



# **NASA Strategic Astrophysics Technology Investments: A Decade of Benefits, Outlook Informed by the 2020 Decadal Survey**

**Opher Ganel\*, Rachel Rivera, and Jay Falker (NASA/GSFC)  
Nicholas Siegler and Brendan Crill (NASA/JPL)  
Mario R. Perez (NASA/HQ)**

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# Strategic Astrophysics Missions



- **Strategic Astrophysics missions as recommended by the 2020 Decadal Survey on Astronomy and Astrophysics**
  - IR/O/UV Flagship to be launched in the first half of the 2040s
  - X-ray Flagship to be launched in a later decade than IR/O/UV Flagship
  - Far-IR Flagship to be launched in a later decade than IR/O/UV Flagship
  - X-ray Probe to compete for 2030s launch
  - Far-IR Probe to compete for 2030s launch
  - CMB Probe to compete for 2040s launch vs. above Probe that hasn't launched
  - TDAMM (program recommendation at this point, but missions may be added)





# Strategic Astrophysics Technology Development



- **The three Program Offices serve the critical function of developing concepts and technologies for strategic missions and facilitating science investigations derived from them, specifically:**
  - Assess and prioritize technology gaps, collecting inputs from the community and technology activities.
  - Manage projects maturing technologies for strategic missions from TRLs of 3, 4, or 5.
  - Promote infusion of technologies into missions and projects.
  - Conduct mission studies and develop mission concepts enabling future scientific discoveries.
  - Communicate progress to and coordinate with the scientific community.
  - Inform the general public about progress achieved by the Programs.
- **Astrophysics Program Office technology maturation process comprises three cycles**
  - A 10-year Decadal Survey cycle to recommend the highest-priority missions and activities for the coming decade.
  - A biennial technology gap prioritization and Program Office reporting cycle.
  - An annual process of soliciting, reviewing, and funding SAT proposals.





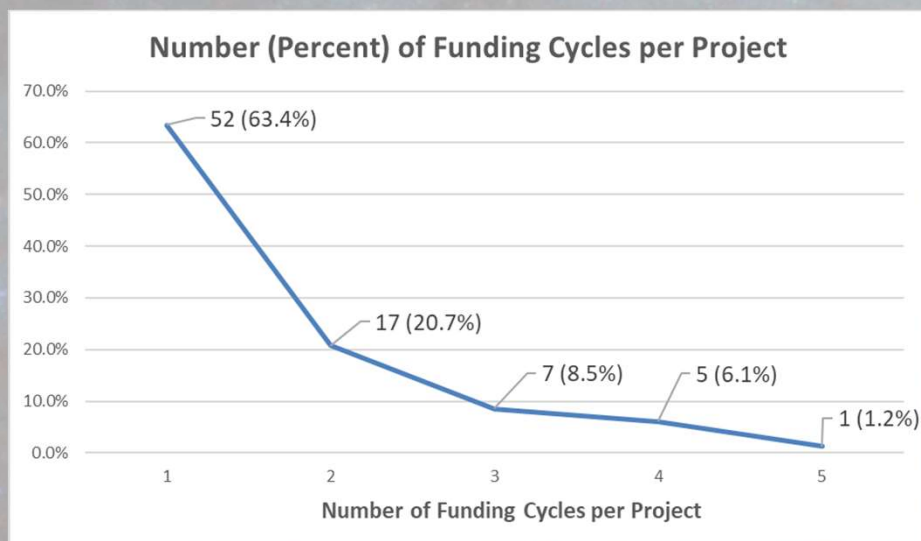
# Astrophysics Strategic Technology Portfolio



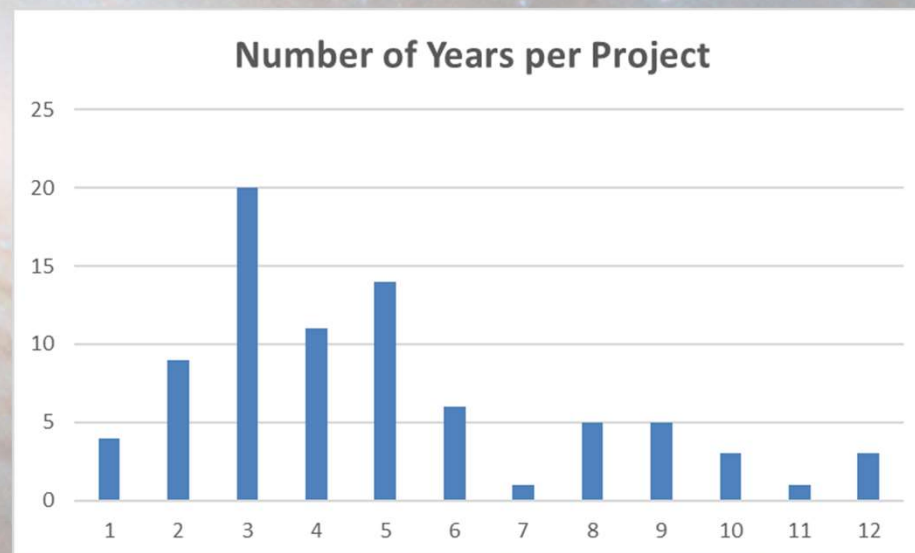
Current COR	Technology Development Projects	PI Name	Org.	Tech Area		
Current ExEP	Technology Development Projects	PI Name	Org.	Tech Area		
Photon-Counting NIR LmAPD Arr	Development of a Method for Exop	Current PCOS Technology Development Projects		PI Name	Org.	Tech Area
Ultrasensitive Bolometers for Far	Laboratory Demonstration of High	Advanced X-ray Microcalorimeters: Magnetically Coupled Calorimeters		Bandler, Simon	GSFC	Detector
Ultrastable Large Telescope Rese	Laboratory Demonstration of High	High-Speed, Low-Noise, Radiation-Tolerant CCD Image Sensors		Bautz, Mark	MIT	Detector
A Single-Photon-Sensing and Pho	Laboratory Demonstration of Multi	Microwave SQUID Readout to Enable Lynx and Other Future Observatories		Bennett, Douglas	NIST	Electronics
Electron Beam Lithography Rule	MEMS Deformable Mirror Technolo	Direct Fabrication of Full-Shell X-ray Optics		Bongiorno, Stephen	MSFC	Optics
Scalable Micro-Shutter Systems	MEMS Deformable Mirror Technolo	Hybrid X-ray Optics by Additive Manufacturing		Broadway, David	MSFC	Optics
High-Performance, Stable, and S	Segmented Coronagraph Design an	Low-Stress Mirror Coatings for X-ray Optics		Broadway, David	MSFC	Opt. Coatings
Development of High-Resolution	Segmented Coronagraph Design an	UV-LED-Based Charge Management System		Conklin, John	UF	Electronics
Development of Digital Micromir	Linear Wavefront Control for High-C	Computer-Controlled Polishing of High-Quality X-ray Optics Mandrels		Davis, Jacqueline	MSFC	Optics
Technology Maturation for Astro	Optimal Spectrograph and Wavefro	High-Density Readout Technology for Superconducting Sensor Arrays for Spaceflight		Frisch, Josef	SLAC	Electronics
Electron-Beam Generated Plasm	Radiation-Tolerant, Photon-Countin	Differential Deposition for Figure Correction in X-ray Optics		Kilaru, Kiran	MSFC	Optics
Ultra-Stable Structures Developm	Broadband Light Rejection with the	Advanced TES Microcalorimeters		Kilbourne, Caroline	GSFC	Detector
High Performance Sealed Tube C	Broadband Light Rejection with the	Enabling and Enhancing Technologies for ATHENA X-IFU Demonstration Model		Kilbourne, Caroline	GSFC	Detector
Process to Produce Scalable, Sup	Vortex Coronagraph High-Contrast	Advancing the Focal Plane TRL for LiteBIRD and Other Next-Generation CMB Missions		Lee, Adrian	UCB	Detector
Precision Thermal Control (PTC)	Vortex Coronagraph High-Contrast	Telescopes for Space-Based Gravitational-Wave Observatories		Livas, Jeffrey	GSFC	Telescope
High-Efficiency Continuous Cooli	First System-level Demonstration o	Development of RFSoc-Based Readout Electronics		Mauskopf, Philip	ASU	Electronics
Large-Format, High-Dynamic-Ran	First System-level Demonstration o	Superconducting Antenna-Coupled Detectors for CMB Polarimetry		O'Brient, Roger	JPL	Detector
	Ultra-Stable Mid-Infrared Detector	Advanced X-ray Microcalorimeters: Lab Spectroscopy for Space Atomic Physics		Porter, Scott	GSFC	Detector
	Super Lyot ExoEarth Coronagraph	X-ray Testing and Calibration		Ramsey, Brian	MSFC	Optics
	A Novel Optical Etalon for Precision	Development of Adjustable X-ray Optics with 0.5 Arcsecond Resolution		Reid, Paul	SAO	Optics
	Starshade Large-Structure Precision	High Resolution High Efficiency X-ray Transmission Grating Spectrometer		Schattenburg, Mark	MIT	Optics
	Starshade Starlight Suppression	Manufacturability & Alignment of X-ray Gratings and Optics for Space Applications		Smith, Randall	SAO	Optics
		Phase Measurement System for Interferometric Gravitational Wave Detectors		Ware, Brent	JPL	Electronics
		Space-based Gravitational Wave Laser Technology Development for the LISA Mission		Yu, Anthony	GSFC	Laser
		Next Generation X-ray Optics		Zhang, William	GSFC	Optics
		LISA Colloid Microthruster Technology		Ziemer, John	JPL	Micropropulsion

Note that this does not include APRA and RTF projects, which are not strategic.

# Technology Development Projects' Length



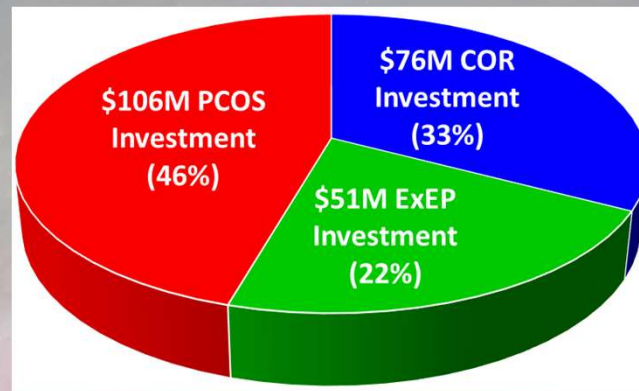
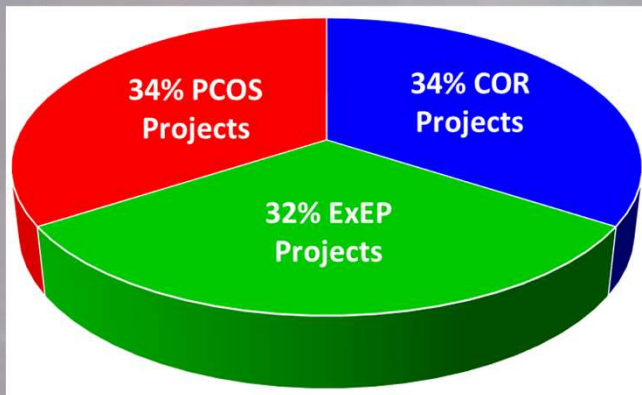
Number of funding cycles per project, with 30 projects extending 1–4 additional cycles.



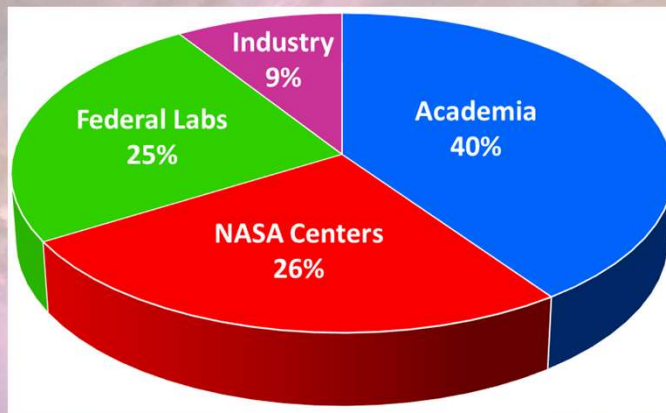
Distribution of projects by total length in years, including any follow-on cycles.



# Technology Development Projects by Program and Org Type

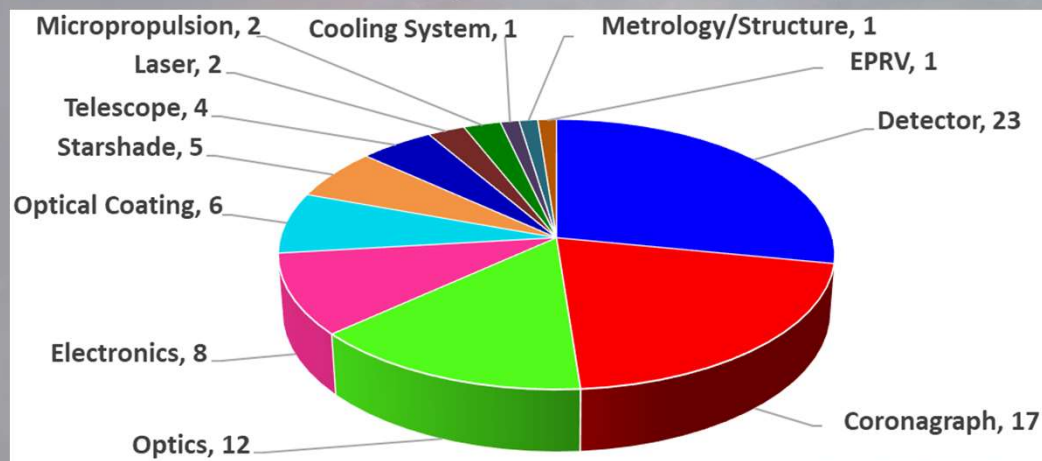


Technology development by Program since 2009: number of projects (left), and total investment (right).



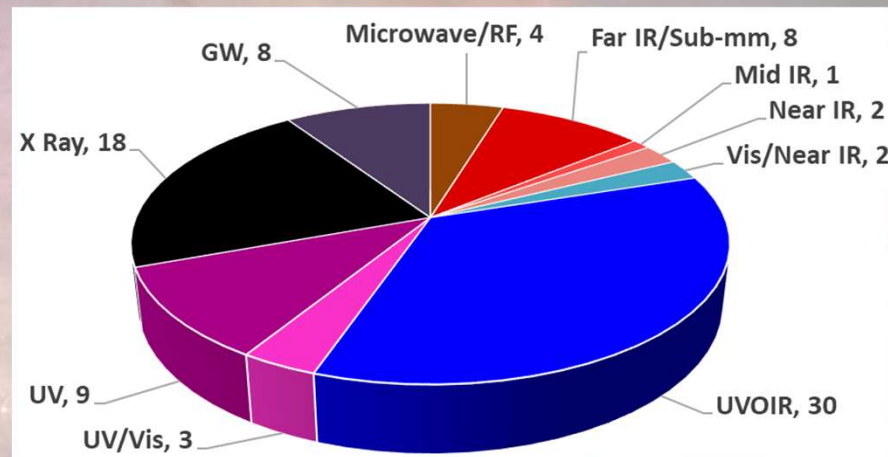
Distribution institutional types receiving awards for strategic technology developments.

# Technology Development Projects by Technology Area and Signal Type

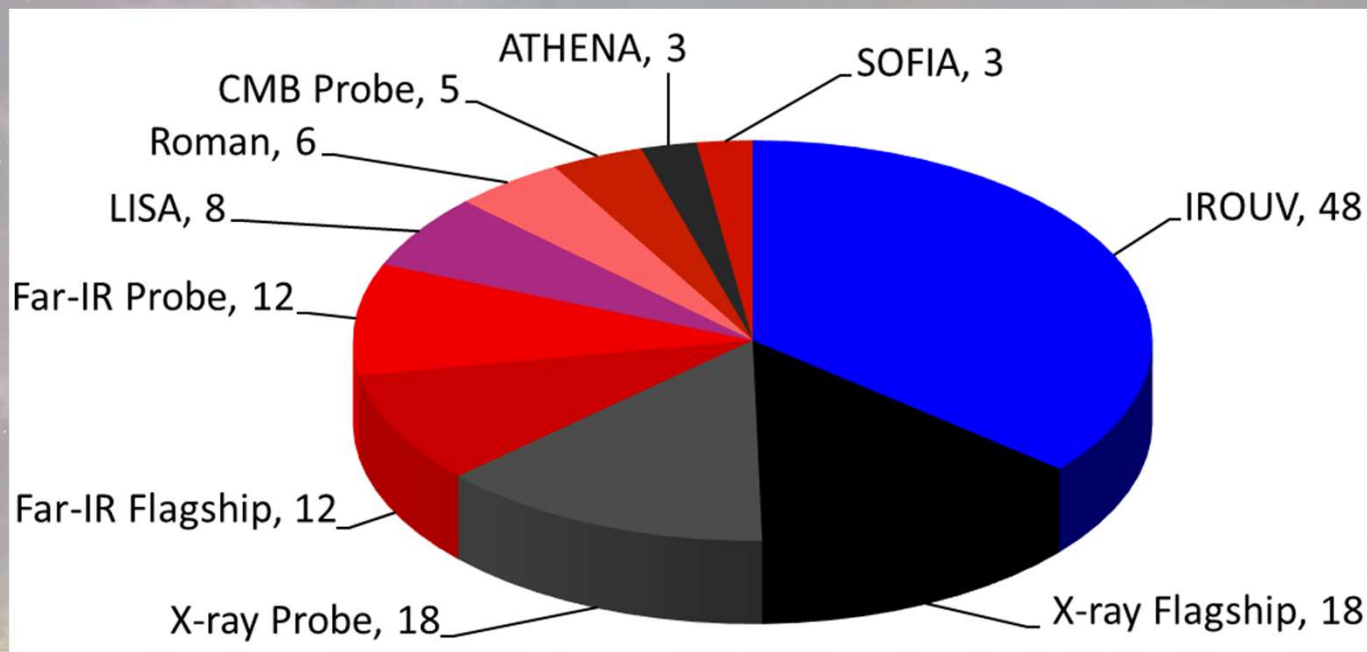


Distribution of the 82 projects in the strategic technology portfolio by topic.

Distribution of the 82 projects in the strategic technology portfolio by signal type



# Technology Development Projects by Strategic Mission Supported

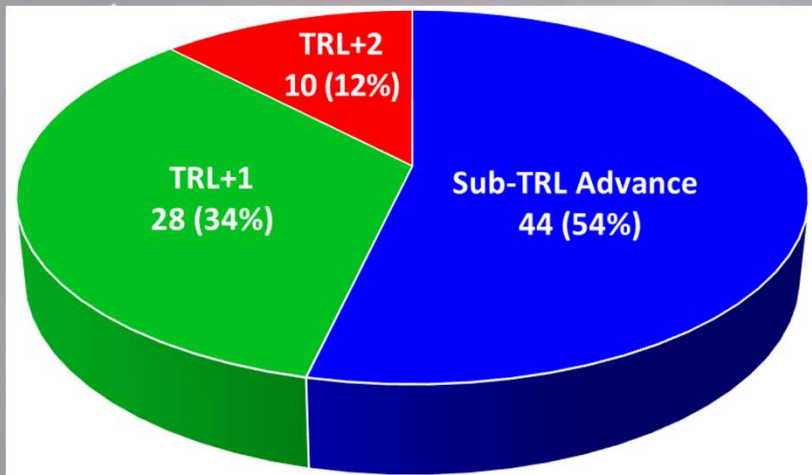


**Distribution of the 82 projects in the strategic technology portfolio by strategic mission(s) supported.**  
**Note that many projects support multiple missions, so the total is greater than 82.**

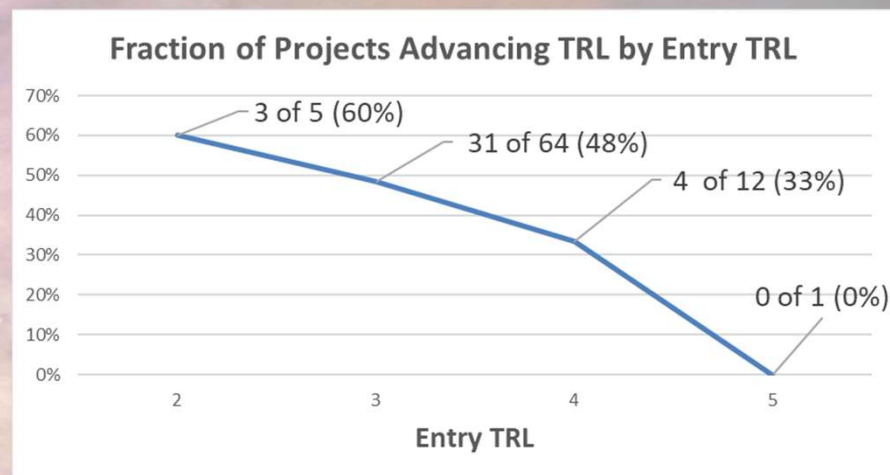


# Technology Readiness Level (TRL) Advancements

Distribution of TRL advancement (while in the program) for strategic technology development projects. Some continue advancing in TRL after project completion, especially those infused into spaceflight projects.



Percentage of technology development projects achieving TRL advancement vs.  $TRL_{in}$ . Not shown in the graph, nine of 31 projects advancing from  $TRL_{in}$  of 3 advanced by two levels, to 5.



# Technology Infusions

	Space	Rocket	Balloon	Airborne	Ground	Total
<b>Implemented</b>	8	14	8	2	25	<b>57</b>
<b>Upcoming</b>	21	14	5	1	15	<b>56</b>
<b>Concepts</b>	61	-	-	-	-	<b>61</b>
<b>Total</b>	<b>90</b>	<b>28</b>	<b>13</b>	<b>3</b>	<b>40</b>	<b>174</b>

**Summary of infusions, already implemented/upcoming/conceptual by mission/project type.**  
 (Implemented = in existing/past mission/project; Upcoming = selected by mission/project in development; Concepts = baselined by reference design of strategic mission)

Discipline	Space	Rocket	Balloon	Airborne	Ground	Total
<b>Astrophysics</b>	22	15	10	1	23	<b>71</b>
<b>Planetary</b>	4	1	-	-	-	<b>5</b>
<b>Heliophysics</b>	4	3	-	-	-	<b>7</b>
<b>Earth Science</b>	3	-	-	-	-	<b>3</b>
<b>Other</b>	-	-	-	1	1	<b>2</b>
<b>Total</b>	<b>33</b>	<b>19</b>	<b>10</b>	<b>2</b>	<b>24</b>	<b>88</b>

**Summary of missions/projects infused into, by discipline and type.**





# Astrophysics Technology Gaps



- **Technology gaps are what separates the current state of the art (SOTA) from what's needed to enable and/or enhance future strategic missions**
- **Technology gap entries need to:**
  - Focus on strategic Astrophysics missions
  - Be directly applicable to Program objectives
  - Require technology development (i.e., TRL < 6 for stated requirements)
  - Describe gap between SOTA and what's required for targeted science objective
- **Technology gaps entries should not:**
  - Be outside our purview (e.g., for launch vehicles, spacecraft systems, etc.)
  - Be a specific solution, or advocate for one





# Strategic Technology Gap Priorities – Tier 1



1. Advanced Cryocoolers
2. Coronagraph Contrast and Efficiency
3. Coronagraph Stability
4. Cryogenic Readouts for Large-Format Far-IR Detectors
5. Heterodyne Far-IR Detector Systems
6. High-Performance, Sub-Kelvin Coolers
7. High-Reflectivity Broadband Far-UV-to-Near-IR Mirror Coatings
8. High-Resolution, Large-Area, Lightweight X-ray Optics
9. High-Throughput Bandpass Selection for UV/VIS
10. High-Throughput, Large-Format Object Selection Technologies for Multi-Object and Integral Field Spectroscopy
11. Large Cryogenic Optics for the Mid IR to Far IR
12. Large-Format, High-Resolution Focal Plane Arrays
13. Large-Format, Low-Darkrate, High-Efficiency, Photon-Counting, Solar-blind, Far- and Near-UV Detectors
14. Large-Format, Low-Noise and Ultralow-Noise Far-IR Direct Detectors
15. Long-Wavelength-Blocking Filters for X-ray Micro-Calorimeters
16. Low-Stress, High-Stability, X-ray Reflective Coatings
17. Mirror Technologies for High Angular Resolution (UV/Vis/Near IR)
18. Stellar Reflex Motion Sensitivity – Astrometry
19. Stellar Reflex Motion Sensitivity – Extreme Precision Radial Velocity
20. Vis/Near-IR Detection Sensitivity





## Strategic Technology Gap Priorities – Tier 2



1. Broadband X-ray Detectors
2. Compact, Integrated Spectrometers for 100 to 1000  $\mu\text{m}$
3. Far-IR Imaging Interferometer for High-Resolution Spectroscopy
4. Far-IR Spatio-Spectral Interferometry
5. Fast, Low-Noise, Megapixel X-ray Imaging Arrays with Moderate Spectral Resolution
6. High-Efficiency X-ray Grating Arrays for High-Resolution Spectroscopy
7. High-Resolution, Direct-Detection Spectrometers for Far-IR Wavelengths
8. Improving the Calibration of Far-IR Heterodyne Measurements
9. Large-Aperture Deployable Antennas for Far-IR/THz/sub-mm Astronomy for Frequencies over 100 GHz
10. Large-Format, High-Spectral-Resolution, Small-Pixel X-ray Focal-Plane Arrays
11. Polarization-Preserving Millimeter-Wave Optical Elements
12. Precision Timing for Space-Based Astrophysics
13. Rapid Readout Electronics for X-ray Detectors
14. Starshade Deployment and Shape Stability
15. Starshade Starlight Suppression and Model Validation
16. UV Detection Sensitivity

# Strategic Technology Gap Priorities – Tiers 3 to 5

## Tier 3:

1. Advancement of X-ray Polarimeter Sensitivity
2. Detection Stability in Mid-IR
3. Far-UV Imaging Bandpass Filters
4. High-Efficiency Far-UV Mirror
5. High-Efficiency, Low-Scatter, High- and Low-Ruling-Density, High- and Low-Blazed-Angle UV Gratings
6. High-Quantum-Efficiency, Solar-Blind, Broadband Near-UV Detector
7. Photon-Counting, Large-Format UV Detectors
8. Short-Wave UV Coatings
9. Warm Readout Electronics for Large-Format Far-IR Detectors

## Tier 4:

1. Advanced Millimeter-Wave Focal-Plane Arrays for CMB Polarimetry
2. Improving Photometric and Spectro-Photometric Precision of Time-Domain and Time-Series Measurements
3. UV/Opt/Near-IR Tunable Narrow-Band Imaging Capability
4. Very-Wide-Field Focusing Instrument for Time-Domain X-ray Astronomy

## Tier 5:

1. Complex Ultra-Stable Structures for Future Gravitational-Wave Missions
2. Disturbance Reduction for Gravitational-Wave Missions
3. Gravitational Reference Sensor
4. High-Performance Spectral Dispersion Component/Device
5. High-Power, High-Stability Laser for Gravitational-Wave Missions
6. Laser Phase Measurement Chain for a Decihertz Gravitational-Wave Mission
7. Micro-Newton Thrusters for Gravitational Wave-Missions
8. Stable Telescopes for Gravitational Wave-Missions





## What do These Priorities Mean?



- **The technology gap list informs SAT solicitation and selections; historically the focus has been on the first two priority tiers**
- **However, gaps in lower tiers are not ignored**
  - Technologies that address any gap, whether solicited in SAT or not, may fit in APRA
  - Gaps in lower tiers have at times moved to higher tiers in later cycles (a new Astrophysics Implementation Plan, AIP, is expected to be released by the end of 2022, and will inform the next cycle)
  - Astrophysics Division may decide to direct-fund technologies they deem important enough after considering programmatic aspects and technology strategy for the future great observatories

## Gaps by Program & Mission Supported

Program	Tier 1	Tier 2	Tier 3	Tier 4	Tier 5	Total	2019
COR	11	7	7	2	1	<b>28</b>	23
ExEP	6	3	1	0	0	<b>10</b>	12
PCOS	3	6	1	2	7	<b>19</b>	13
<b>Total</b>	<b>20</b>	<b>16</b>	<b>9</b>	<b>4</b>	<b>8</b>	<b>57</b>	<b>48</b>

Mission	Tier 1	Tier 2	Tier 3	Tier 4	Tier 5	Total
IROUV	11	3	6	1		<b>21</b>
Far-IR Flagship	6	8	2			<b>16</b>
X-ray Flagship	3	5	1			<b>9</b>
Far-IR Probe	6	8	2			<b>16</b>
X-ray Probe	3	5	1	1		<b>10</b>
CMB Probe				1		<b>2</b>
TDAMM				2		<b>2</b>
None					8	<b>8</b>
<b>Total</b>	<b>29</b>	<b>30</b>	<b>12</b>	<b>5</b>	<b>8</b>	<b>84</b>

Many gaps affect multiple missions, so the total is greater than 57.



# How Can You Participate?

- **If you're from a US institution, you can propose to any of the following opportunities:**

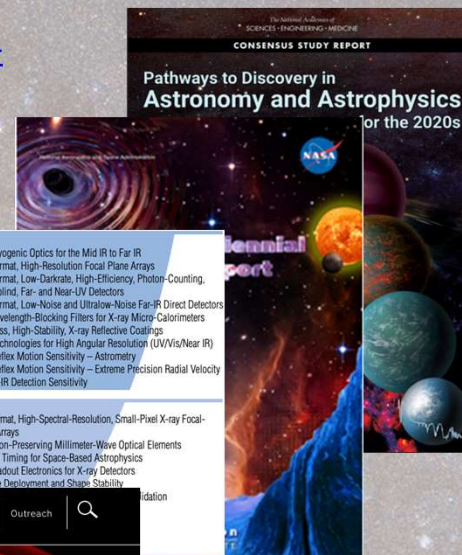
- APRA (<https://nspires.nasaprs.com/external/solicitations/summary.do?solId={8378F31B-1E87-B357-98A2-9E4237B0F1FB}&path=&method=init>)
- SAT (<https://nspires.nasaprs.com/external/solicitations/summary.do?solId={14958525-AA09-111E-59C2-5BBF6585F7AA}&path=&method=init>)
- RTF (<https://nspires.nasaprs.com/external/solicitations/summary.do?solId={1EF0CA93-73C0-F40B-CE95-96647D403CA6}&path=&method=init>)
- FINESST (<https://nspires.nasaprs.com/external/solicitations/summary!init.do?solId=%7b87947100-56AE-C4DC-C511-0349862D658A%7d&path=open>)

- **Important Due Dates for APRA and SAT**

- Mandatory Notice of Intent (NOI) is due by October 21, 2022
- Proposals are due by December 15, 2022



- Decadal Survey: <https://www.nationalacademies.org/our-work/decadal-survey-on-astronomy-and-astrophysics-2020-astro2020>
- ABTR: [https://apd440.gsfc.nasa.gov/images/tech/2022\\_ABTR.pdf](https://apd440.gsfc.nasa.gov/images/tech/2022_ABTR.pdf)
- Tech gap list: [https://apd440.gsfc.nasa.gov/tech\\_gap\\_priorities.html](https://apd440.gsfc.nasa.gov/tech_gap_priorities.html)
- Tech gaps details: [https://apd440.gsfc.nasa.gov/tech\\_gap-descriptions.html](https://apd440.gsfc.nasa.gov/tech_gap-descriptions.html)



<b>Tier 1 Technology Gaps</b> Advanced Cryocoolers Coronagraph Contrast and Efficiency Coronagraph Stability Cryogenic Readouts for Large-Format Far-IR Detectors Heterodyne Far-IR Detector Systems High-Performance, Sub-Kelvin Coolers High-Reflectivity Broadband Far-UV-to-Near-IR Mirror Coatings High-Resolution, Large-Area, Lightweight X-ray Optics High-Throughput Bandpass Selection for UV/Vis High-Throughput, Large-Format Object Selection Technologies for Multi-Object and Integral Field Spectroscopy	Large Cryogenic Optics for the Mid IR to Far IR Large-Format, High-Resolution Focal Plane Arrays Large-Format, Low-Darkrate, High-Efficiency, Photon-Counting, Solar-blind, Far- and Near-UV Detectors Large-Format, Low-Noise and Ultralow-Noise Far-IR Direct Detectors Long-Wavelength-Blocking Filters for X-ray Micro-Calorimeters Low-Stress, High-Stability, X-ray Reflective Coatings Mirror Technologies for High Angular Resolution (UV/Vis/Near IR) Stellar Reflex Motion Sensitivity – Astrometry Stellar Reflex Radial Velocity Ultra-Near-IR Detection Sensitivity
<b>Tier 2 Technology Gaps</b> Broadband X-ray Detectors Compact, Integrated Spectrometers for 100 to 1000 $\mu$ m Far-IR Imaging Interferometer for High-Resolution Spectroscopy Far-IR Spatio-Spectral Interferometry Fast, Low-Noise, Megapixel X-ray Imaging Arrays with Moderate Spectral Resolution	Large-Format, High-Spectral-Resolution, Small-Pixel X-ray Focal-Plane Arrays Polarization-Preserving Millimeter-Wave Optical Elements Precision Timing for Space-Based Astrophysics Rapid Readout Electronics for X-ray Detectors Standardized Deployment and Storage Stability

 <div>Overview   Technology   Outreach   Search</div>							
Astrophysics Program Offices <b>2022 Astrophysics Strategic Technology Gaps</b> <small>TECHNOLOGY GAPS / OVERVIEW / TECH GAP PRIORITIES / PRIORITIZATION PROCESS / TECH GAP DESCRIPTIONS</small>							
Gap Name	Description	Current State-of-the-Art	TRL	Performance Goals and Objectives	Scientific, Engineering, and/or Programmatic Benefits	Applications and Potential Relevant Astrophysics Missions	Urgency
Cryogenic Readouts for Large-format Far-IR Detectors	Readout schemes including cryogenic multiplexing for arrays of large-format Far-IR detectors need to be developed	Readout schemes using HEMT or SiGe amplifiers and frequency-multiplexed resonant circuits are in development. A few hundred channels per 1 mm <sup>2</sup> HEMT amplifier have been demonstrated. Low power dissipation at 4 K is required. For TES-based detectors a microwave SQUID multiplexer using frequency division, time division or code division multiplexing is needed. Frequency division multiplexing is well advanced and can meet the needs of a far-IR flagship when scaled to 2000 pixels (resonators) per a GHz channel.	3	Near-term, this scheme should result in 2000 pixels per amplifier channel (enabling) 3000 pixels/channel (enhancing) HEMT amplifiers from Low Noise Factory can achieve 10 dB with 0.38 mW of dissipation at 4 K.	Sensitivity reduces observing times from many hours to a few minutes (100x faster), while array format increases areal coverage by ~10-100x. Overall mapping speed can increase by factors of thousands. Sensitivity enables measurement of low-surface-brightness debris disks and protoplanets with an interferometer. This is enabling technology. Suborbital and ground-based platforms can be used to validate technologies and advance TRL of new detectors.	Far-IR Flagship Far-IR Probe FIR detector technology is an enabling aspect of all future FIR mission concepts, and is essential for future progress. This technology can improve science capability at a fixed cost much more rapidly than larger telescope sizes. This development serves Astrophysics almost exclusively (with some impact on planetary and Earth studies). Many synergies exist with similar developments for x-ray microcalorimeters	Required TRL 6 by mission PDR. Extreme stretch. Single tech