

Investigating Core Induction Effect on Atmospheric Escape from Rocky Exoplanets and Implications for Habitability

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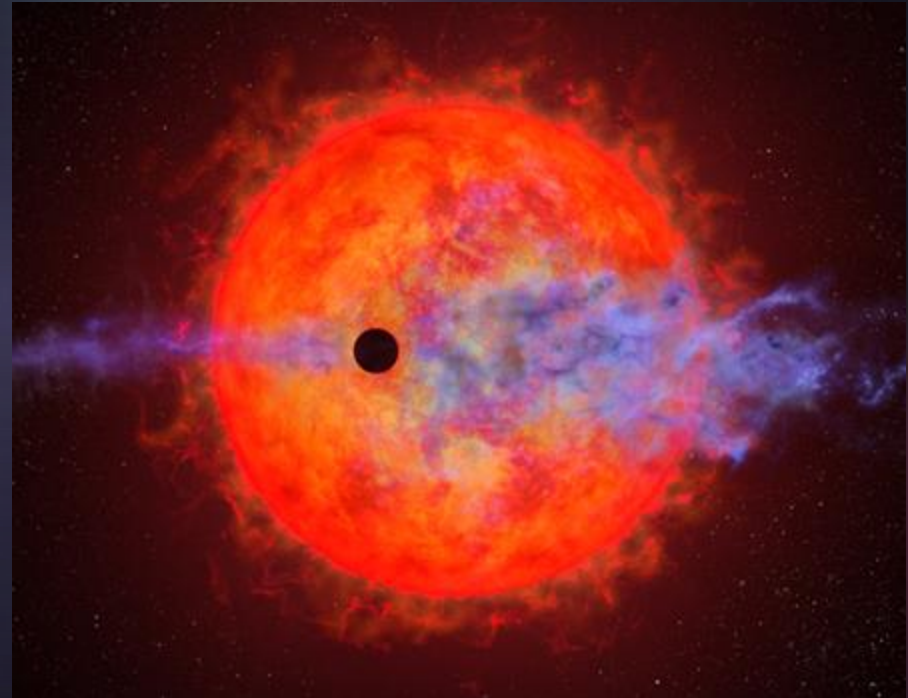
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ExoPAG 33 Early Career Talk

Jan. 3, 2026

Background & Motivation

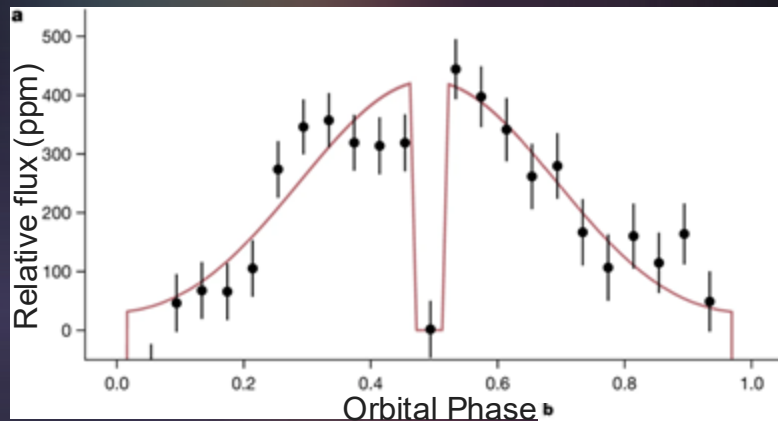
- Atmospheres are a key characteristic of habitable rocky worlds
 - Atmosphere must be sustained over long time periods
- **Nonthermal atmospheric ion escape**
 - Stellar wind erodes ions from upper atmosphere



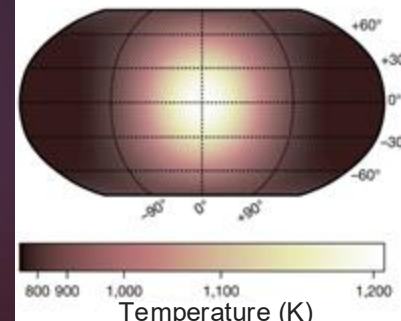
Artist's concept of atmospheric loss from an exoplanet orbiting an M-dwarf host star. NASA, ESA, Joseph Olmsted (STScI)

Background & Motivation

- Statistical trends emerge from 6,000+ confirmed exoplanets
 - Radius valley
 - Trends for terrestrial ($R < 2R_E$) planets?
- Campaigns to observe rocky exoplanet atmospheres (JWST)
- **What types of rocky exoplanets are most vulnerable to atmospheric ion loss?**

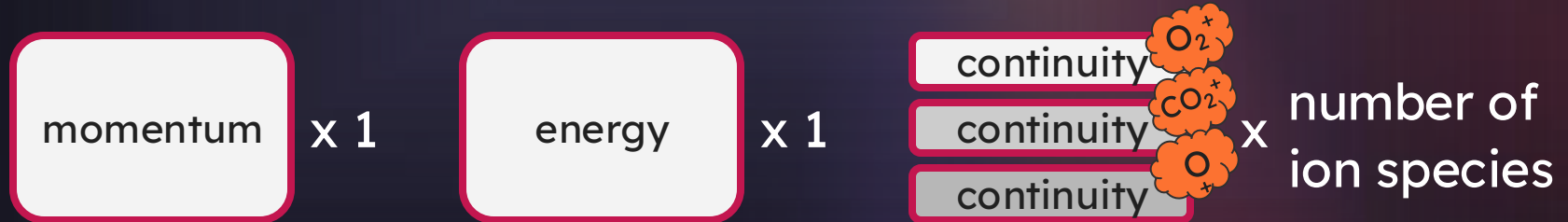


Thermal phase curve and best-fit temperature map of LHS 3844b, a close-in super-Earth orbiting an M-dwarf host. [Kreidberg+ 2019](#)



Methods

- MHD code: **BATS-R-US**
 - MHD component of Space Weather Modeling Framework (SWMF)
- Multi-species Magnetohydrodynamics (MS-MHD)
 - Chemistry: photoionization, charge exchange, electron recombination
 - Separate ion species
 - Assume same bulk velocity + temperature in all plasma



Planetary System Parameters

Planetary Radius

Conducting Core Size

THE ASTROPHYSICAL JOURNAL LETTERS

OPEN ACCESS

Role of Planetary Radius on Atmospheric Escape of Rocky Exoplanets

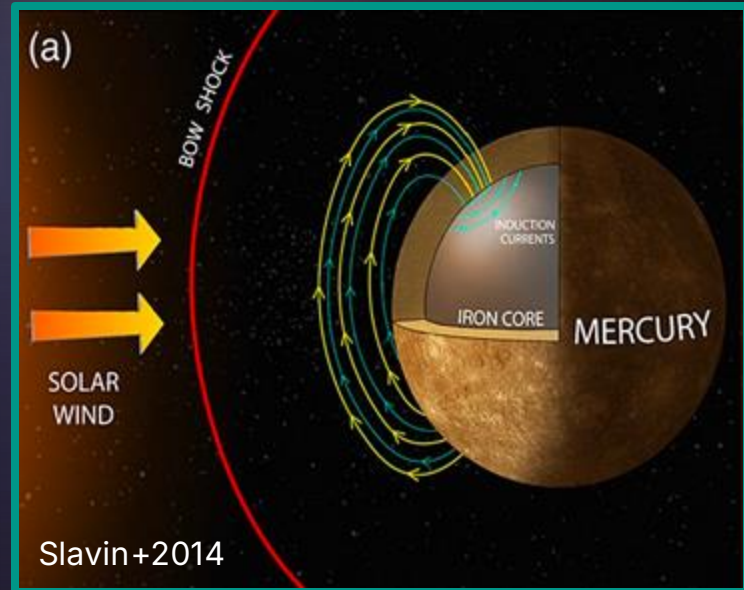
Laura Chin¹ , Chuanfei Dong¹ , and Manasvi Lingam^{2,3} 

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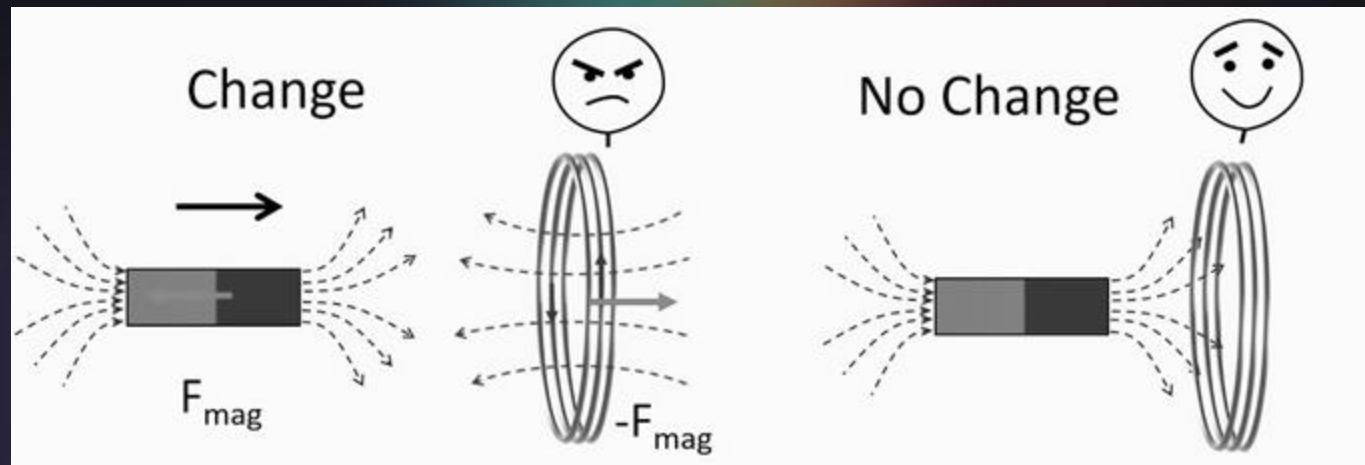
[The Astrophysical Journal Letters](#), Volume 963, Number 1

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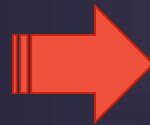
DOI 10.3847/2041-8213/ad27d8



Lenz's Law



loop of conducting material experiences change in magnetic flux



conservation of flux



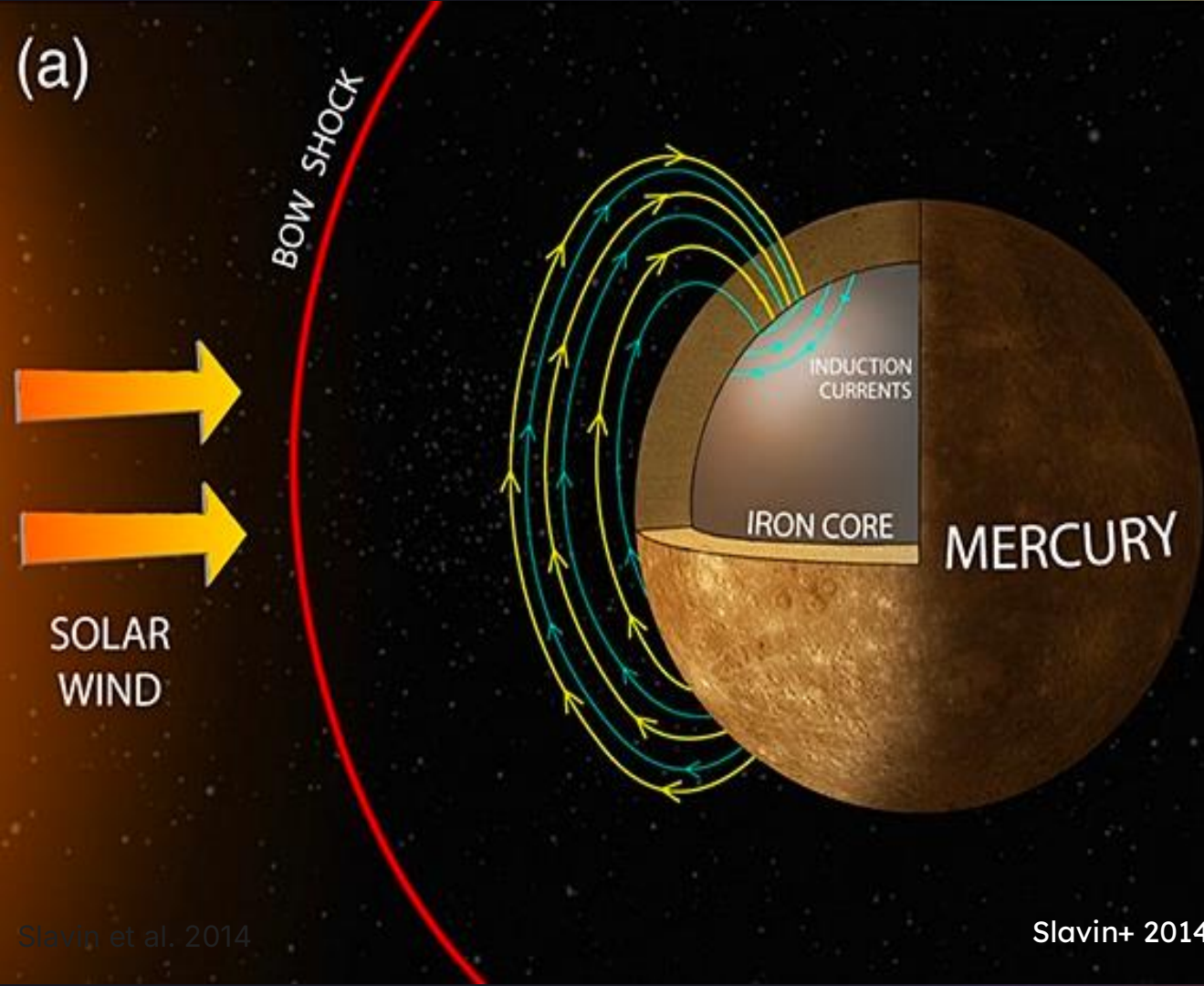
induced current flows through loop, generates opposing magnetic field



= conducting "loop"



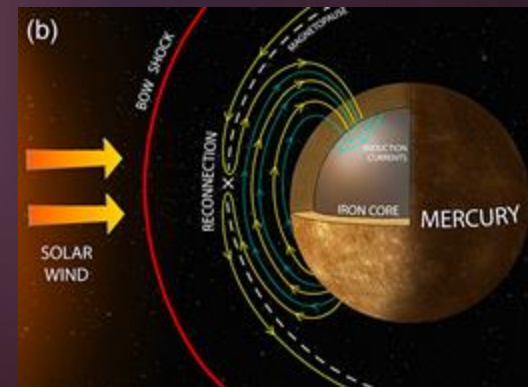
= external magnetic field



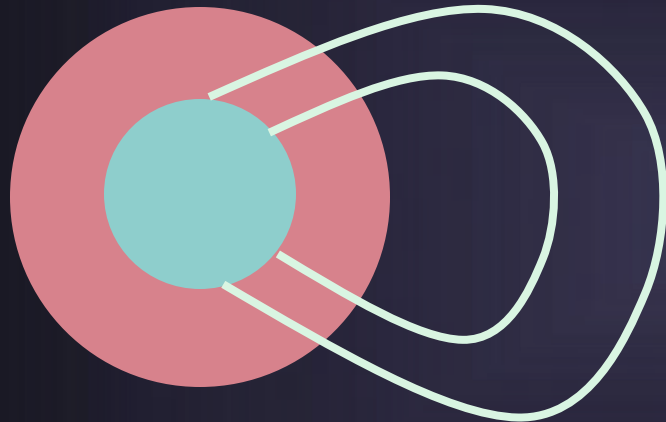
yellow lines =
intrinsic B field

green lines =
induced B field

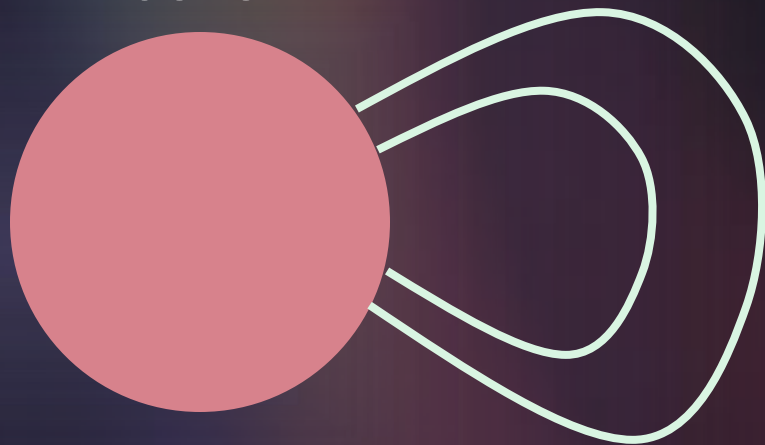
note: reconnection can
combat induction effect



Conducting
core



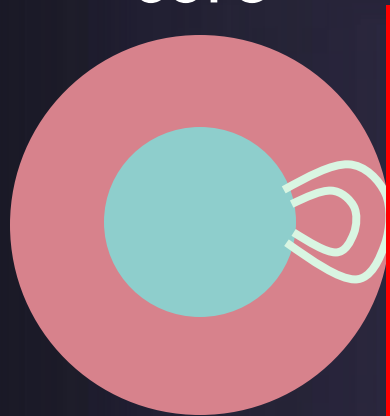
No conducting
core



IMF

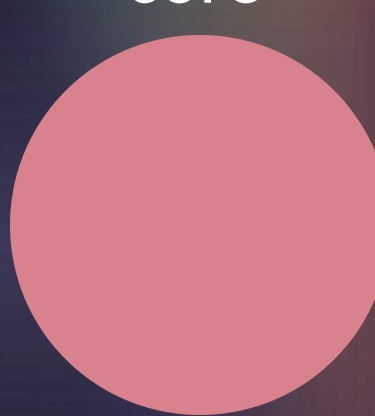
IMF

Conducting
core



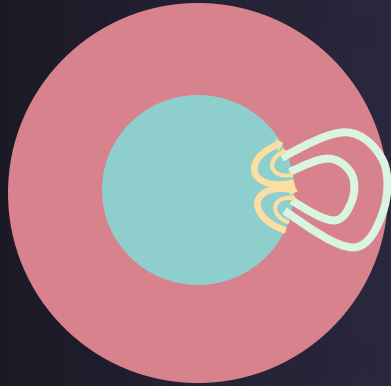
IMF

No conducting
core



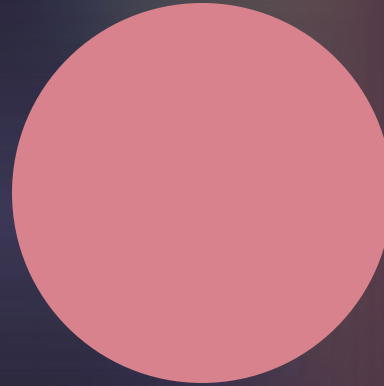
IMF

Conducting
core

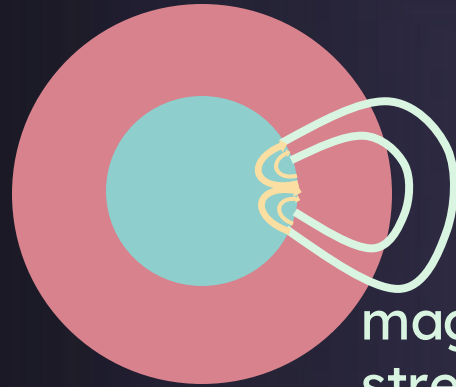


Induced
current

No conducting
core



Conducting
core

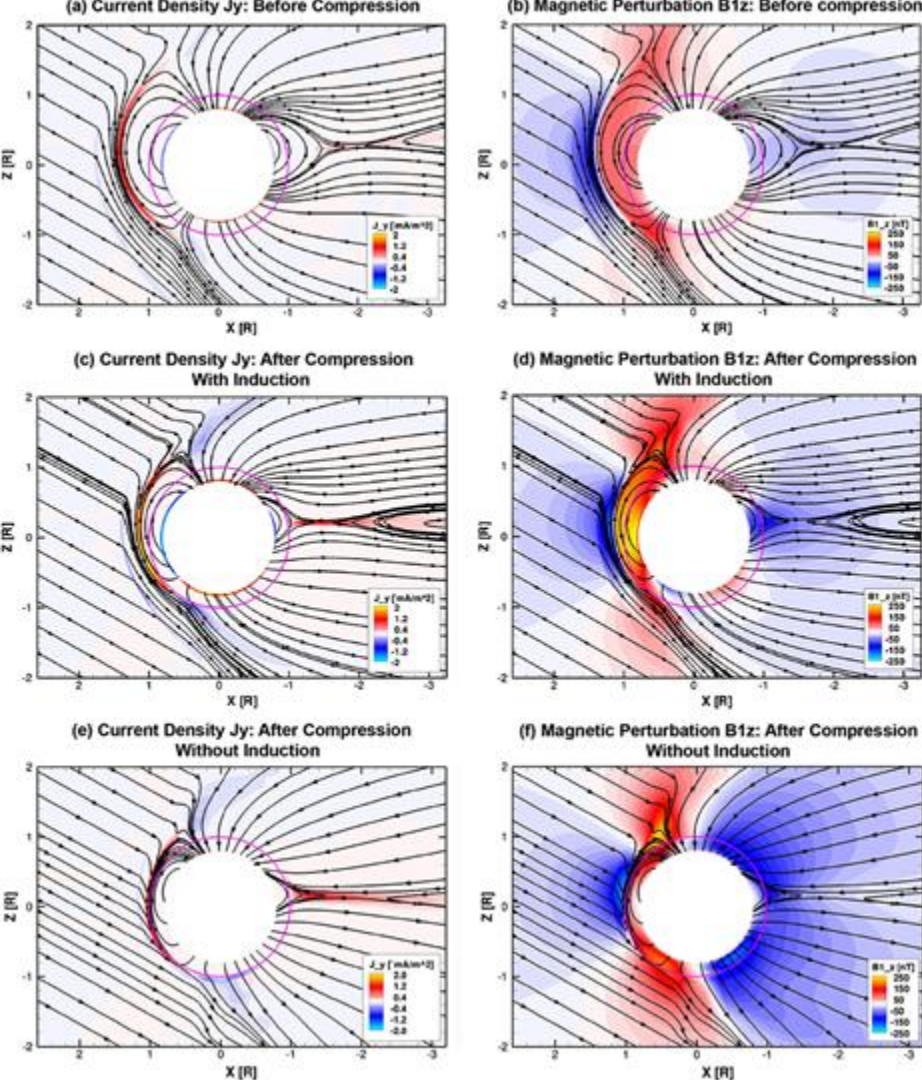


Dayside
magnetosphere
strengthened

Induced
current

No conducting
core





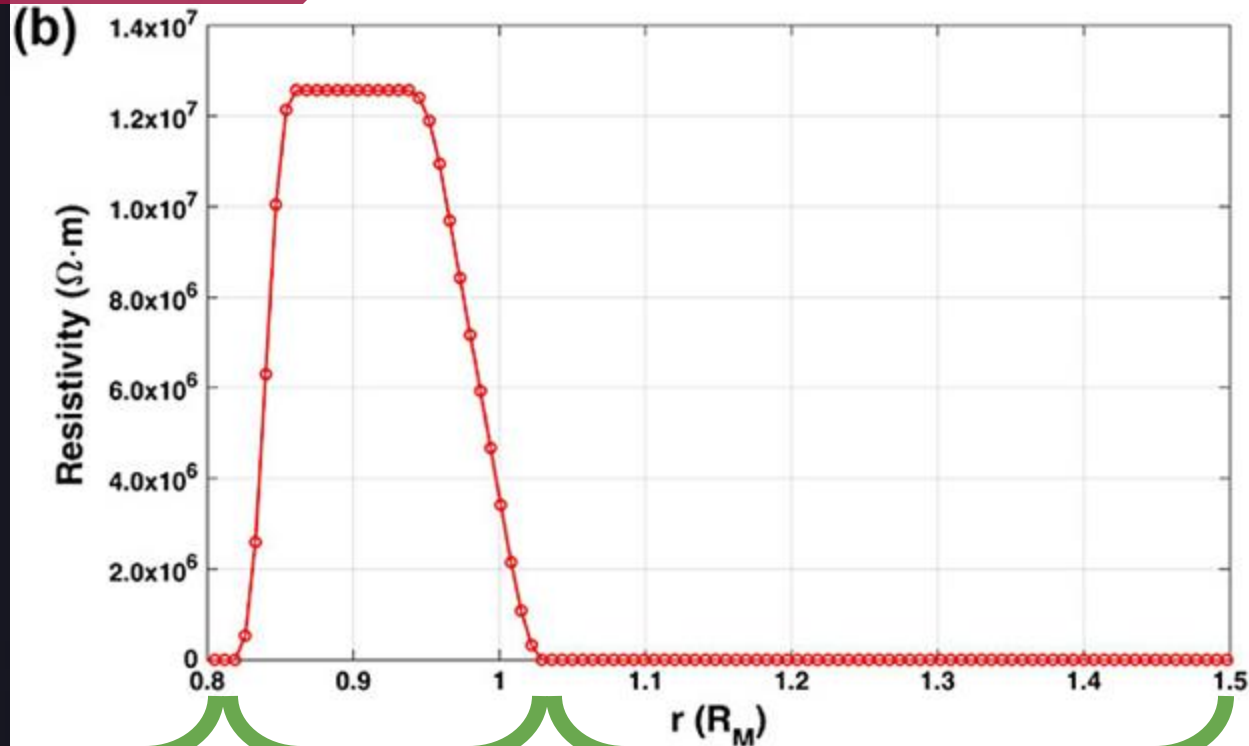
MHD study of induction effect
in context of Mercury
(no atmosphere)

Result: **induction effect enhances magnetosphere** in response to solar wind compression

Implications for atmospheric loss:
are planets with larger cores better
at protecting their atmospheres?

Jia+ 2015

Conducting Core



Resistivity profile

Inside planet
(core)

Mantle
high resistivity

Outside planet

Stellar Wind

$$n = 4245 \text{ cm}^{-3}$$

$$T = 10.0e5 \text{ K}$$

$$v = (-1937.5, 6.7, -13.0) \text{ km/s}$$

$$B = (0, 23.0, -194.3) \text{ nT}$$

(Exo)Planet Analog

Early (~4 Gya) Solar CME event with 10x higher density

Mass* Radius

$$M = 1.29 M_{\text{Earth}}$$

$$R = 1.1 R_{\text{Earth}}$$

Proxima Centauri b (PCb)
Earth, Venus, PCb

Core

$$R_C = 20\% \quad 50\% \quad 80\% R_p$$

$$\text{Dipole} = \frac{1}{2} \text{ Earth (15000 nT)}$$

Moon, Earth, Mercury
Early (~4 Gya) Earth

Atmosphere

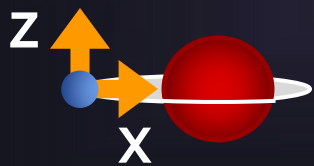
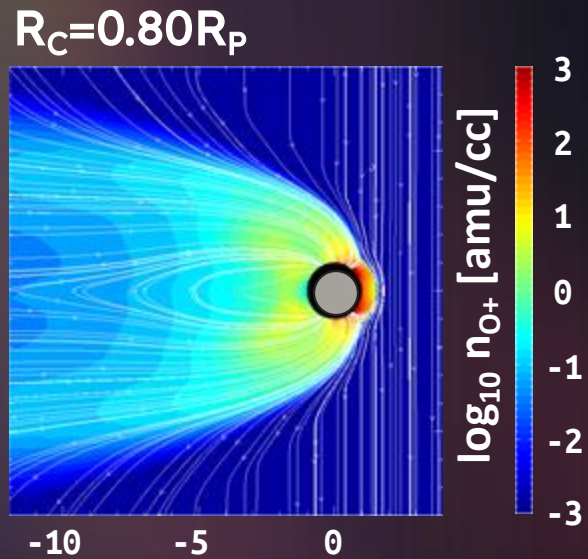
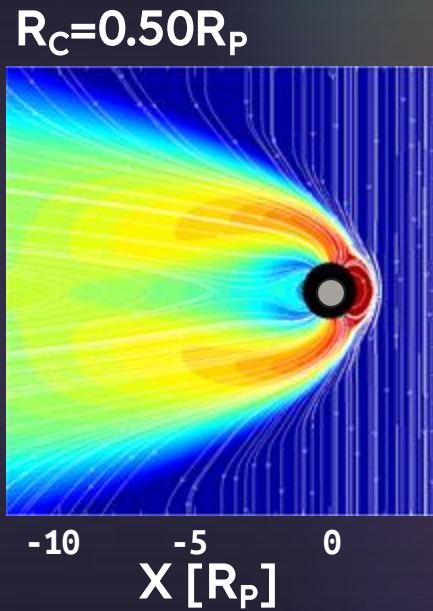
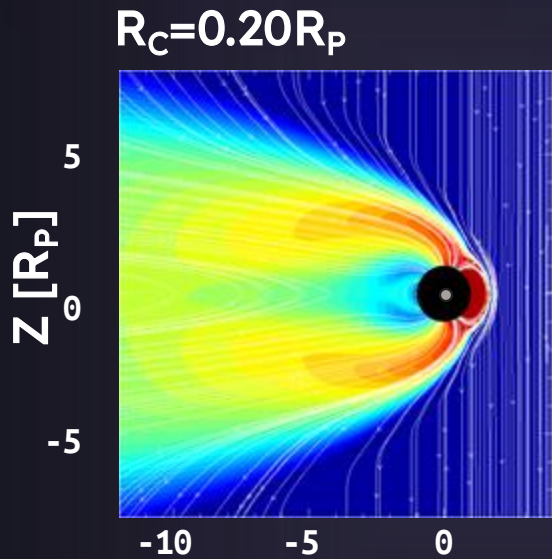
$$\text{Scale height} = 0.85 H_{\text{Venus}}$$

$$\text{Composition} = \text{CO}_2\text{-dominated}$$

PCb
Venus

Conducting Core

Core	$R_C = 0.20$	0.50	0.80
Induction Effect Only	1.29 M_{Earth}	1.29 M_{Earth}	1.29 M_{Earth}
Induction + Gravity Effect	1.11 M_{Earth} 0.99 H_{Venus}	1.29 M_{Earth} 0.85 H_{Venus}	1.91 M_{Earth} 0.58 H_{Venus}



Induction Effect
Only

Induction +
Gravity Effect

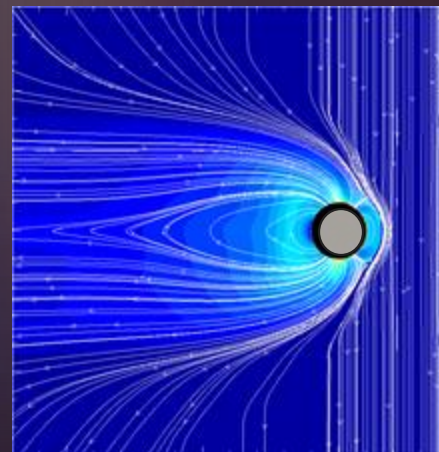
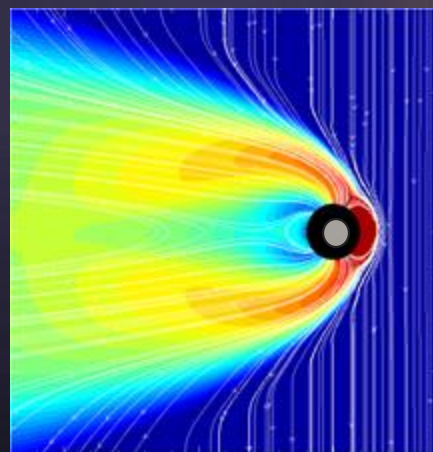
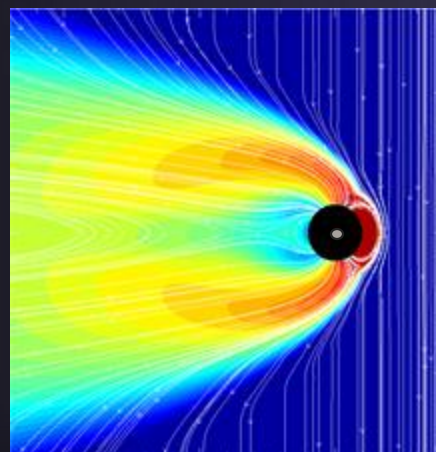
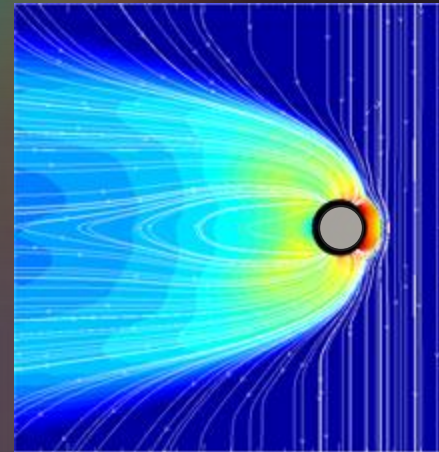
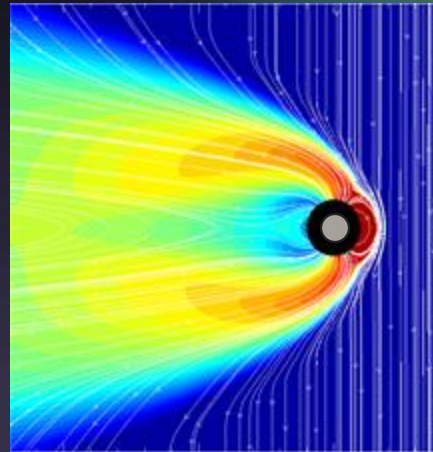
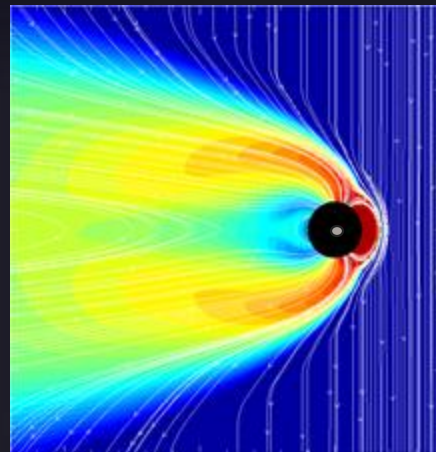
$R_C=0.20R_p$

$R_C=0.50R_p$

$R_C=0.80R_p$

$Z [R_p]$

$Z [R_p]$



-10

-5

0

-10

-5

0

-10

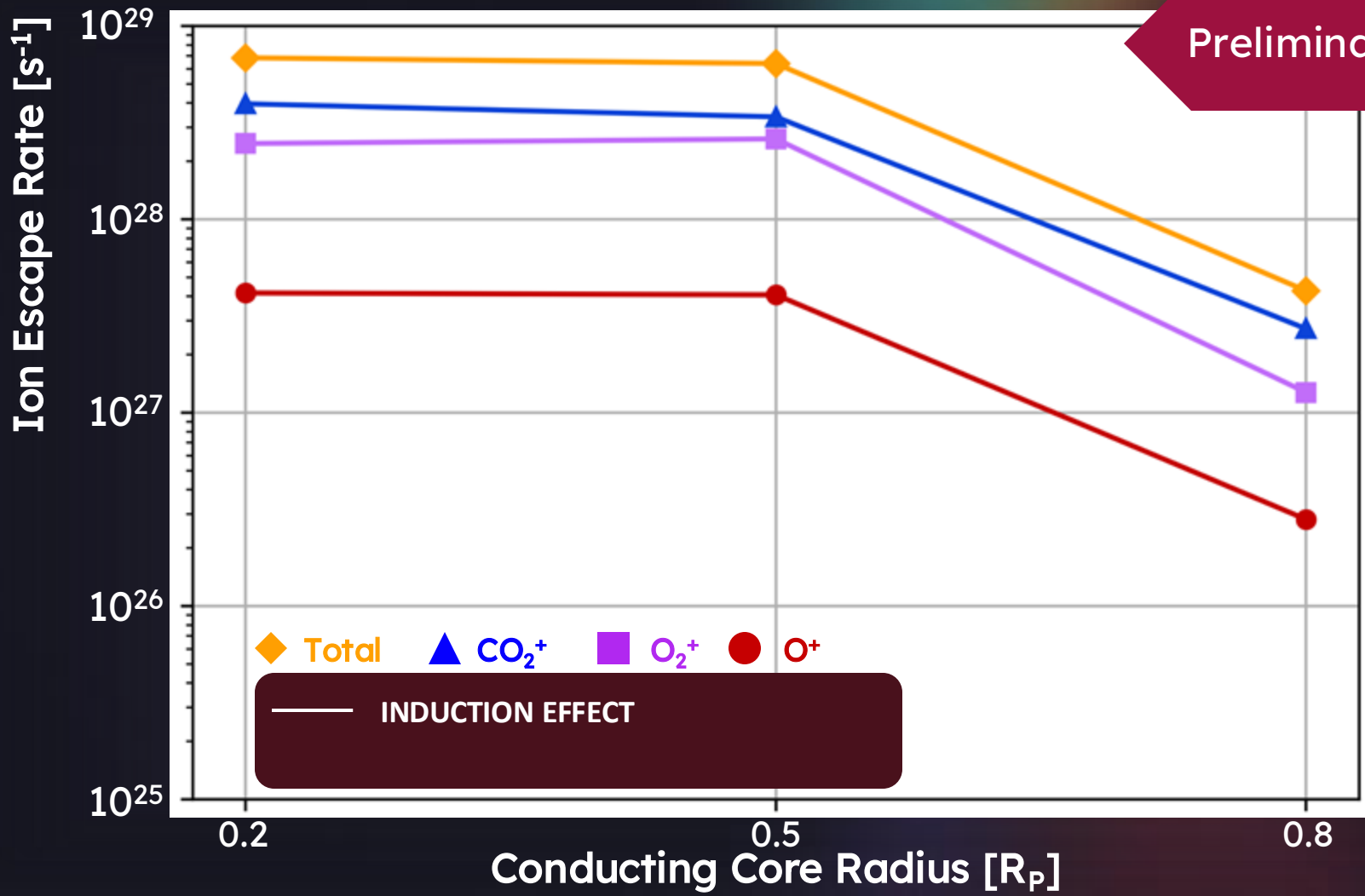
-5

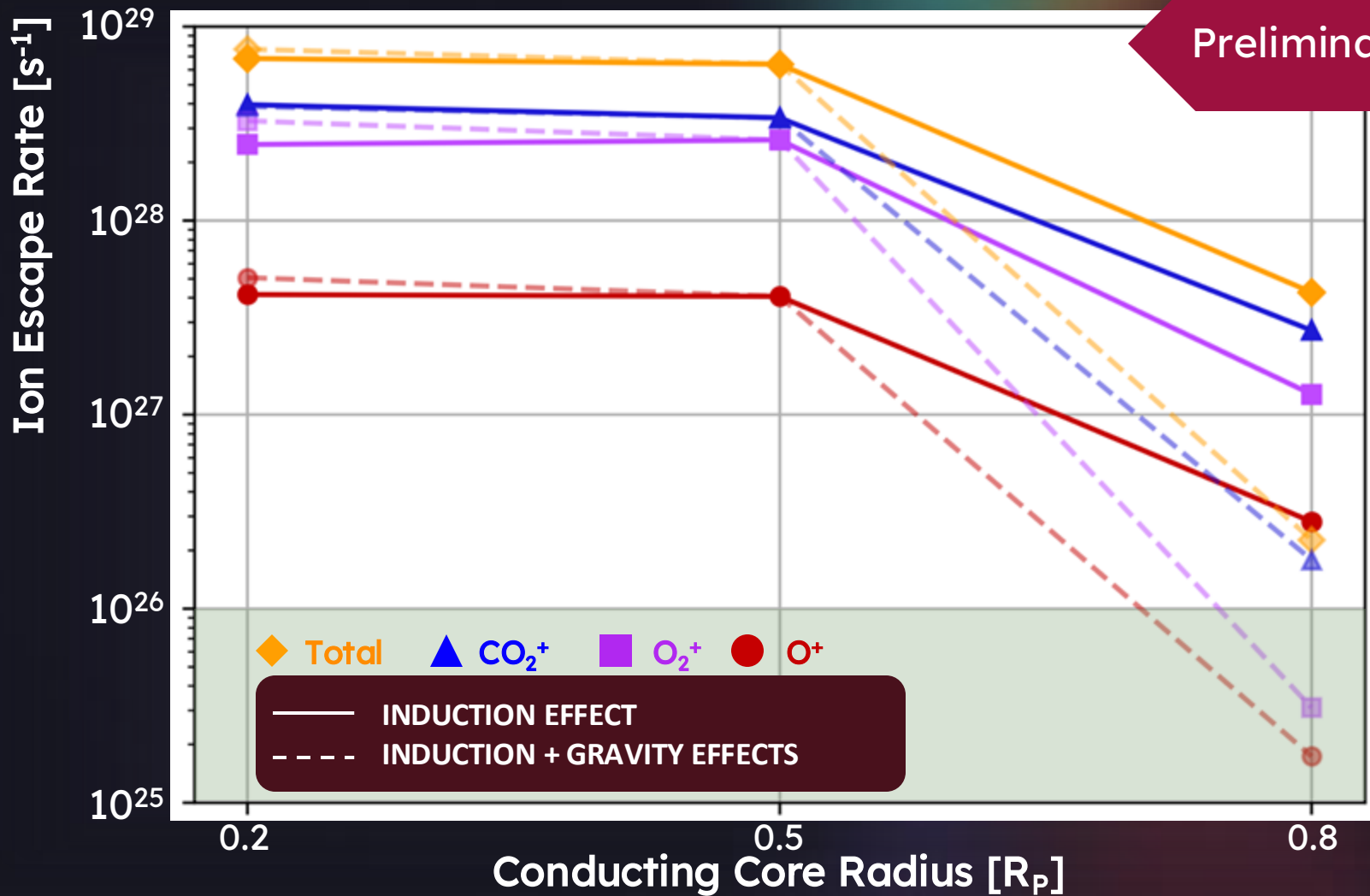
0

$X [R_p]$

$\log_{10} n_{O^+} [\text{amu/cc}]$

3
2
1
0
-1
-2
-3

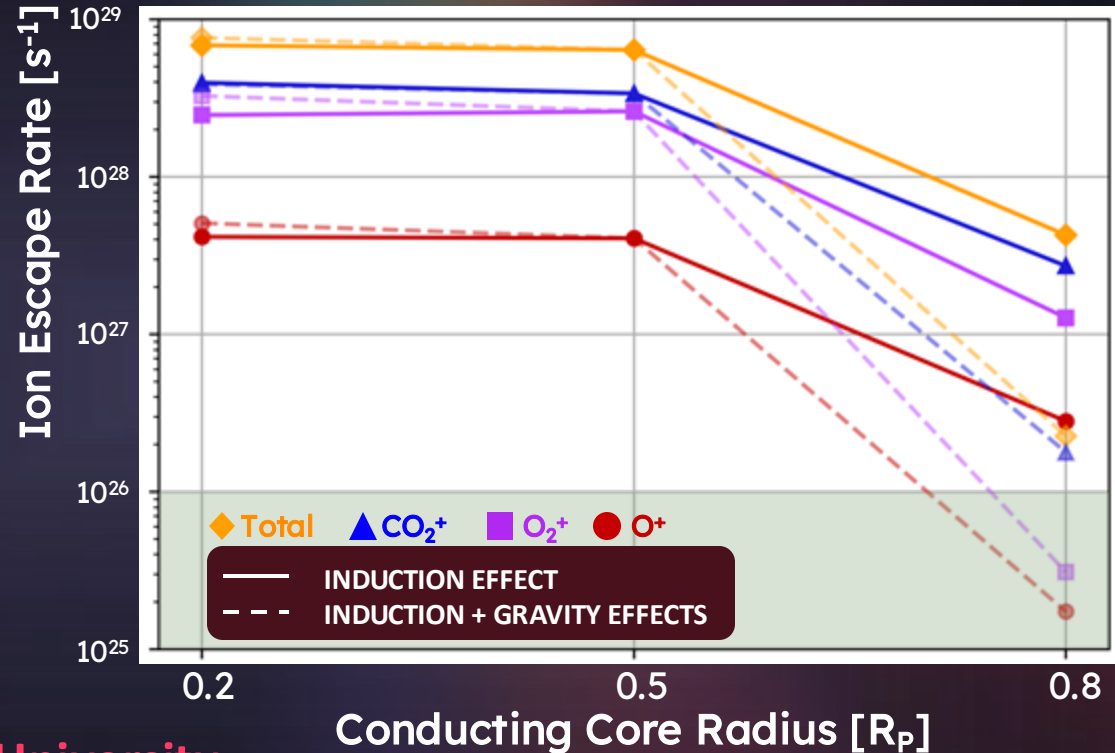




CURRENT
VENUS/EARTH/
MARS ESCAPE
RATES
Gronoff+ 2020

Summary

- The planet with the largest conducting core exhibits the lowest atmospheric ion escape rate across all species.
- The induction effect can help protect planets and reduce their atmospheric escape rate by generating additional core-induced magnetic fields.
- Planets with larger conducting cores may have better prospects for habitability.



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