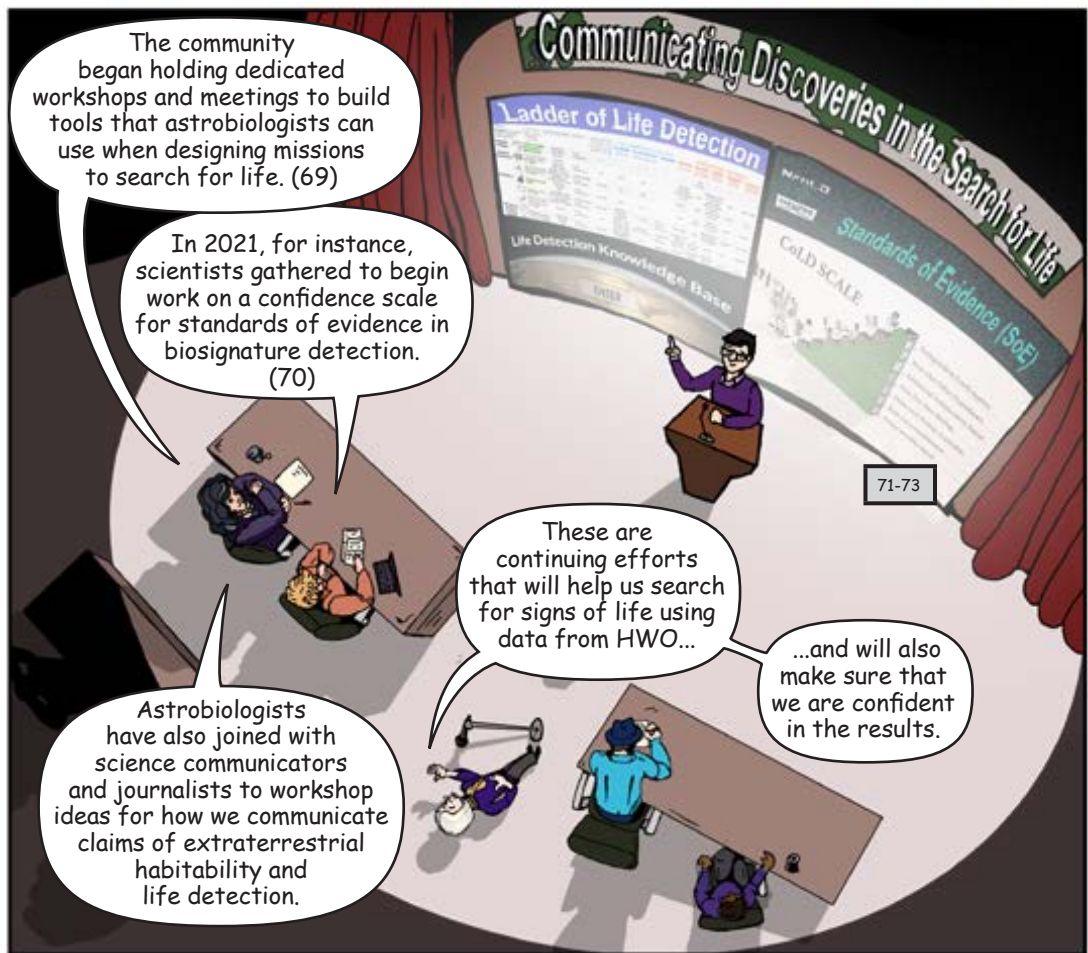
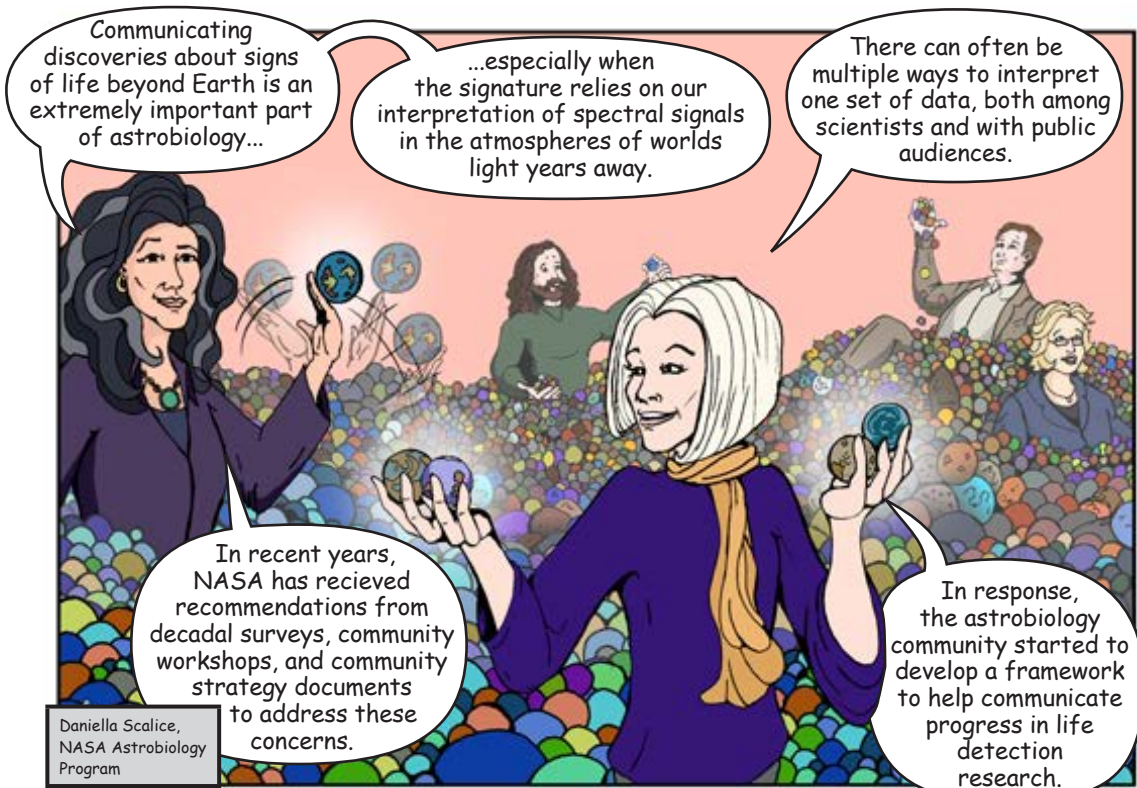
Even some of the first images that the mission sent back highlighted this advance...

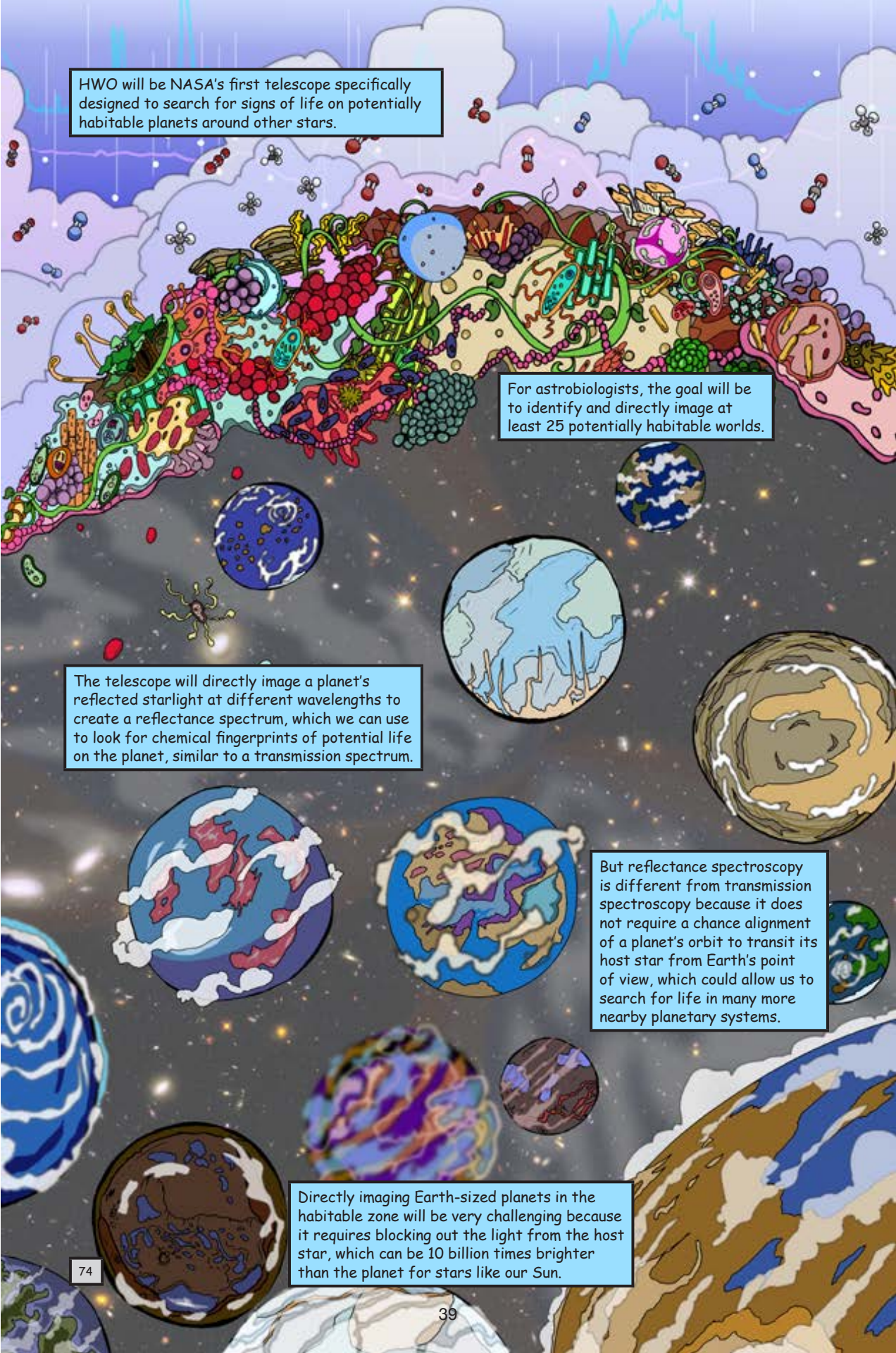












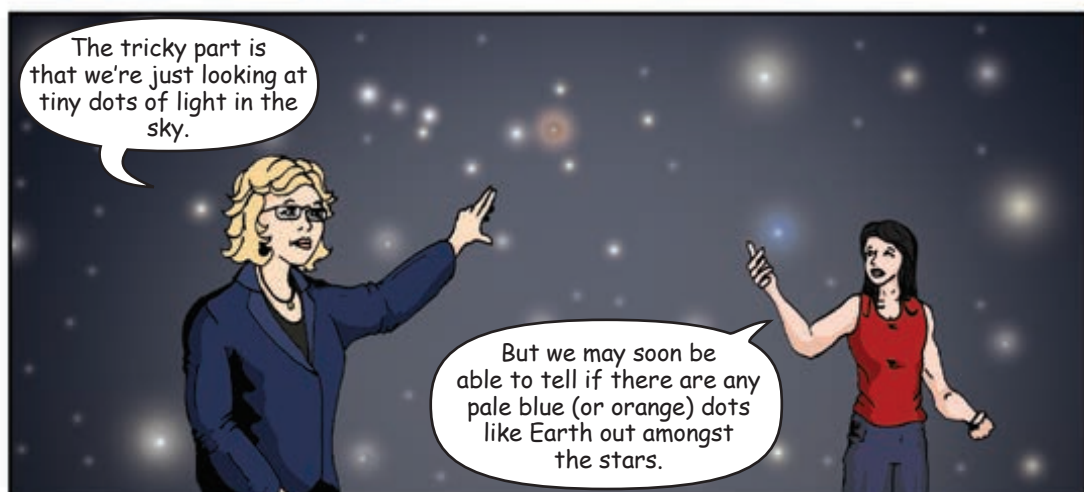
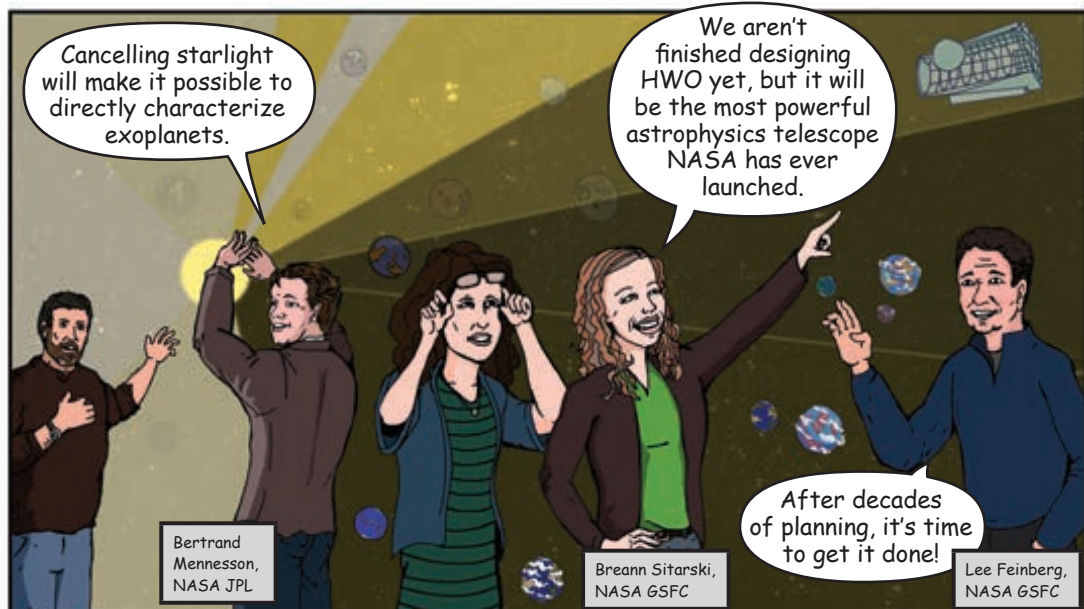
HWO will be NASA's first telescope specifically designed to search for signs of life on potentially habitable planets around other stars.

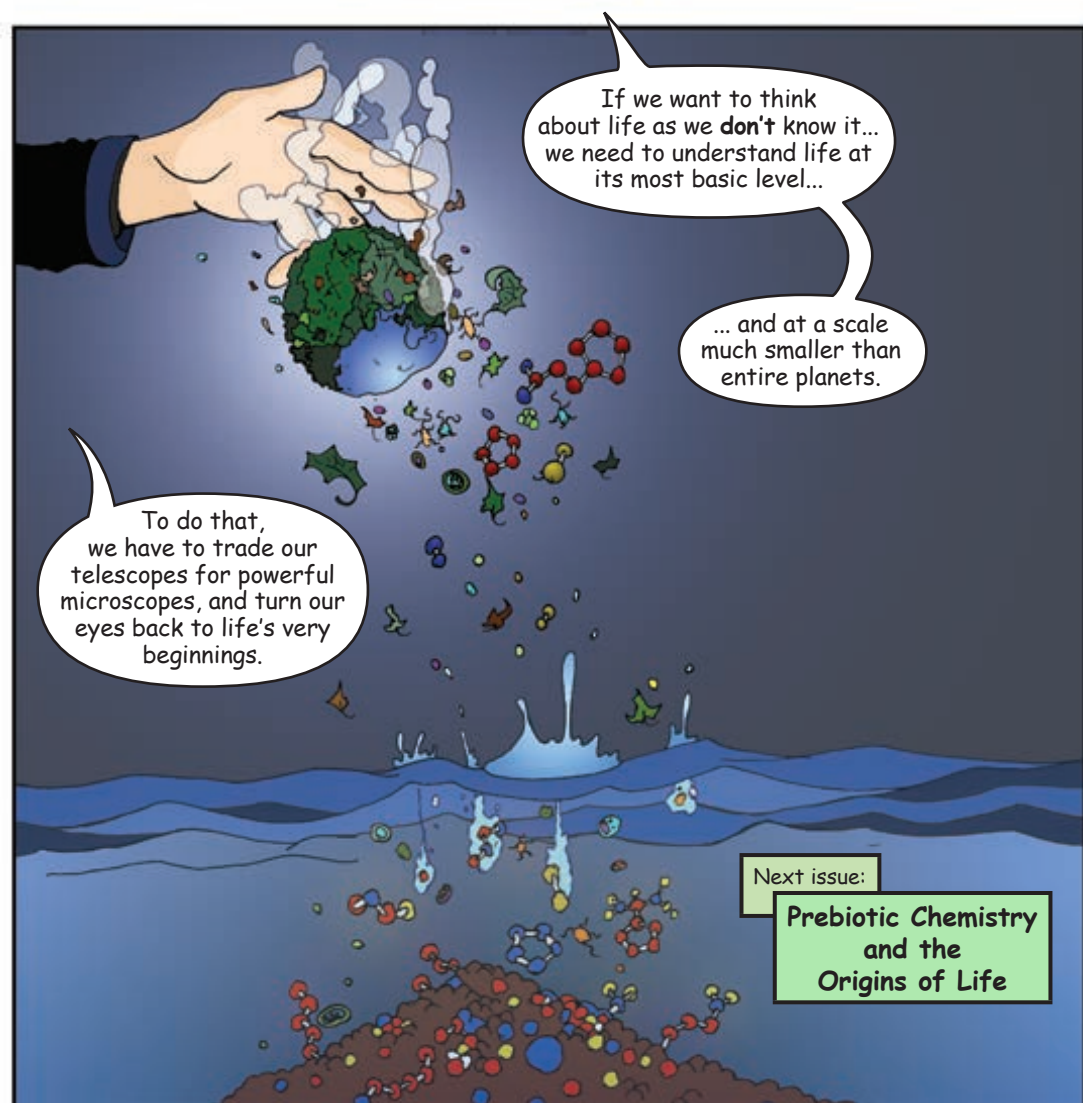
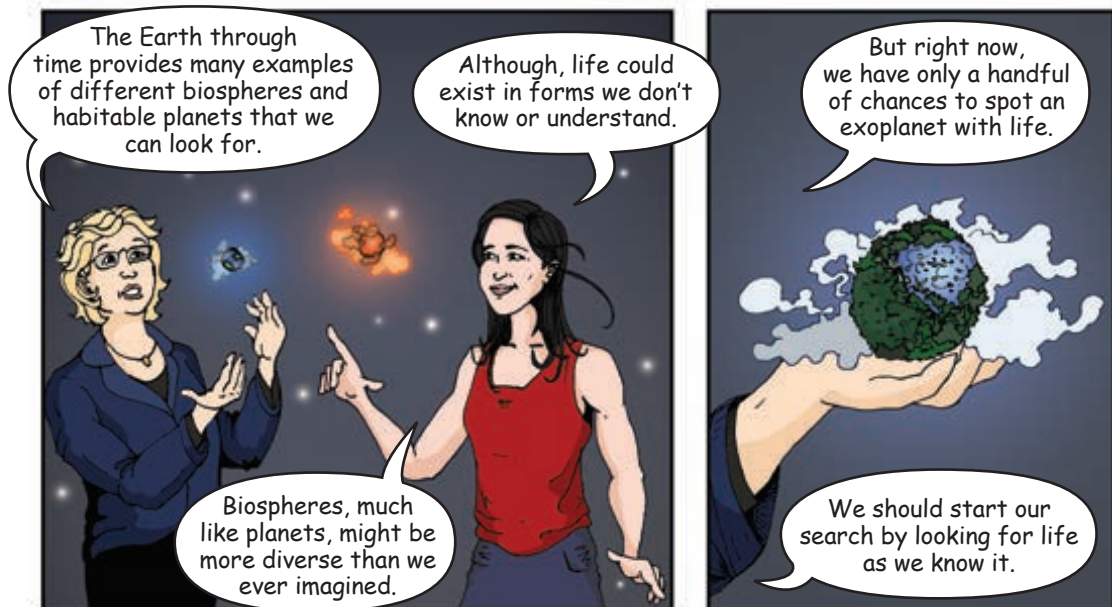
For astrobiologists, the goal will be to identify and directly image at least 25 potentially habitable worlds.

The telescope will directly image a planet's reflected starlight at different wavelengths to create a reflectance spectrum, which we can use to look for chemical fingerprints of potential life on the planet, similar to a transmission spectrum.

But reflectance spectroscopy is different from transmission spectroscopy because it does not require a chance alignment of a planet's orbit to transit its host star from Earth's point of view, which could allow us to search for life in many more nearby planetary systems.

Directly imaging Earth-sized planets in the habitable zone will be very challenging because it requires blocking out the light from the host star, which can be 10 billion times brighter than the planet for stars like our Sun.





Astrobiology

A History of Exobiology and Astrobiology at NASA

Further Resources and References cited in this issue:

1. Bruno, G. (1584) *De l'infinito universon e mondi* (On the Infinite Universe and Worlds)
2. Mayor, M., and Queloz, D. (1995) A Jupiter-mass companion to a solar-type star. *Nature*, Vol. 378, 355 - 359.
3. NASA. (2017) Five ways to find a planet. NASA Exoplanet Exploration Program. Available at: <https://exoplanets.nasa.gov/interactable/11/>
4. NASA Hubble Mission Team (2001) Hubble Makes First Direct Measurements of Atmosphere on World Around another Star. <https://science.nasa.gov>
5. Charbonneau, D., et. al. (2002) Detection of an Extrasolar Planet Atmosphere. *The Astrophysical Journal*, Vol. 568, No. 1.
6. Tsiaras, A. et. al. (2016) Detection of an atmosphere around the super-Earth 55 Cancri E. *The Astrophysical Journal*, Vol. 820, No. 99. DOI: 10.3847/0004-637X/820/2/99
7. NASA. Kepler / K2. Available at: <https://science.nasa.gov/mission/kepler/>
8. The Virtual Planetary Laboratory. University of Washington, Seattle. <https://depts.washington.edu/naivpl/content/welcome-virtual-planetary-laboratory>
9. Seager, S. et. al. (2007) Mass-radius relationships for solid exoplanets. *The Astrophysical Journal*, Volume 669, Number 2
10. Adams, E.R., et. al. (2008) Ocean planet or thick atmosphere: On the mass-radius relationship for solid exoplanets with massive atmospheres. *The Astrophysical Journal*, Volume 673, Number 2
11. Rogers, L. (2014) Most 1.6 Earth-radius planets are not rocky. *The Astrophysical Journal*, Vol. 801, No. 41.
12. Kopparapu, R.K., et. al. (2013) Habitable zones around main-sequence stars: New estimates. *The Astrophysical Journal*, Vol. 765, No. 2
13. Haqq-Misra, J. D., et. al. (2009) A revised, hazy methane greenhouse for the Archean Earth. *Astrobiology*, Vol. 8, No. 6, 1127-1137.
14. Snellen, I. (2010) Exoplanet atmospheres at high spectral resolution: A CRIRES survey of hot-Jupiters. *Proceedings of the International Astronomical Union*, Vol. 6, 208-211
15. Seager, S. and Sasselov, D. D. (2000) Theoretical transmission spectra during extrasolar giant planet transits. *The Astrophysical Journal* Vol. 537, 916-921.
16. Seager, S., & Sasselov, D. D. (1998) Extrasolar giant planets under strong stellar irradiation. *The Astrophysical Journal*, Vol. 502, No. 1, 157-161.
17. Seager, S., & Sasselov, D. D. (2000) Theoretical Transmission Spectra During an Extrasolar Giant Planet Transit. *The Astrophysical Journal*, Vol. 537, 916-921.
18. Deming, D., et. al. (2005) Detection of infrared radiation from an extrasolar planet. *Nature*, Vol. 434, 740-743.
19. Seager, S., et. al. (2005) Exoplanet atmospheres and photochemistry. *Proceedings of the International Astronomical Union*, Vol. 1, No. S231, 491-198.
20. Seager, S. and Deming, D. (2010) Exoplanet Atmospheres. *Annual Review of Astronomy and Astrophysics*, Vol. 48, 631-672.
21. Cockell (2015) *Astrobiology: Understanding Life in the Universe*. John Wiley & Sons. 472 pages.
22. Arney, G., et. al. (2016) The pale orange dot: The spectrum and habitability of hazy archean Earth. *Astrobiology*, Vol. 16, No. 11, 873-899.

23. Luger, R., et. al. (2016) The pale green dot: A method to characterize Proxima Centauri b using exo-auroae. *The Astrophysical Journal*, Vol. 837, No. 1
24. Kiang, N., et. al. (2007) Spectral signatures of photosynthesis. I. Review of Earth organisms. *Astrobiology*, Vol. 7, No. 1, 222-251.
25. Kiang, N., et. al. (2007) Spectral signatures of photosynthesis. II. Coevolution with other stars and the atmosphere of extrasolar worlds. *Astrobiology*, Vol. 7, No. 1, 252-274.
26. Schopf, J.W. (Ed.) (1983) *Earth's earliest biosphere: Its origin and evolution*. Princeton University Press, Princeton, NJ. 543 pgs.
27. Schwieterman, E., et. al. (2015) Nonphotosynthetic pigments as potential biosignatures. *Astrobiology*, Vol. 15, No. 5, 341-361.
28. Domagal-Goldman, S.D., et. al. (2014) Abiotic ozone and oxygen in atmospheres similar to prebiotic Earth. *The Astrophysical Journal*, Vol. 792, No. 2
29. Domagal-Goldman, S.D. and Meadows, V. (2010) Abiotic buildup of ozone. *Pathways towards habitable planets*, ASP Conference Series, Vol. 430.
30. Luger, R. and Barnes, R. (2015) Extreme water loss and abiotic O₂ buildup on planets throughout the habitable zones of M dwarfs. *Astrobiology*, Vol. 15, No. 2, 119-143.
31. Cockell, C.S., et. al. (2009) Cryptic photosynthesis—Extrasolar planetary oxygen without a surface biological signature. *Astrobiology*, Vol. 9, No. 7, 623-636.
32. Harman, C.E., et. al. (2015) Abiotic O₂ levels on planets around F, G, K and M stars: Possible false-positives for life? *The Astrophysical Journal*, Volume 812, Number 2.
33. Nexus for Exoplanet System Science (NExSS). <https://nexss.info>
34. NASA. The Habitable Zone. <https://science.nasa.gov>
35. Guo, J., et al. (2009) Probability distribution of terrestrial planets in habitable zones around host stars. *Astrophysics and Space Science*, Vol. 323, p. 367-373.
36. Berta, Z.K., et. al. (2011) The flat transmission spectrum of the super-Earth GJ1214b from wide field camera 3 on the Hubble Space Telescope. *The Astrophysical Journal*, Vol. 747, No. 1.
37. Barnes, R., et. al. (2009) Tidal limits to planetary habitability. *The Astrophysical Journal Letters*, Vol. 700, No. 1.
38. Barnes, R., et. al. (2013). Tidal Venuses: Triggering a climate catastrophe via tidal heating. *Astrobiology*, Vol. 13, No. 3, 225-250.
39. Jackson, B., et. al. (2008) Tidal heating of extrasolar planets. *The Astrophysical Journal*, Vol. 681, No. 2.
40. Barnes, R., et. al. (2010) CoRoT-7b: Super-Earth or super-Io? *The Astrophysical Journal Letters*, Vol. 709, No. 2.
41. Barnes, J.W. (2007) Effects of orbital eccentricity on extrasolar planet transit detectability and light curves. *Publications of the Astronomical Society of the Pacific*, Vol. 119, No. 859.
42. Shields, A., et. al. (2016) The Effect of orbital configuration on the possible climates and habitability of Kepler-62 f. *Astrobiology*. Vol. 16, No. 6, 443-464.
43. Heller, R., et. al. (2011) Tidal obliquity evolution of potentially habitable planets. *Astronomy and Astrophysics*, Vol. 528.
44. Shields, A., et. al. (2013) The effect of host star spectral energy distribution and ice- albedo feedback on the climate of extrasolar planets. *Astrobiology*, Vol. 13, No. 8.
45. Shields, A., et. al. (2016) The habitability of planets orbiting M-dwarf stars, *Physics Reports*, Vol. 663, 1-38.
46. Yang, J., et. al. (2014) Strong dependence of the inner edge of the habitable zone on planetary rotation rate. *The Astrophysical Journal Letters*, Volume 787, Number 1.

47. Shields, A., et. al. (2011) Earth as an extrasolar planet: Comparing polar and equatorial views of modern day and snowball Earth. *Bulletin of the American Astronomical Society*, Vol. 43.
48. This image shows the first flash of sunlight reflected off a lake on Saturn's moon Titan. The glint off a mirror-like surface is known as a specular reflection, and was detected by the visual and infrared mapping spectrometer (VIMS) on NASA's Cassini spacecraft on July 8, 2009. Image credit: NASA/JPL/University of Arizona/DLR
49. A sun glint on Earth is captured (center of the black circle) in the middle frame of this series of images taken by NASA's Deep Impact spacecraft as it looked at the north pole. Image Credit: Don Lindler, Sigma Space Corporation/GSFC
50. Robinson, T., et. al. (2010) Detecting oceans on extrasolar planets using the glint effect. *Astrophysical Journal Letters*, Vol. 721, No. 1.
51. NASA GSFC. (2016) Transiting Exoplanet Survey Satellite. Available at: <https://exoplanets.nasa.gov/tess/>
52. Dragomir, D., et. a. (2019) TESS Delivers Its First Earth-sized Planet and a Warm Sub-Neptune. *The Astrophysical Journal Letters*, Vol. 875, No. 2. DOI: 10.3847/2041-8213/ab12ed
53. Kruse, E. (2023) TESS 2022 Sky Views. NASA Scientific Visualization Studio, NASA's Goddard Space Flight Center.
54. NASA. James Webb Space Telescope. <https://science.nasa.gov/mission/webb/>
55. Gillon, M., et. al. (2017) Seven temperate terrestrial planets around the nearby ultracool dwarf star TRAPPIST-1. *Nature* 542, 456-460.
56. Lim, O., et. al. (2023) Atmospheric Reconnaissance of TRAPPIST-1 b with JWST/NIRISS: Evidence for Strong Stellar Contamination in the Transmission Spectra. *The Astrophysical Journal*, Vol. 955, No. 1. DOI: 10.3847/2041-8213/acf7c4
57. Greene, T.P., et. al. (2023) Thermal emission from the Earth-sized exoplanet TRAPPIST-1 b using JWST. *Nature*, Vol. 618, 39-42. DOI: 10.1038/s41586-023-05951-7
58. Zieba, S., et. al. (2023) No thick carbon dioxide atmosphere on the rocky exoplanet TRAPPIST-1 c. *Nature*, Vol. 620, 746-749. DOI: 10.1038/s41586-023-06232-z
59. Thomas, T.B., et. al. (2025) Statistical Geochemical Constraints on Present-day Water Outgassing as a Source of Secondary Atmospheres on the TRAPPIST-1 Exoplanets. *The Planetary Science Journal*, Vol. 6, No. 126. DOI: 10.3847/PSJ/add261
60. Nixon, C.N., et. al. (2025) The atmosphere of Titan in late northern summer from JWST and Keck observations. *Nature Astronomy*. DOI: 10.1038/s41550-025-02537-3
61. Cartwright, R.J., et. al. (2025) JWST Reveals Spectral Tracers of Recent Surface Modification on Europa. *The Planetary Science Journal*, Vol. 6, No. 125. DOI: 10.3847/PSJ/adcab9
62. NASA's Goddard Space Flight Center (2022) Carina Nebula Webb vs Hubble. <https://svs.gsfc.nasa.gov/>. Hubble image of NGC 3324 (2008) Credit: NASA, ESA, and The Hubble Heritage Team (STScI/AURA)
63. NASA Webb Telescope Team (2022) NASA's Webb Reveals Cosmic Cliffs, Glittering Landscape of Star Birth. <https://www.nasa.gov/>
64. NASA (2018) The Pillars of Creation. Credits: NASA, ESA and the Hubble Heritage Team (STScI/AURA). <https://www.nasa.gov>
65. NASA Webb Telescope Team. (2022) NASA's Webb Takes Star-Filled Portrait of Pillars of Creation. Science Credit: NASA, ESA, CSA, STScI. Image Processing: DePasquale, J., Koekemoer, A.M., Pagan, A. <https://www.nasa.gov>

66. NASA. Nancy Grace Roman Space Telescope. <https://science.nasa.gov/mision/roman-space-telescope/>
67. Batista, V., et. al. (2013) MOA-2011-BLG-293LB: First microlensing planet possibly in the habitable zone. *The Astrophysical Journal*, Vol. 780, No. 1.
68. National Academies of Sciences, Engineering, and Medicine. (2023) *Pathways to Discovery in Astronomy and Astrophysics for the 2020s*. Washington, DC: The National Academies Press. DOI: 10.17226/26141
69. NASA Astrobiology. Resources for Life Detection Missions. Available at: <https://science.nasa.gov/astrobiology/researchers/life-detection-resources/>
70. NfOLD/NExSS. (2021) Standards of Evidence for Life Detection Community Workshop. Available at: <https://www.nfold.org/standards-of-evidence-workshop>
71. Neveu, M. et. al. (2018) The Ladder of Life Detection. *Astrobiology*, Vol. 18, No. 11. DOI: 10.1089/ast.2017.177
72. NASA Astrobiology (2025) NASA's Life Detection Knowledge Base: A Tool for Life Detection Mission Planning. Available at: <https://science.nasa.gov/astrobiology/researchers/life-detection-resources/ldkb/>
73. Green, J., et. al. (2021) Call for a framework for reporting evidence for life beyond Earth. *Nature*, Vol. 598, p. 575–579. DOI: 10.1038/s41586-021-03804-9
74. Image adapted from: NASA, ESA, Windhorst, R., McCarthy, P., O'Connell, R., WFC3 Science Oversight Committee (2011) GOODS South Deep (GSD). Hubble Uncovers Tiny Galaxies Bursting with Star Birth in Early Universe.

