



Moon Burst Energetics All-sky Monitor

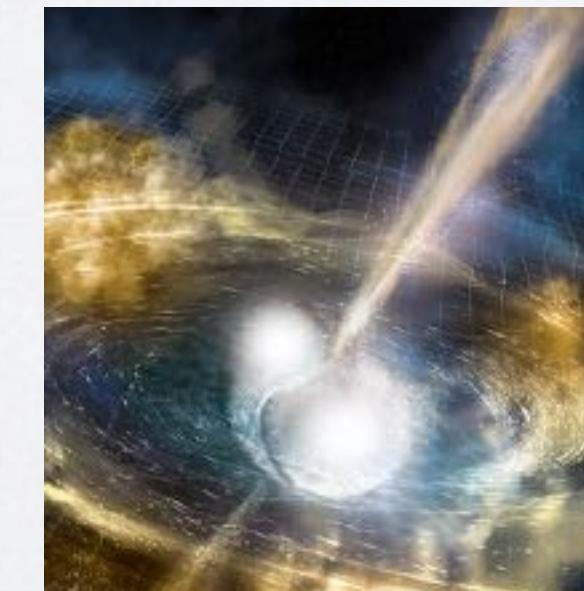


OVERVIEW

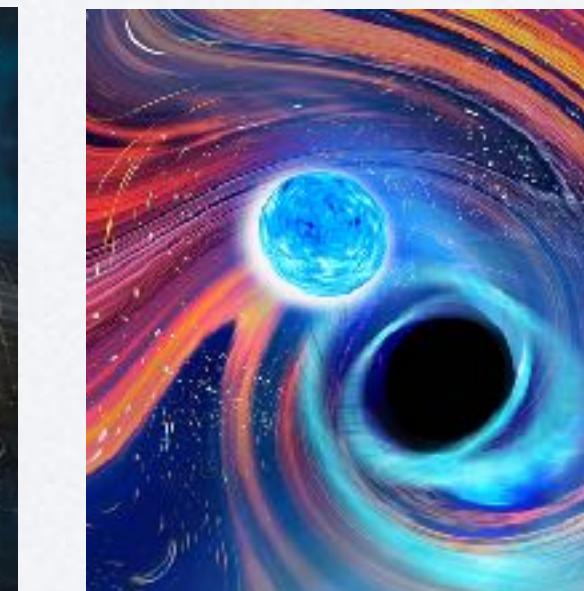


- Moon Burst Energetics All-sky Monitor is 3-year gamma-ray mission in cislunar orbit to explore the behavior of matter and energy under extreme conditions by observing relativistic astrophysical explosions.
- MoonBEAM provides key capabilities that are difficult to achieve in Low Earth Orbit:
 - instantaneous all-sky gamma-ray field of view
 - uninterrupted observations with >96% duty cycle
 - background radiation stability
- 3 years of mission operation will provide observations of:
 - 1600 binary compact mergers
 - 5900 optically discovered core collapse supernovae
 - 140 magnetar giant flares
 - and enables 55 very high energy gamma-ray and 360 optical follow-up

**Astro2020 Decadal Survey:
Astronomical Transient Events**
 “Higher sensitivity all-sky monitoring of the high-energy sky, complemented by capabilities in the optical such as from Kepler and TESS, is a critical part of our vision for the next decade in transient and multi-messenger astronomy.”



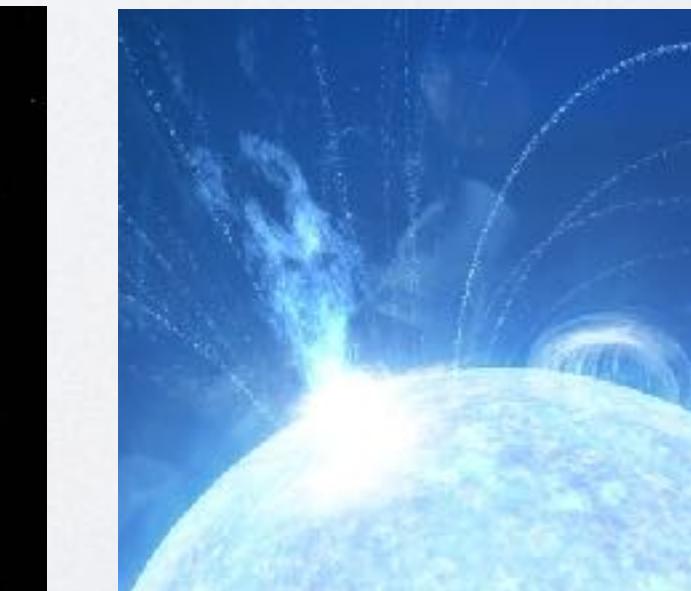
Credit: NSF/LIGO/Sonoma State Univ./A. Simonnet



Credit: Carl Knox, OzGrav-Swinburne University



Credit: CNRS/C



Credit: NASA GSFC/Chris Smith (USRA/GESTAR)

RELATIVISTIC TRANSIENTS



• Gamma-ray Bursts (GRBs)

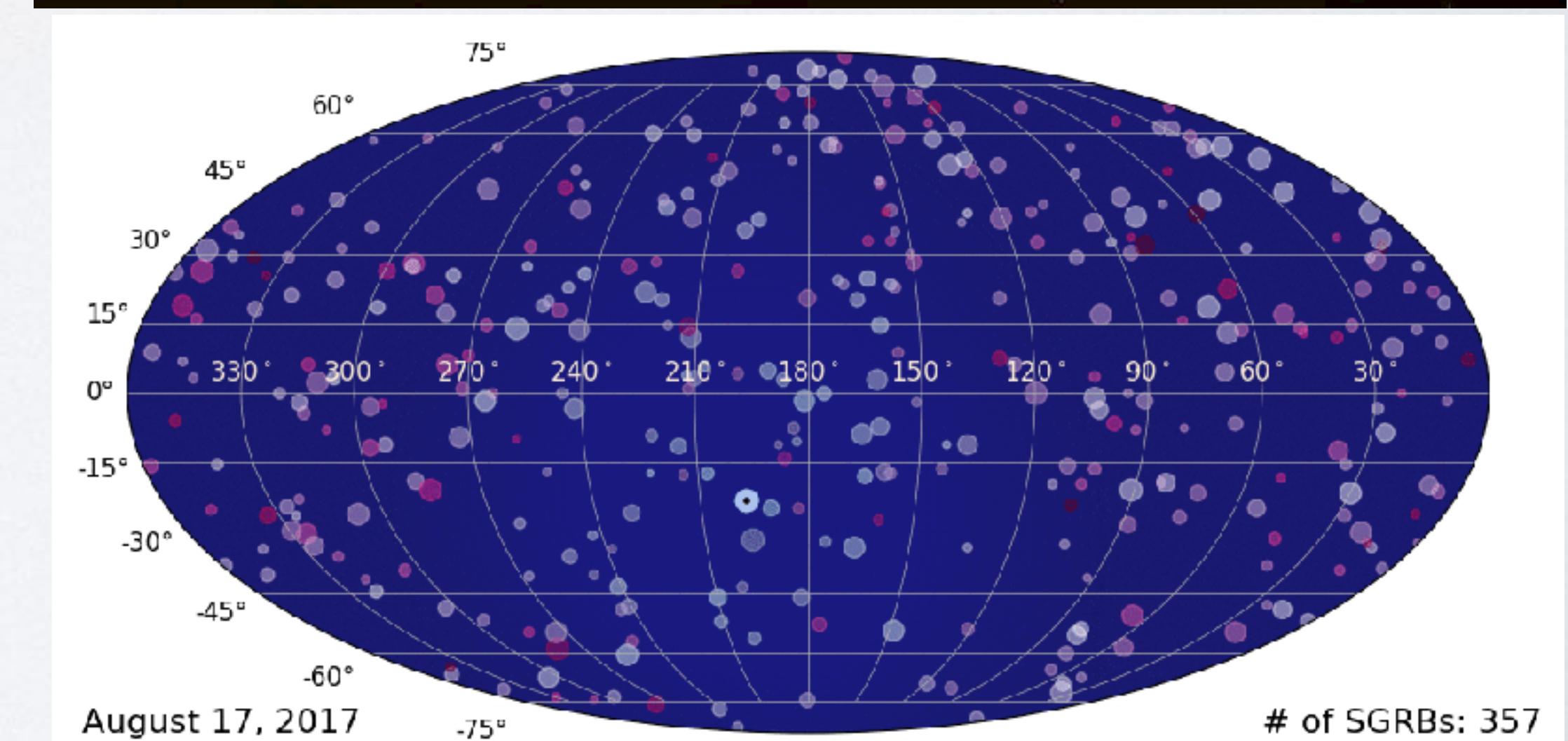
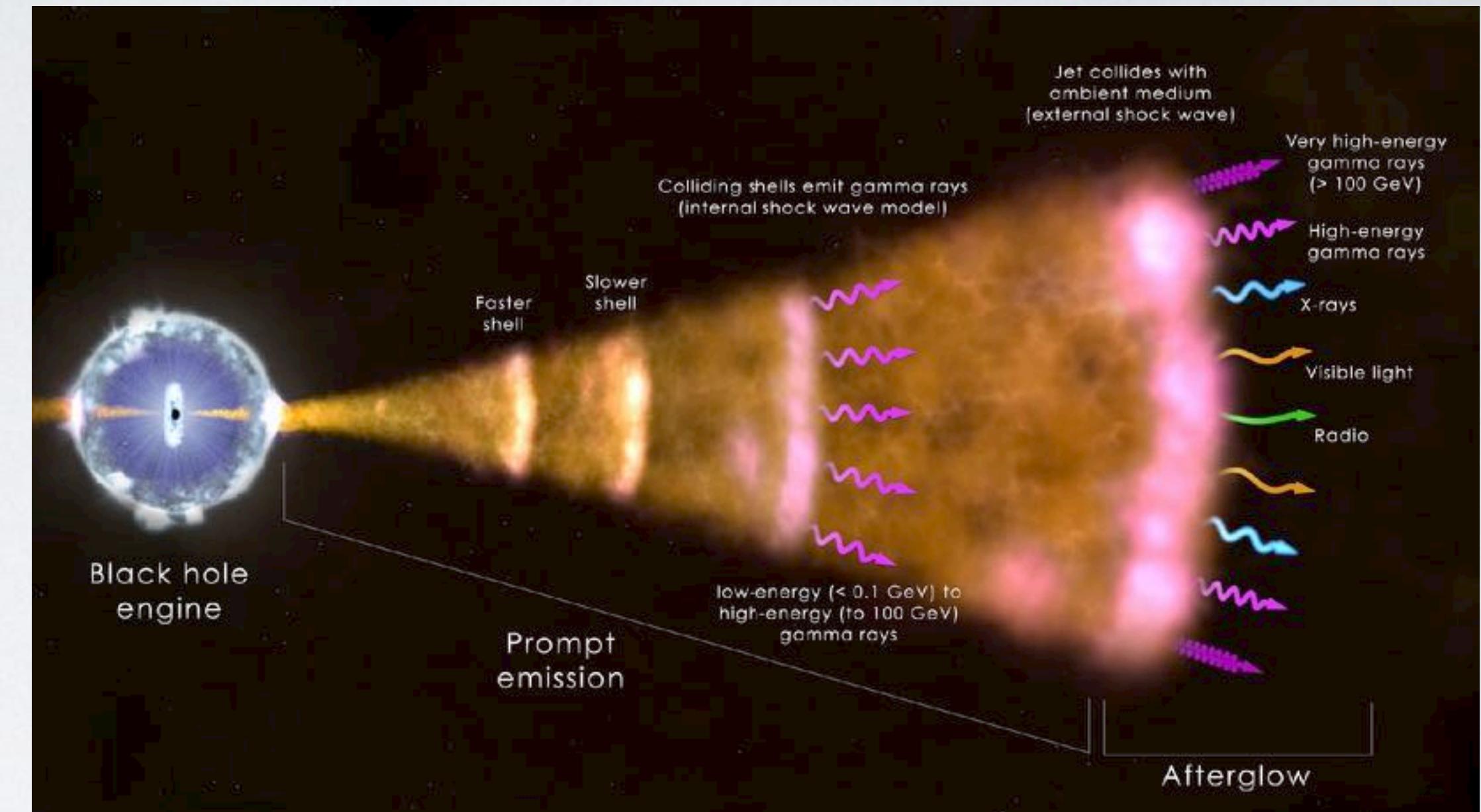
- most energetic explosions in the Universe.
- detectable in all wavelengths from radio to gamma rays.
- can generate multi-messenger signals: gravitational waves, neutrinos, and cosmic rays.

• Transient nature

- prompt emission in gamma rays, lasting <1s to >100 s.
- afterglow starting within minutes and can last up to years.
- detectable ~once per day, distributed all over the sky.

• Era of Multi-Messenger Astrophysics

- 2017-08-17: The merger of two neutron stars was detected in both gravitational waves and gamma rays, and subsequent kilonova and afterglow detection across the entire electromagnetic spectrum.
- Open questions remain such as merger and jet geometry, intrinsic properties etc., progress requires a population of joint detections.



RELATIVISTIC TRANSIENTS



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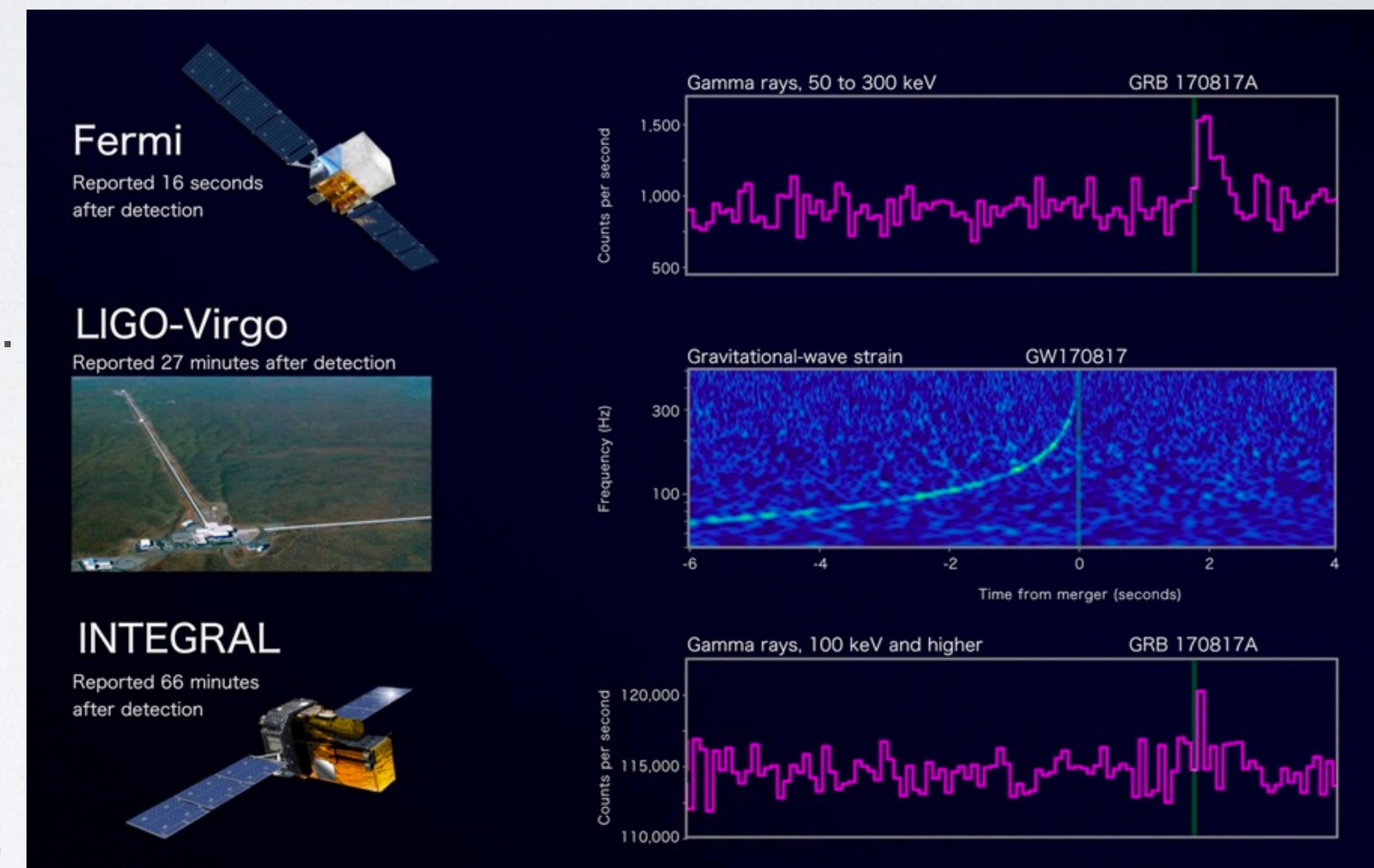
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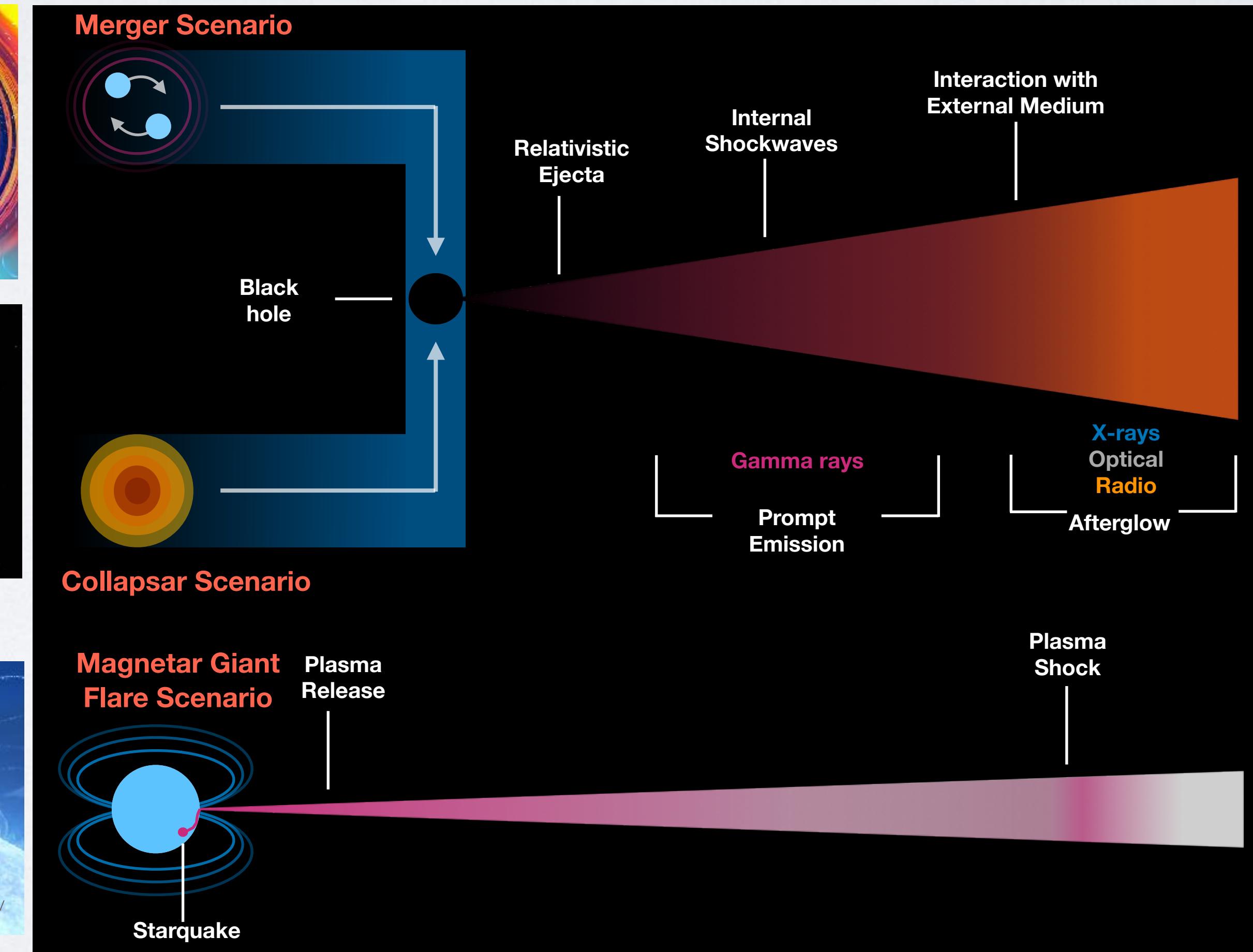
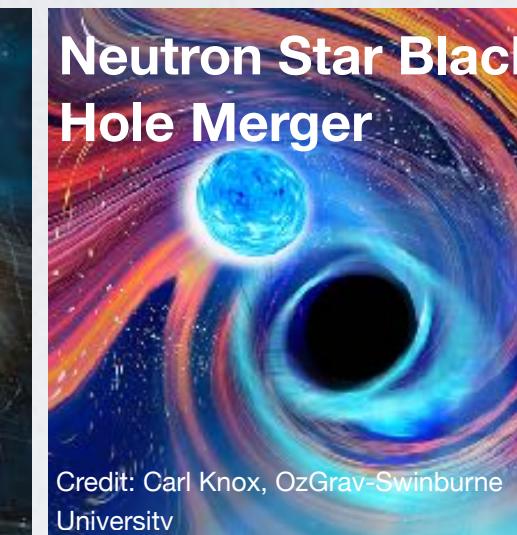
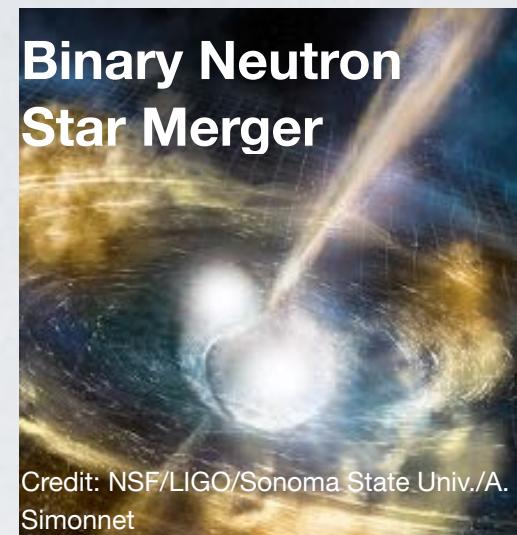
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MISSION GOAL AND OBJECTIVES



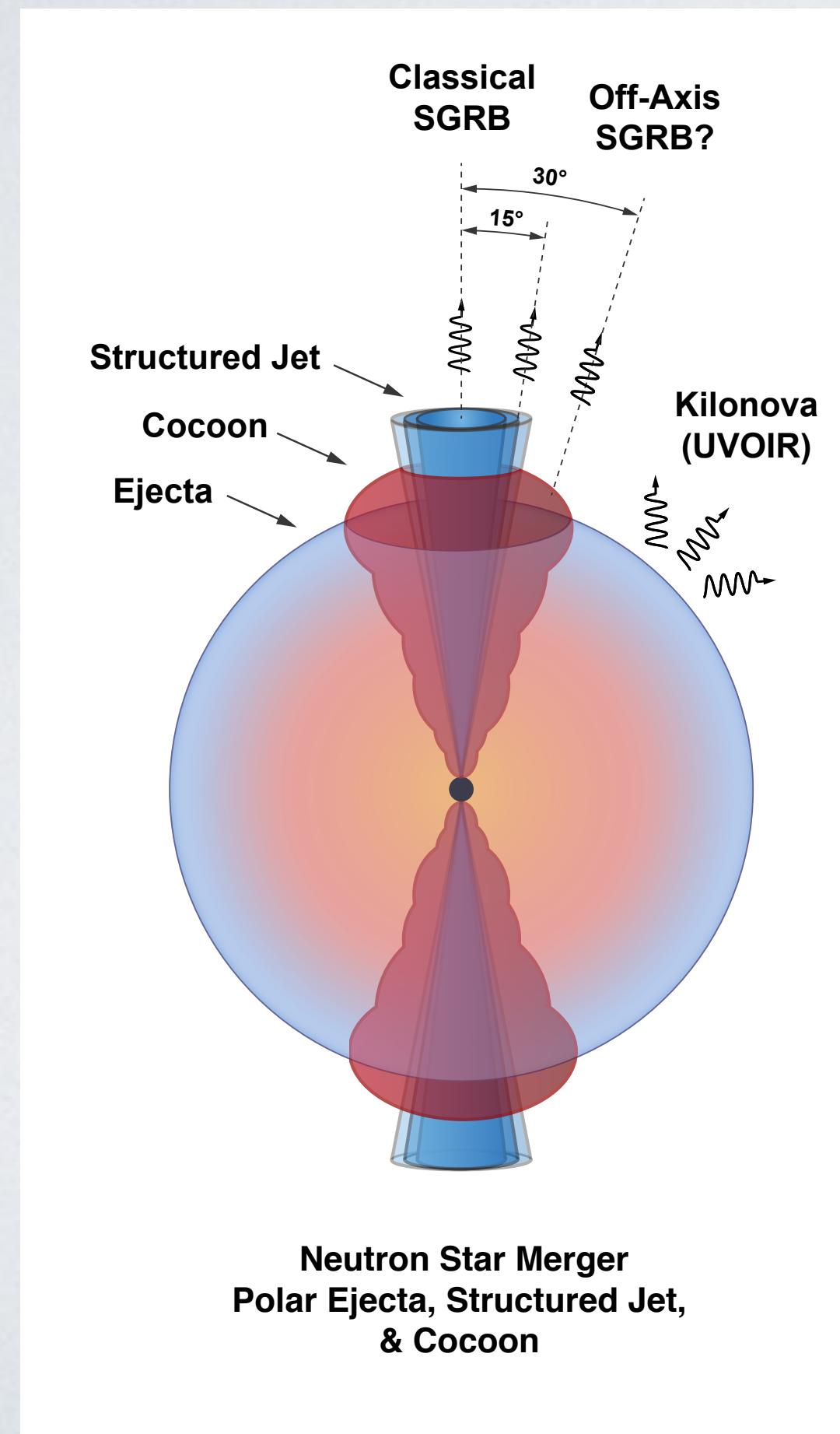
- Mission Goal: Explore the behavior of matter and energy in its most extreme environments
 - Objective 1: Characterize the progenitors of gamma-ray bursts and their multi-messenger and multi-wavelength signals
 - Objective 2: Identify conditions necessary to launch a transient astrophysical jet
 - Objective 3: Determine the origins of the observed high-energy emission within the relativistic outflow



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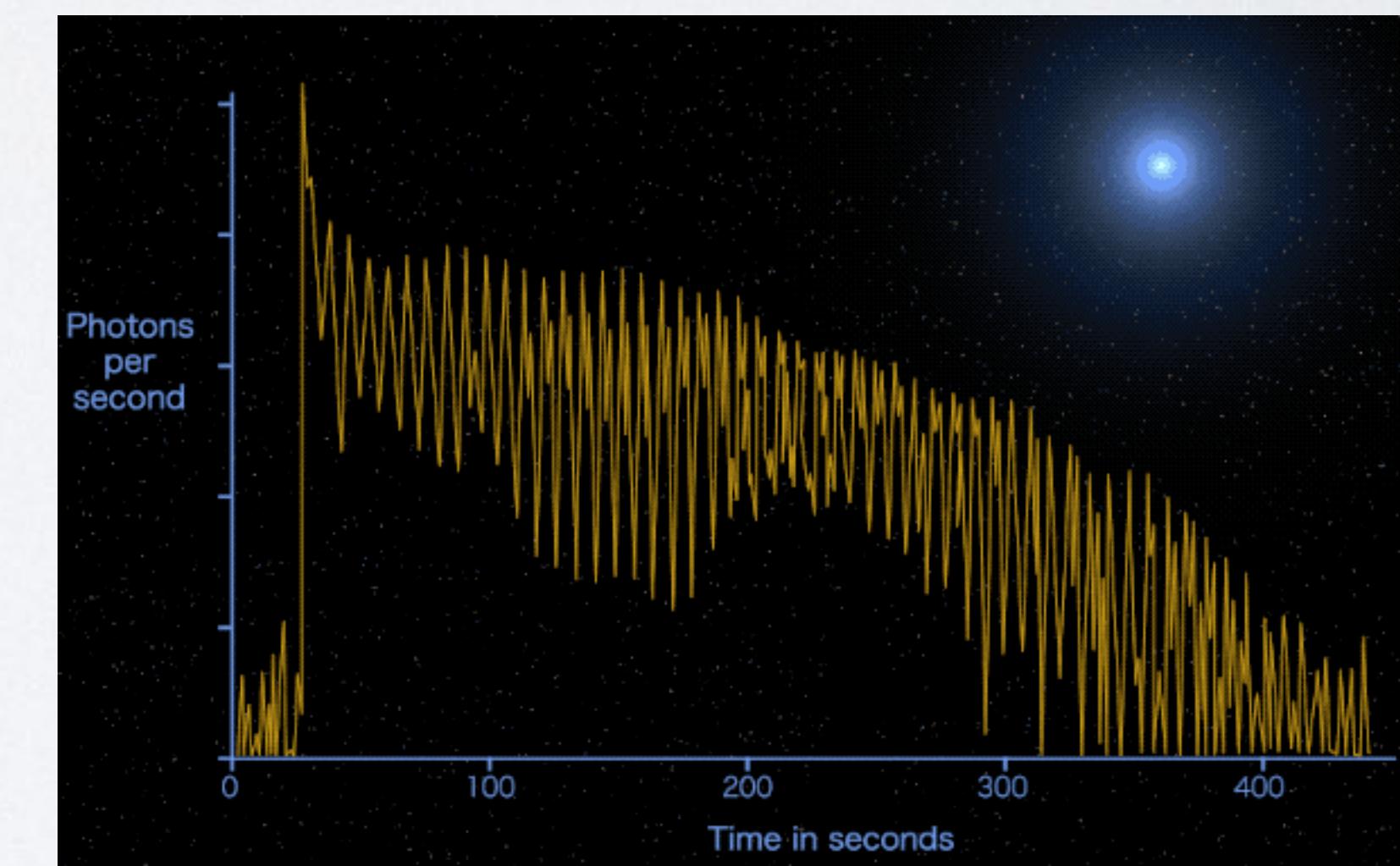
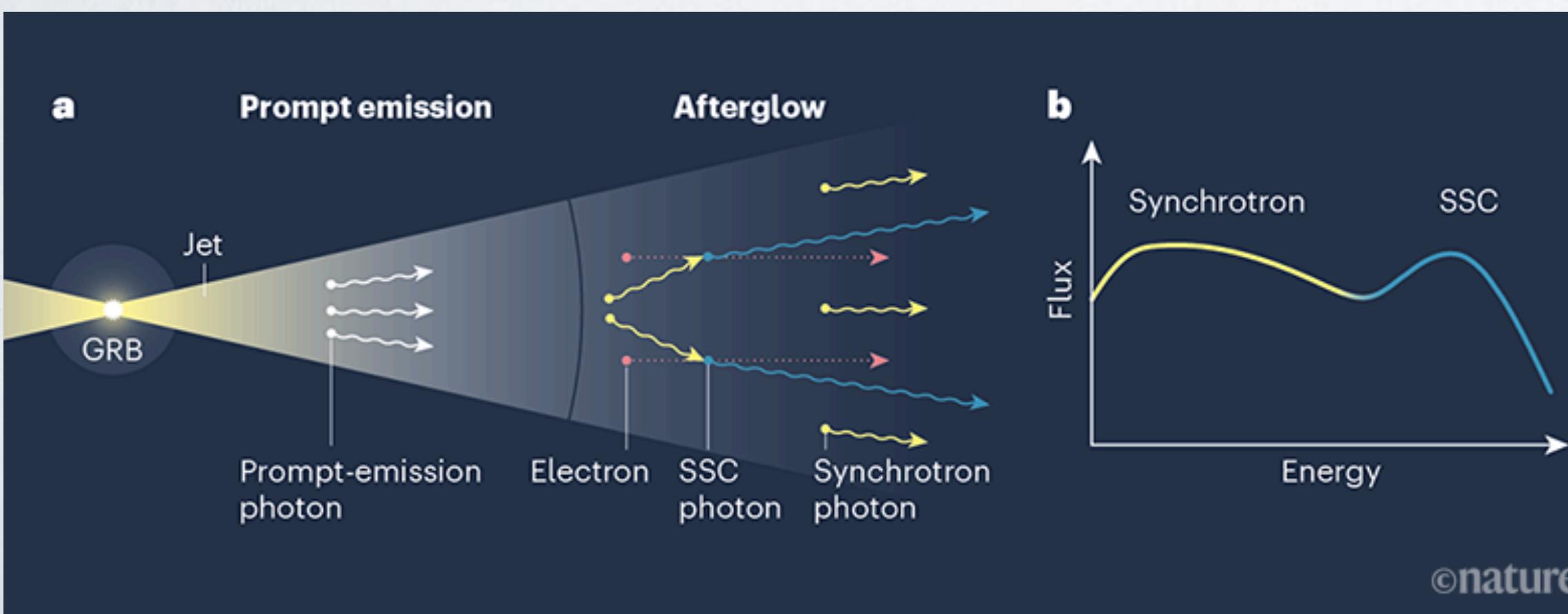
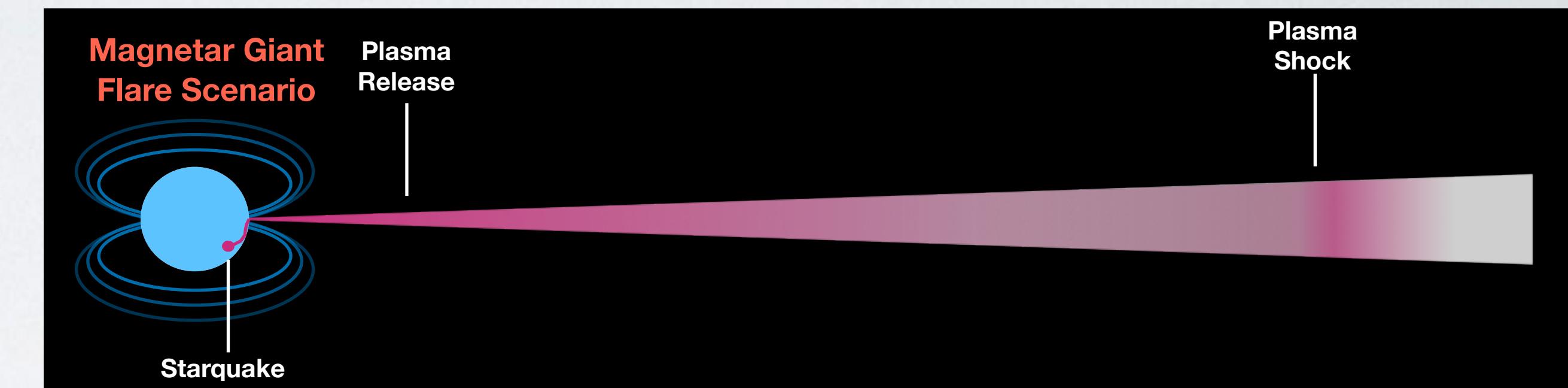
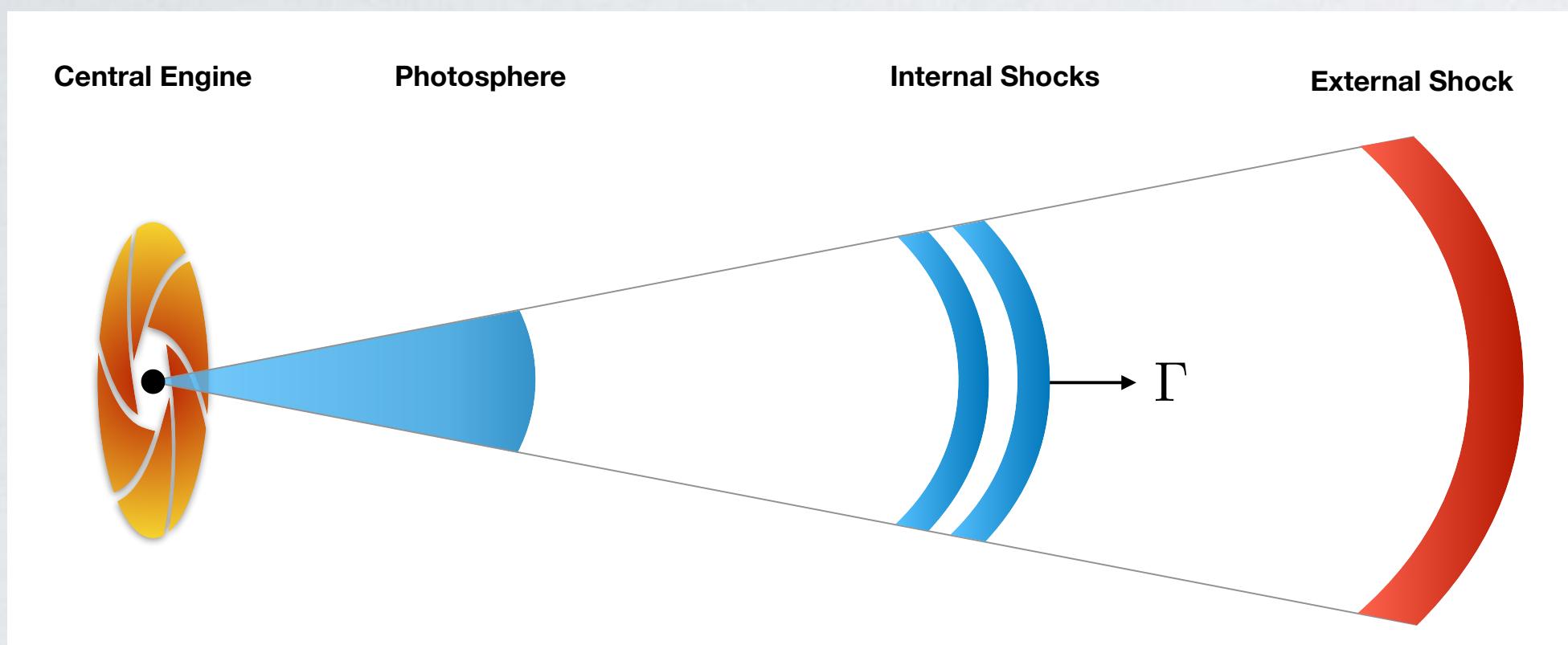
- A spectrum of jets, from completely failed (choked) to ultra-relativistic.
- Jet launch mechanisms:
 - magnetic (Blandford-Znajek mechanism)
 - neutrino - antineutrino annihilation
- Central engine powering the jet with the observed temporal and spectral properties:
 - black hole
 - magnetar?

Astro2020 Decadal Survey: “Understanding the central engines (newly formed compact objects like magnetars and BHs) that power many explosive transients continues to be a fundamental astrophysical challenge.”

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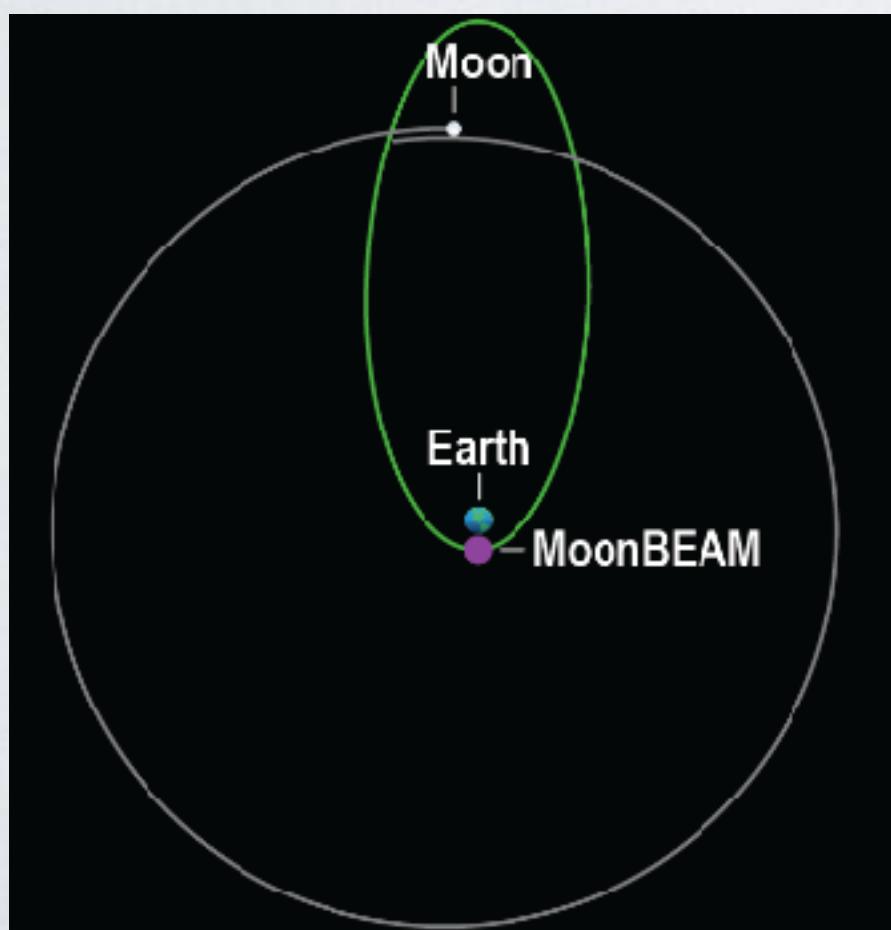
MISSION GOAL AND OBJECTIVES



- Mission Goal: Explore the behavior of matter and energy in its most extreme environments
 - What are the physical characteristics of stellar explosions that lead to a relativistic transient?
 - What conditions lead to the range of jet scenarios, from a failed jet to an ultra-relativistic jet?
 - What are the different emission mechanisms that convert the relativistic outflow into radiation?
 - What is the distribution of outflow widths and what determines the outflow width?
 - What is the velocity distribution of ejecta across the transverse axis of the outflow?
- Key open questions from the 2019 GW-EM task force report:
 - What conditions are necessary to produce relativistic jets, and what is their composition/structure?
 - Do black hole - neutron star and binary black hole mergers produce electromagnetic signals?
 - Can binary neutron star mergers reproduce the relative and total abundances of heavy (r-process) elements?
 - What is the current expansion rate of the Universe (Hubble constant)?
 - What is the equation of state of dense nuclear matter?

Addressed by MoonBEAM
Enabled by MoonBEAM

MISSION DESIGN

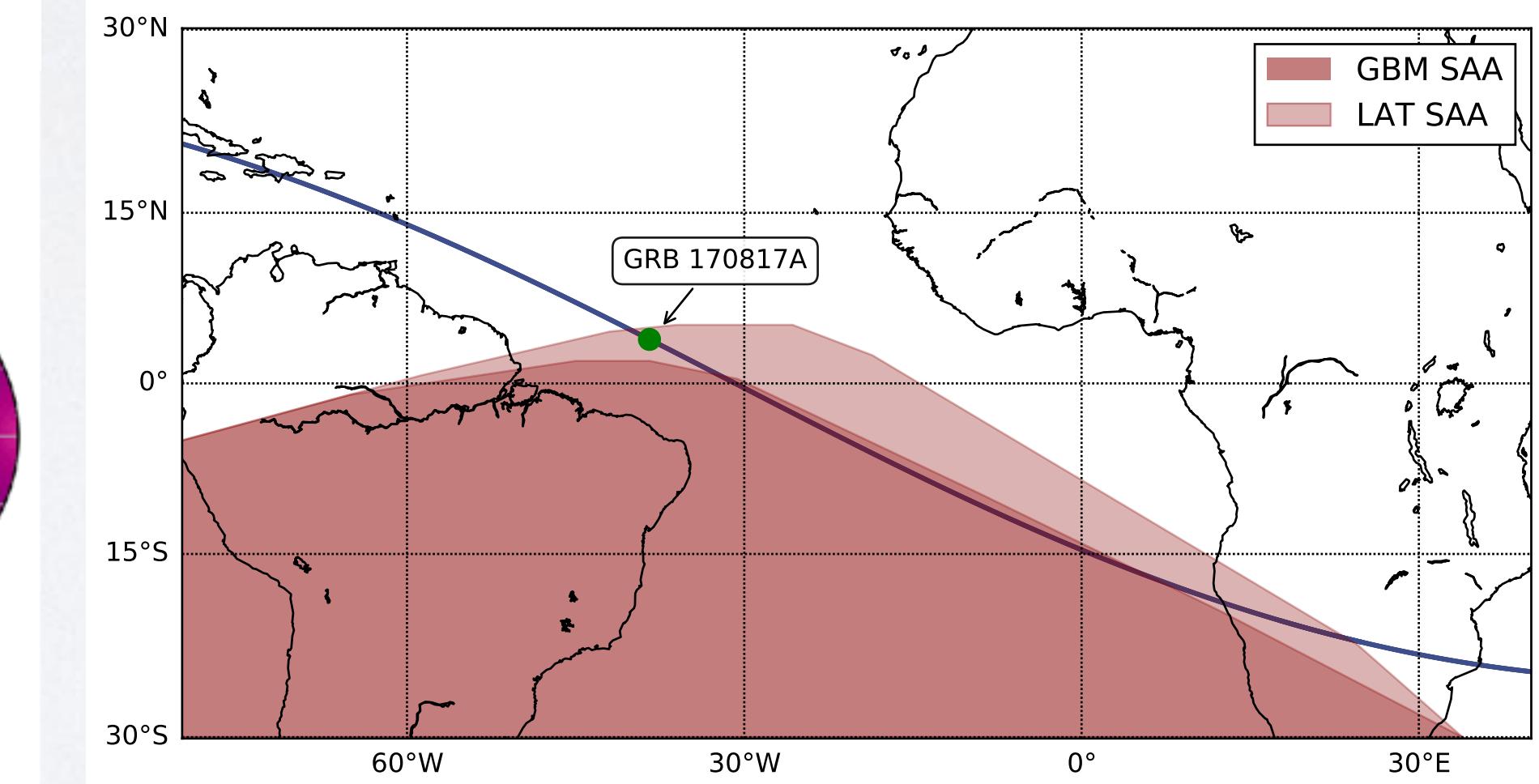
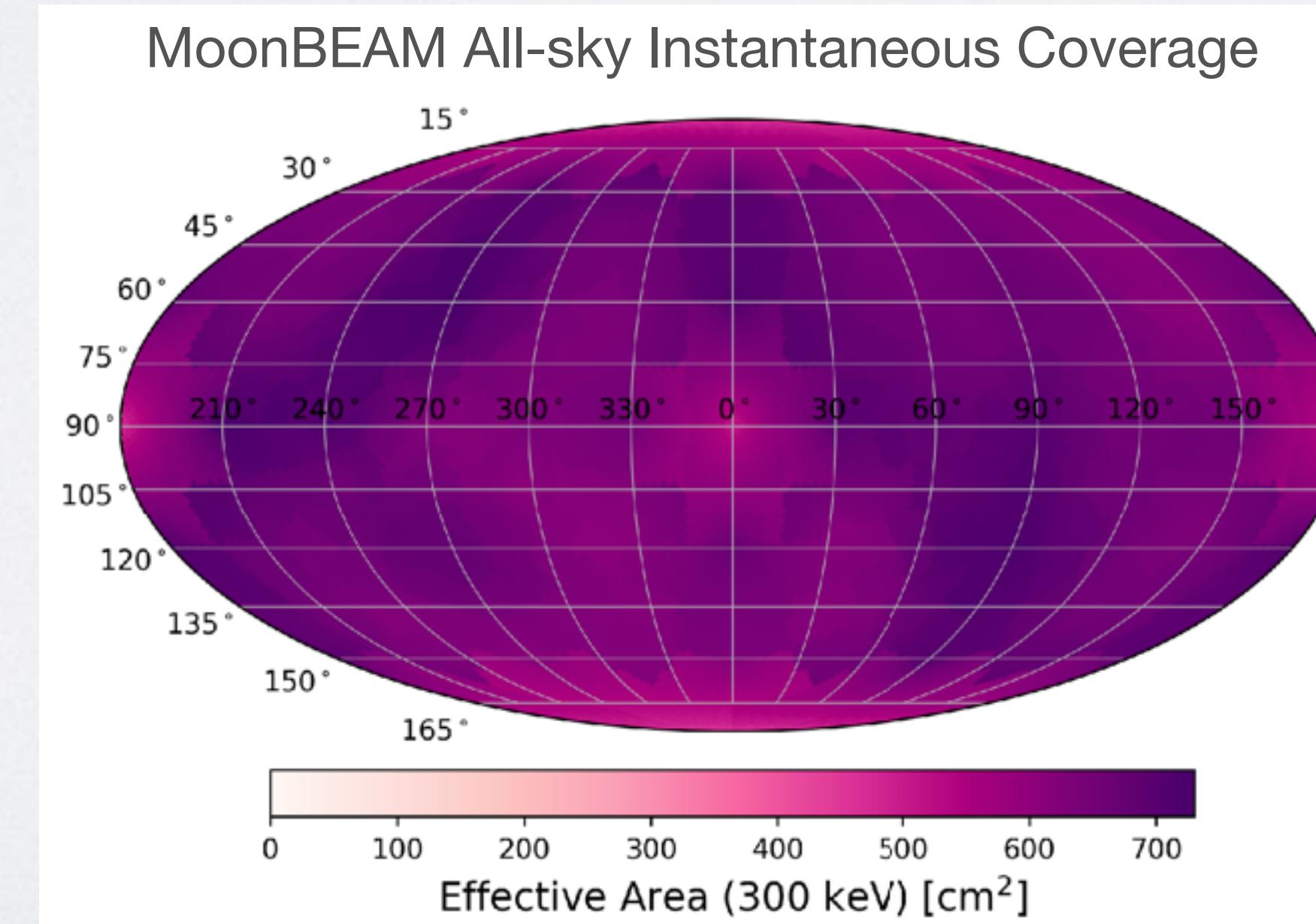
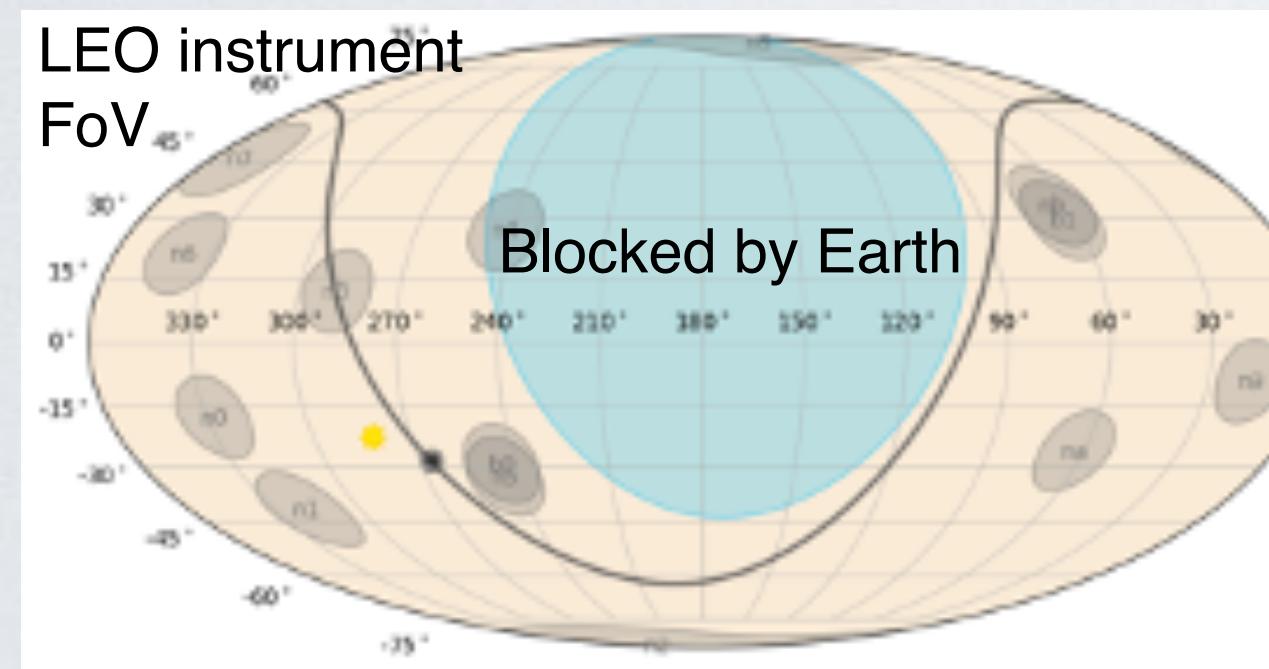
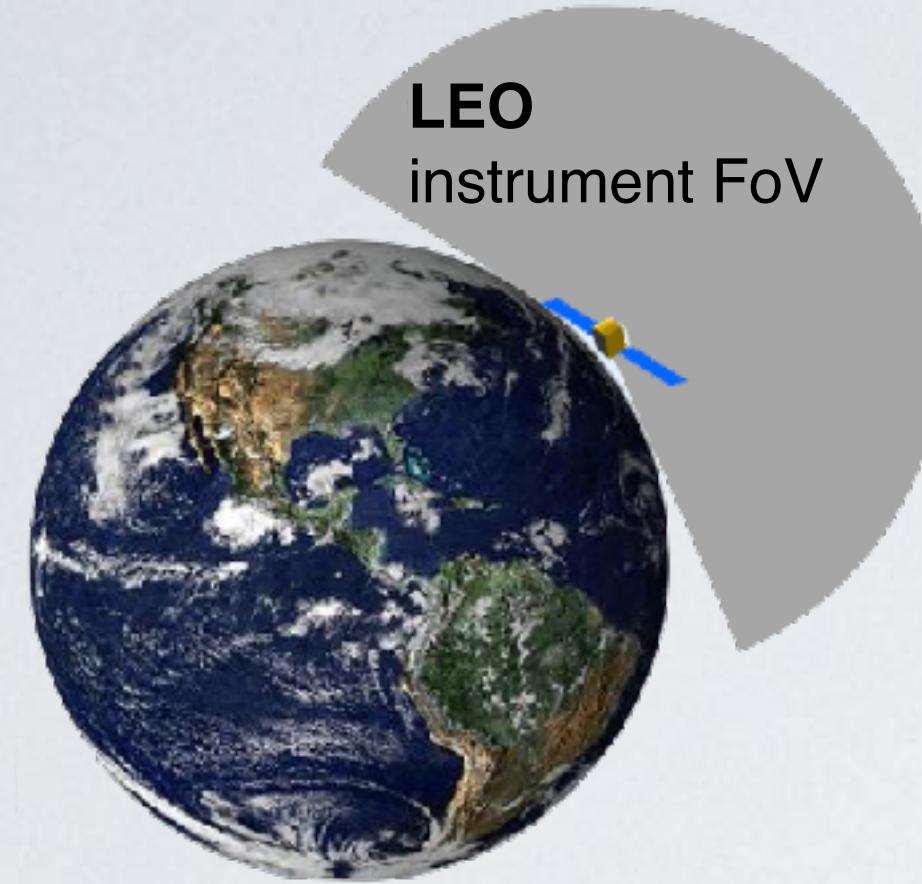


- Lockheed Martin SmallSat spacecraft bus
 - reusing >90% of high-maturity Lunar Trailblazer design.
 - compatible with ESPA Grande mass and volume constraint.
 - high-heritage deep space propulsion approach to lunar resonant orbit from *any* Geosynchronous Transfer Orbit (GTO) rideshare launch.
- Orbital distance up to 460,000km from Earth (**1.5 light-seconds**).
- Orbital period of **13.7 days**.
- Mission lifetime of **3 years**, launch ready 2027.
- Communication
 - **continuous burst alert coverage** with dedicated ground stations.
 - daily data downlink with the Near Space Network.

MISSION CAPABILITY



- Orbital distance 22,000km to 460,000km from Earth (up to 1.5 light-seconds).
 - **Instantaneous all-sky field of view:** Earth occults $\sim 2\%$ of the sky at closest approach, $<<1\%$ on average.
 - **high duty cycle >96%, 13+ days uninterrupted livetime:** no passage through the South Atlantic Anomaly (SAA).
 - **more stable background** compared to Low Earth Orbit: no atmospheric scattering and SAA-related radiation.
 - **additional localization improvement** using timing triangulation technique with other gamma-ray missions.

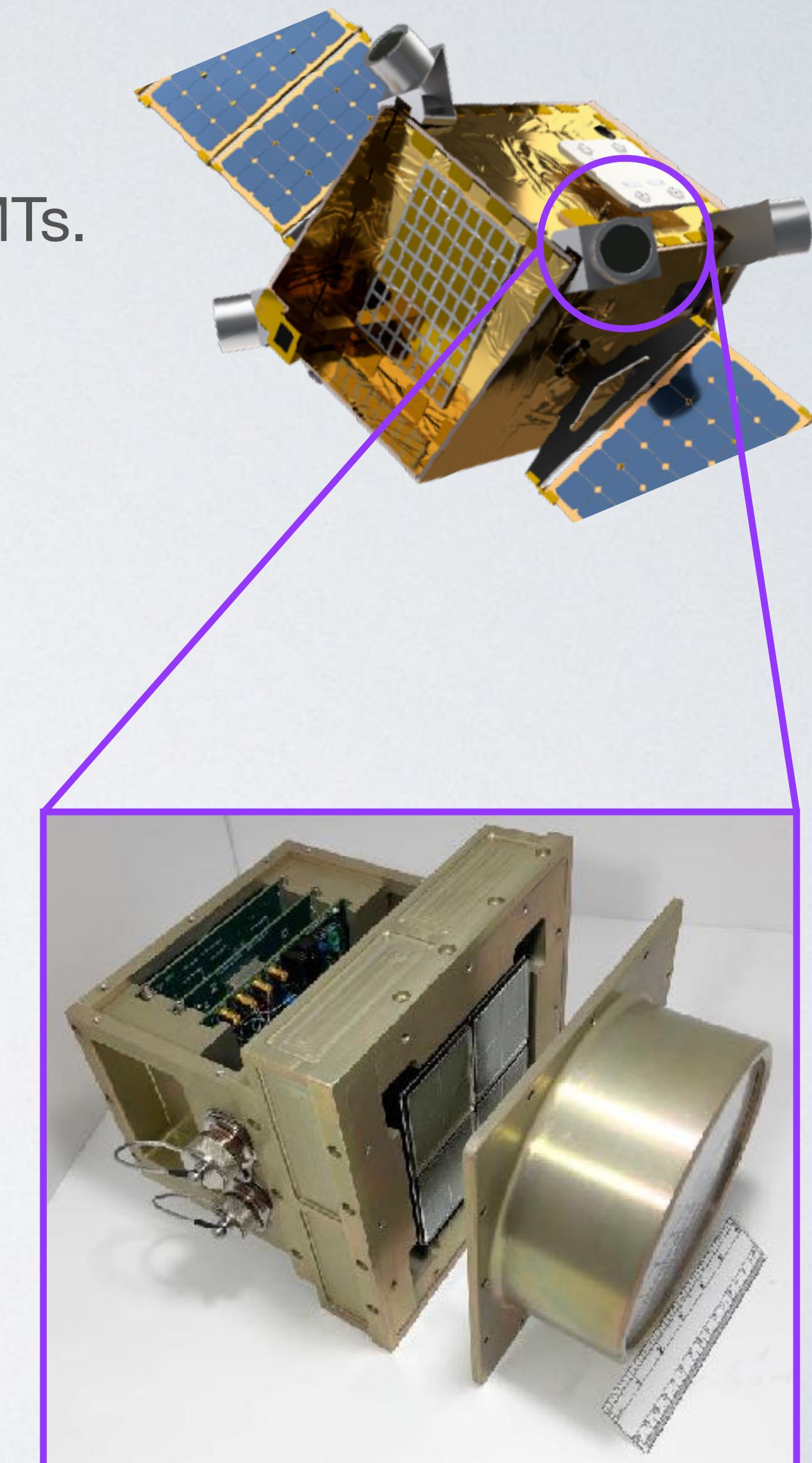
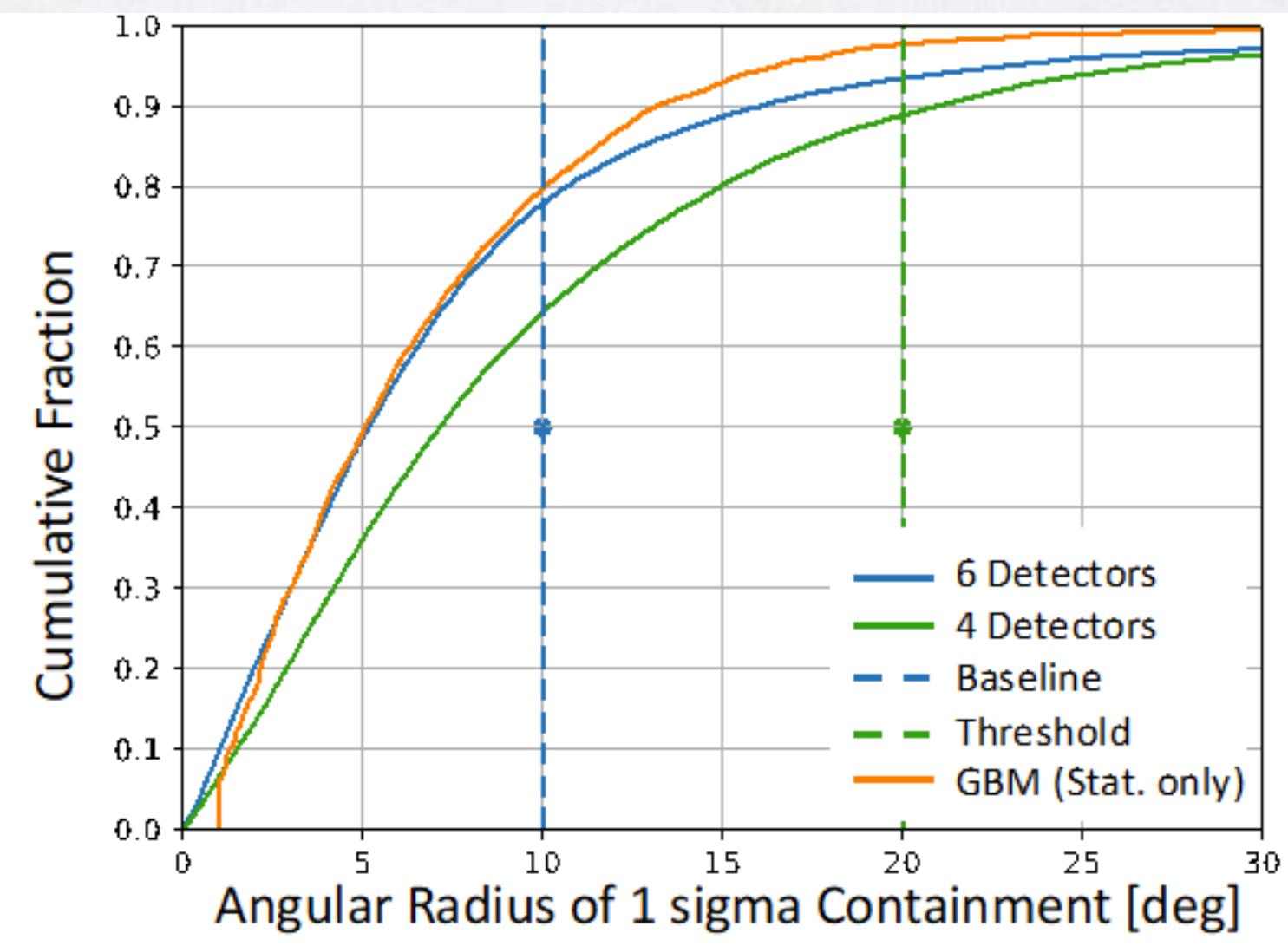
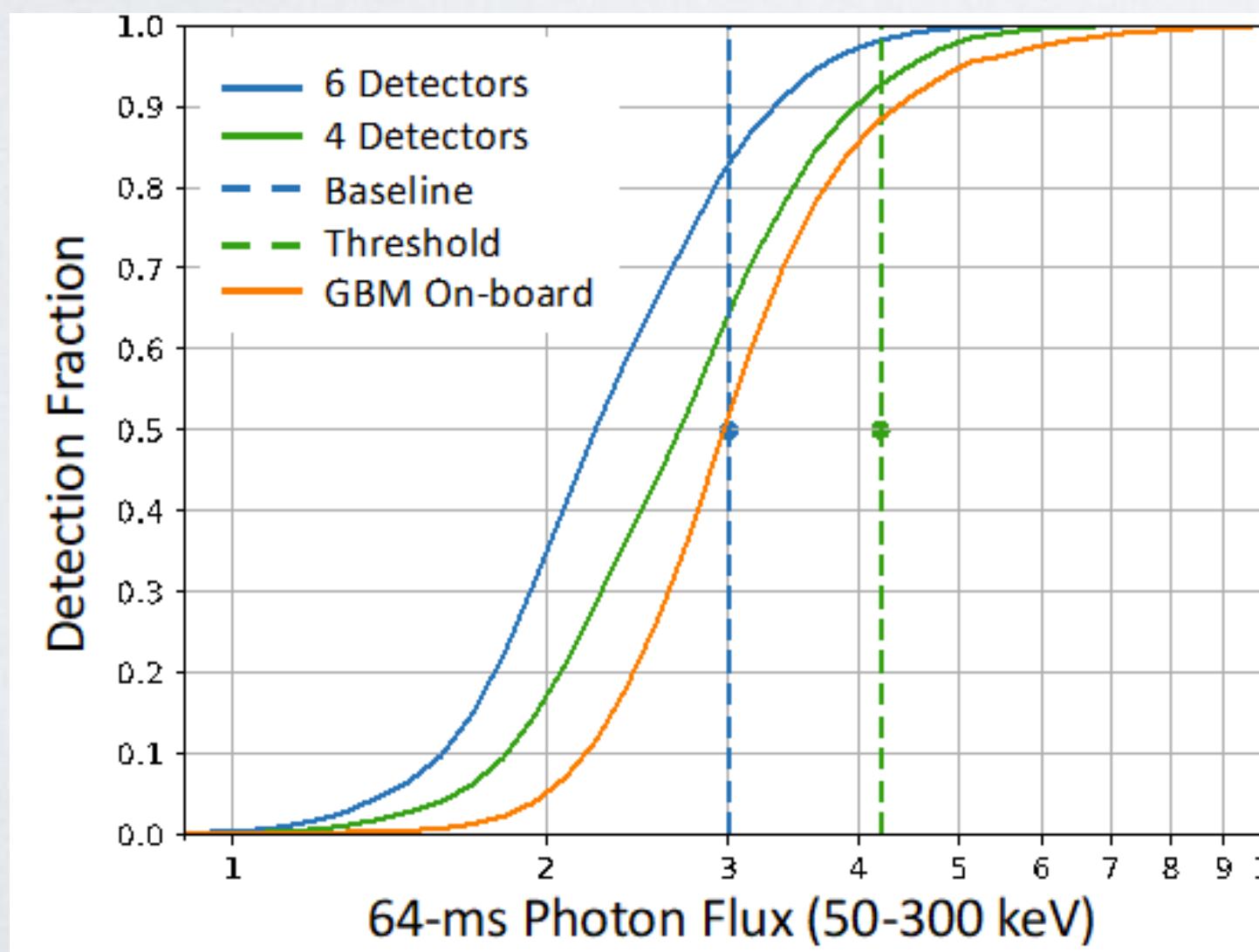


Fermi-GBM turned off for SAA 2 minutes after GRB 170817A / GW170817.

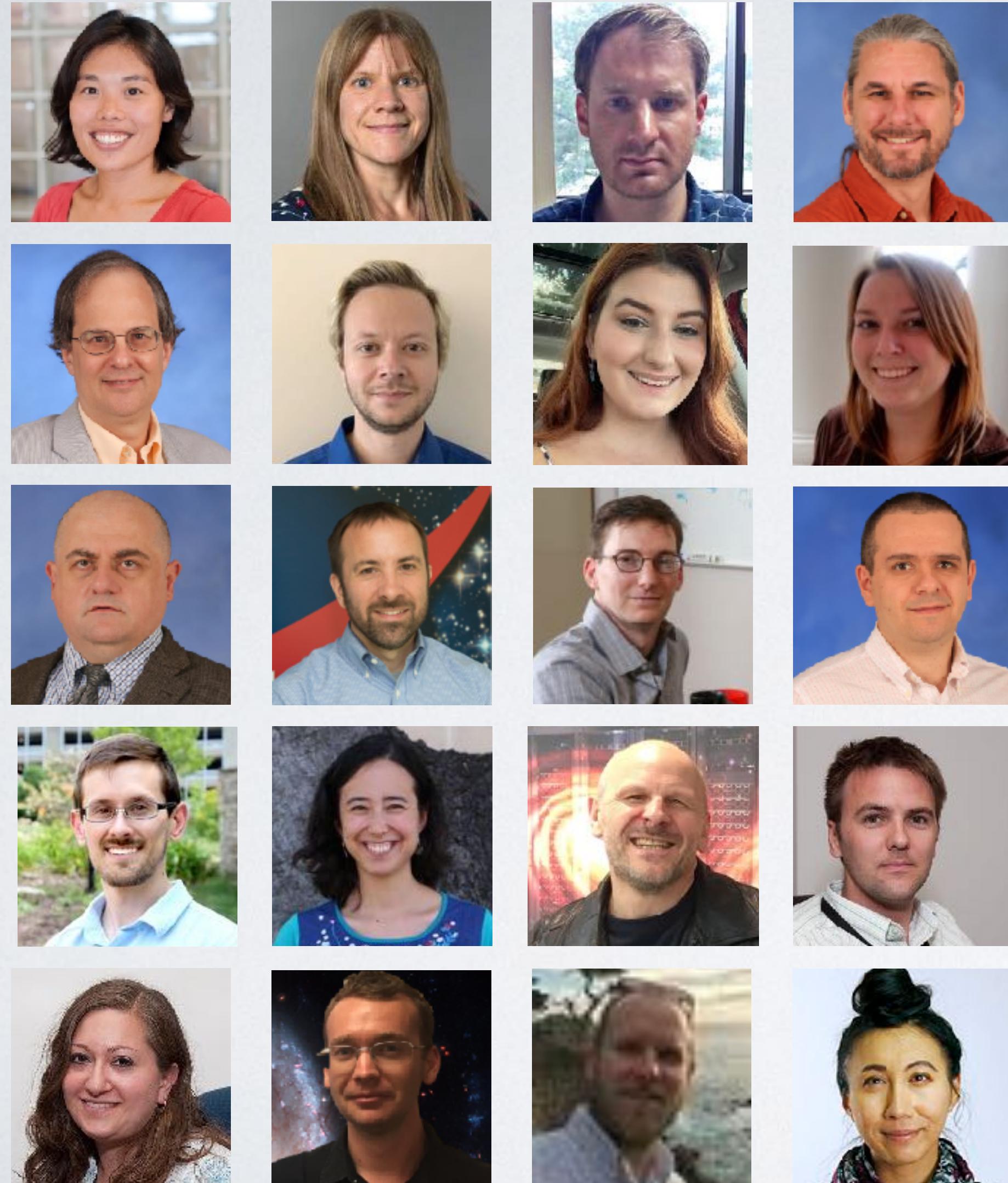
INSTRUMENT PERFORMANCE



- 6 scintillating detectors positioned for **instantaneous all-sky coverage, no pointing needed**.
- each detector module consists of a NaI(Tl)/CsI(Na) phoswich and flat panel PMTs.
- phoswich design enables simultaneous dual-mode observations:
 - low background, direction dependency for localization
 - ▶ pulse discrimination identifies origin >96% for background rejection
 - wide energy range and wide field-of-view for spectroscopy
 - ▶ 10–5000 keV, prompt GRB peak energy range
 - ▶ 10% energy resolution at 662 keV



SCIENCE TEAM



PI: C. Michelle Hui (MSFC)

Deputy PI: Colleen Wilson-Hodge (MSFC)

Project Scientist: Adam Goldstein (USRA)

Instrument Scientist: Peter Jenke (UAH)

Science Team members:

Michael Briggs (UAH)

Eric Burns (LSU)

Sarah Dalessi (UAH)

Corinne Fletcher (USRA)

Boyan Hristov (UAH)

Daniel Kocevski (MSFC)

Tyson Littenberg (MSFC)

Peter Veres (UAH)

Joshua Wood (MSFC)

Anna Ho (Cornell)

Eric Howell (UWA)

Jeremy Perkins (GSFC)

Judith Racusin (GSFC)

Oliver Roberts (USRA)

Jacob Smith (NRL)

Sylvia Zhu (DESY Zeuthen)

SUMMARY



MoonBEAM provides essential gamma-ray observations of relativistic astrophysical transients with the following capabilities:



- **instantaneous all-sky** field of view from lunar resonant orbit.
- **13+ days of uninterrupted livetime.**
- **stable background** for ultra long duration GRBs.
- sensitive to **prompt GRB emission** energy range, with broad coverage for spectroscopy.
- **independent localization and longer baseline** for additional localization improvement with other gamma-ray missions.
- **rapid alerts** to the astronomical community for contemporaneous and follow-up observations.
- planned launch in ~2027, overlapping with upcoming new capabilities identified by the Decadal Survey and others.

Time Domain Astrophysics Program (Highest Priority Sustaining Activity for Space)

“Exploring the cosmos in the multi-messenger and time domains is a key scientific priority for the coming decade, with new capabilities for discovery on the horizon with the Rubin Observatory, Roman, LIGO/Virgo and the Kamioka Gravitational Wave Detector (KAGRA), and IceCube.”